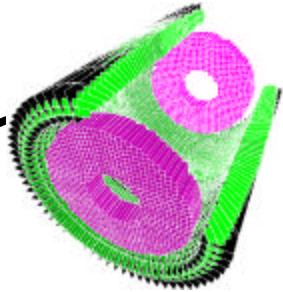


Bottomonium studies via $\Upsilon(3S)$ decays

First observation of $\Upsilon(1D)$

Tomasz Skwarnicki

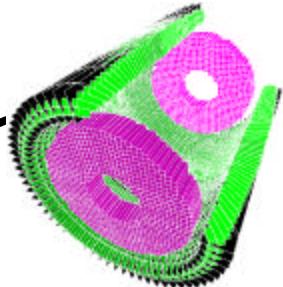
*Representing
the CLEO collaboration*



Onia

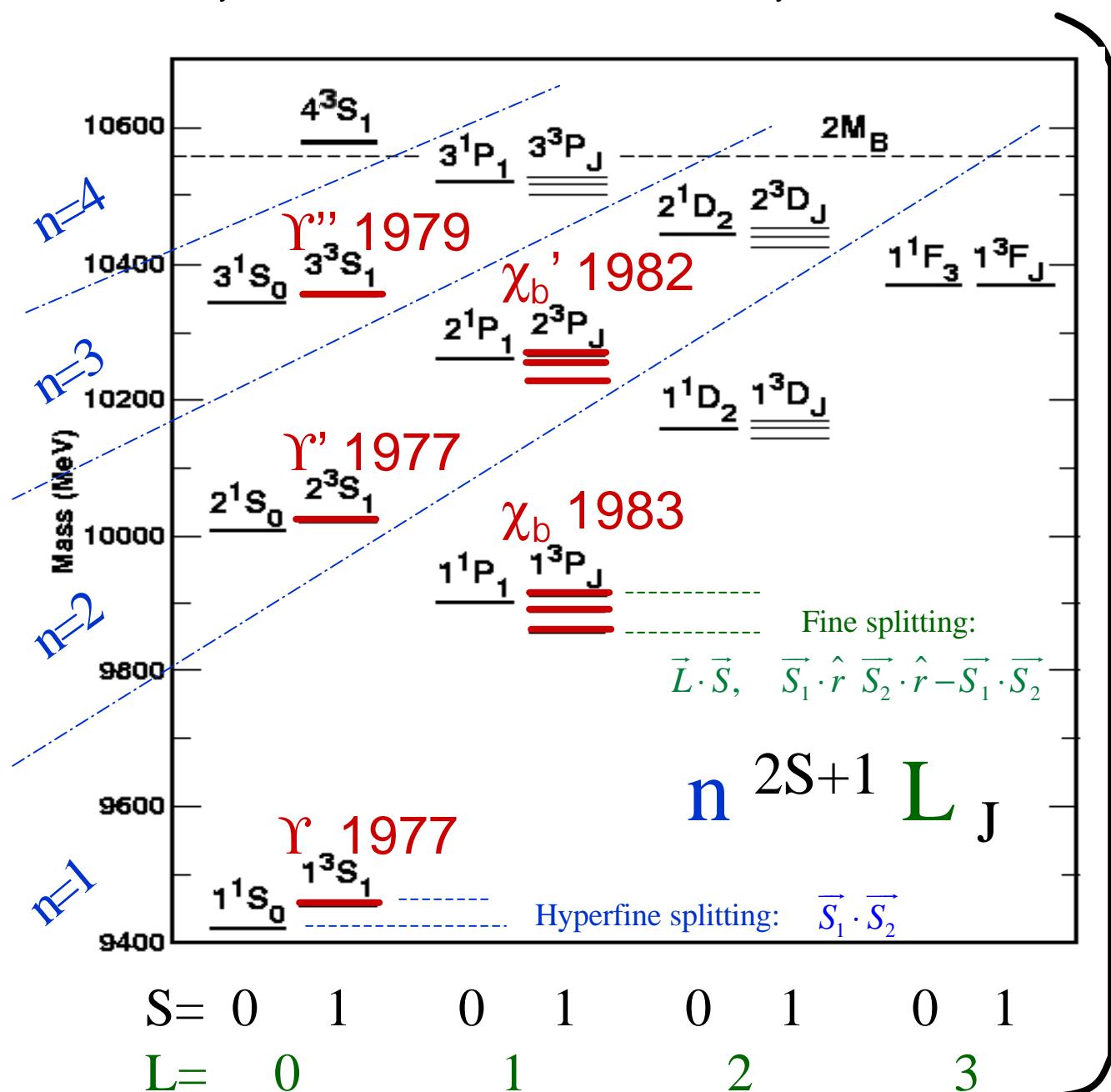
FORCES		Sys- tem	$(v/c)^2$	Ground triplet state 1^3S_1		Number of states below dissociation energy	
binding	decay			Name	Γ (MeV)	n^3S_1	all
POSITRONIUM							
EM	EM	e^+e^-	~ 0.0	Ortho-	$5 \cdot 10^{-15}$	2	8
QUARKONIUM							
STRONG	S	E	$u\bar{u}, d\bar{d}$	~ 1.0	ρ	150.00	0
	T		$s\bar{s}$	~ 0.8	ϕ	4.40	"1"
	R		$c\bar{c}$	~ 0.25	ψ	0.09	2
	O		$b\bar{b}$	~ 0.08	Υ	0.05	8
	N	weak	$t\bar{t}$	<0.01		3000.00	3
	G						30

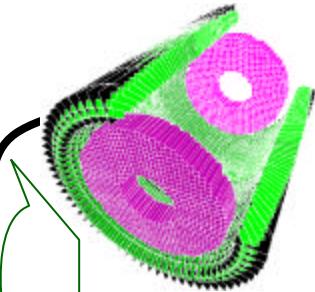
- Heavy quarkonia hold a promise of playing a similar role for QCD as positronium did for QED
 - Upsilonons are the most non-relativistic (i.e. simplest) states among all long-lived quarkonia states
 - The Upsilon system also has the largest number of stable states
- Upsilonons play a special role in probing the strong interactions (tests of lattice QCD, potential models)



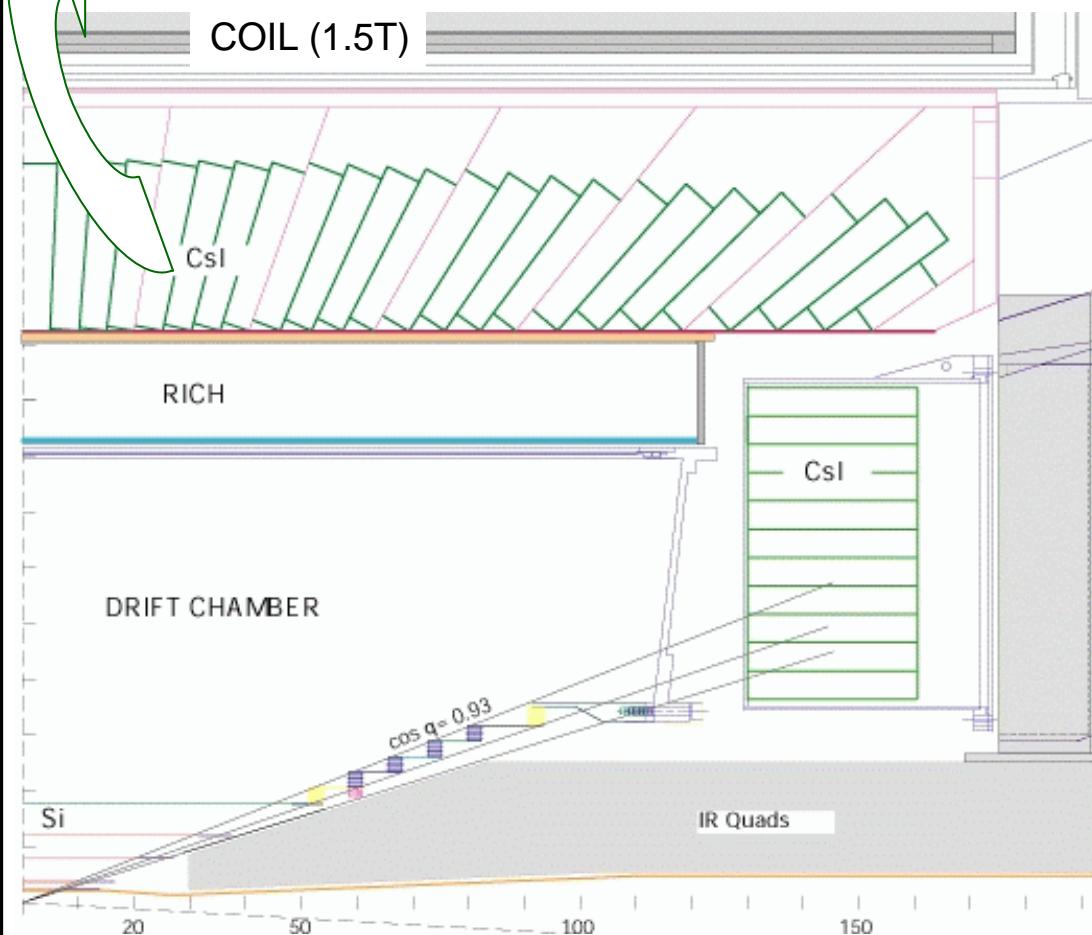
Upsilon States

- Only 9 out of 30 narrow states observed so far
- No spin-singlet states observed
- No new states observed in 19 years

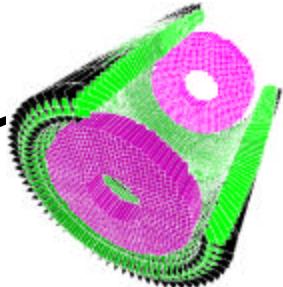




CLEO-III Detector

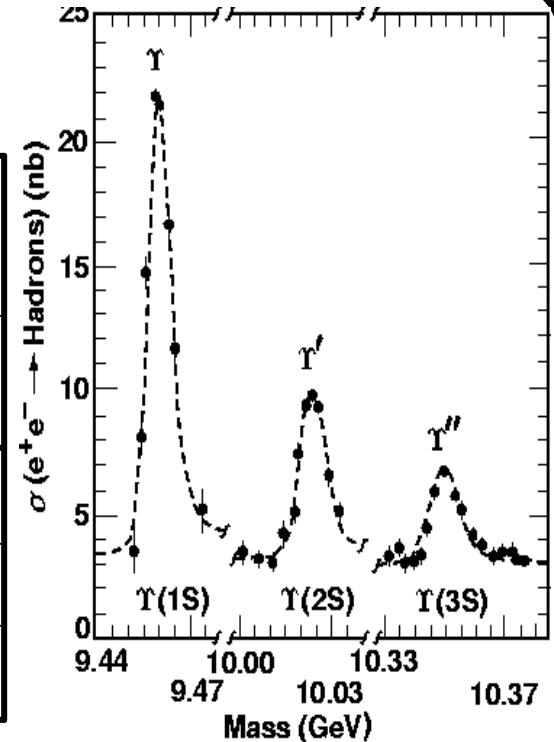


- **EM calorimeter - Essential for this work**
 - ~8000 CsI(Tl) crystals + photo-diodes
 - First crystal calorimeter in magnetic field
 - Changes since CLEO-II:
 - Low-mass DR endplate!
 - Re-stacked endcaps, moved away
 - New readout electronics
 - Some light loss due to the deteriorating glue used to attach the photo-diodes

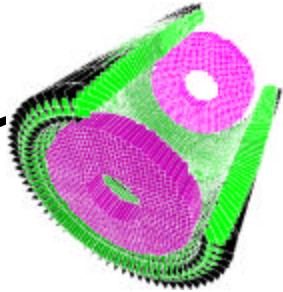


CLEO-III Υ Data

Reso-nance	CLEO-III Integrated Luminosity (pb^{-1})			Number of resonance decays (10^6)		
	ON	OFF	Scan	CLEOIII	CLEOII	CUSB (C.Ball)
$\Upsilon(3S)$	1150	128	100	4.7	0.46	1.3
$\Upsilon(2S)$	550	150	50	~3.6	0.49	(0.19)
$\Upsilon(1S)$	1230	200	100	~29.0	1.9	(0.48)

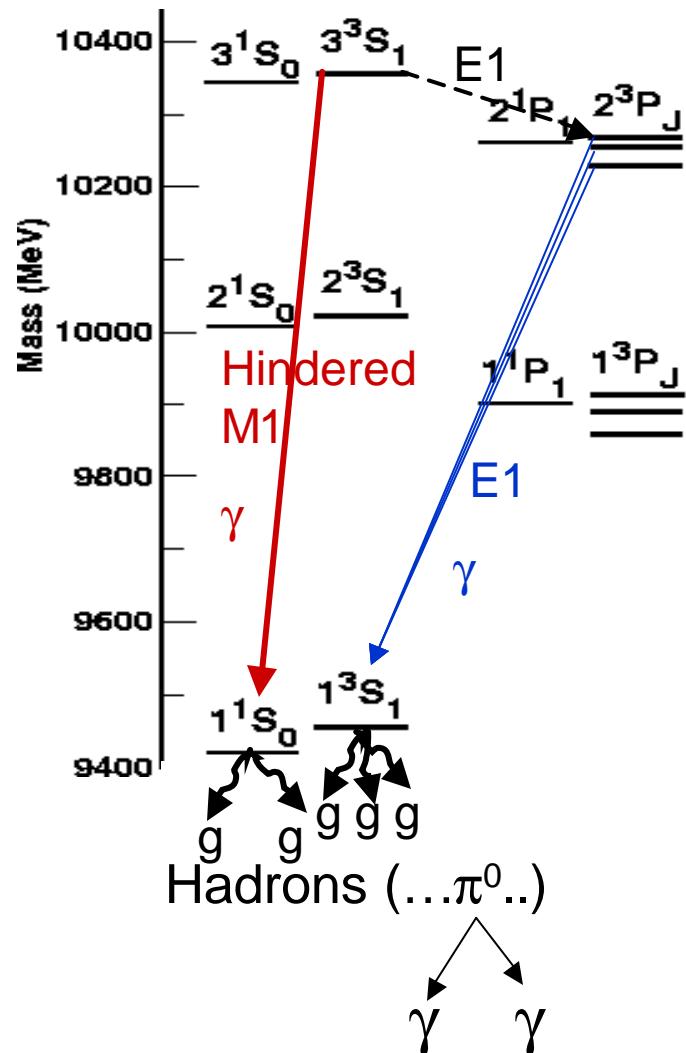


- About 10-fold increase over the CLEO-II statistics
- About 4-times more $\Upsilon(3S)$ data than analyzed by CUSB + 2.5-4.5 times higher efficiency for the final states analyzed here
- The $\Upsilon(3S)$ data already processed off-line
- We will take more Υ data before lowering the beam energy to the charm threshold region next year (\rightarrow CLEO-c)

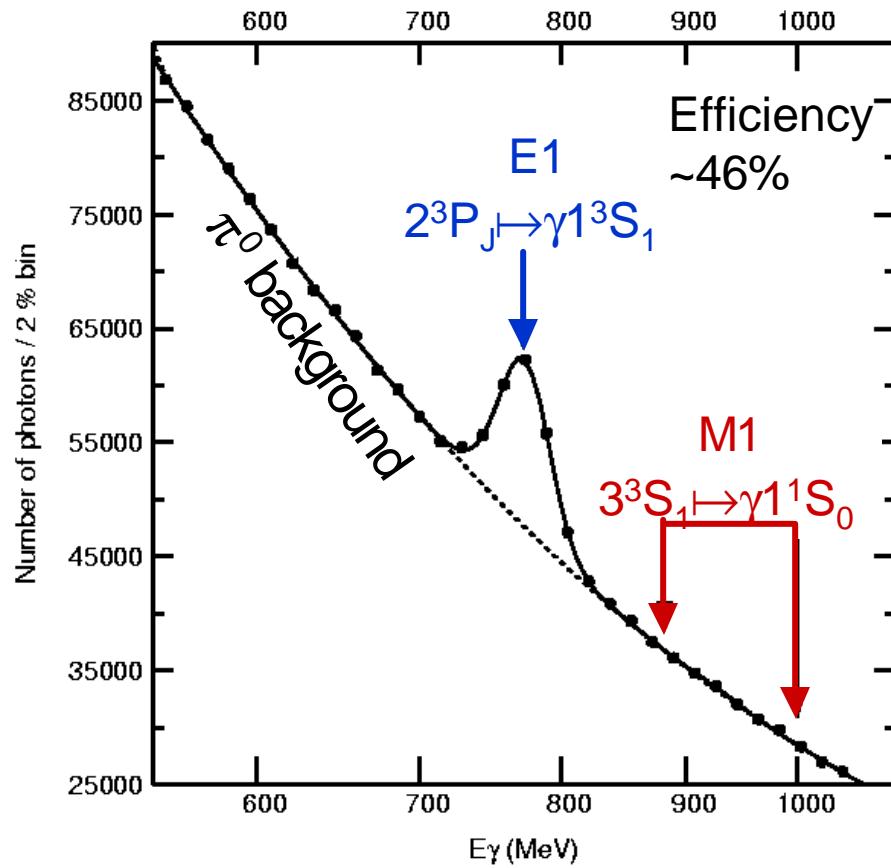


ICHEP ABS947, CLEO CONF 02-05

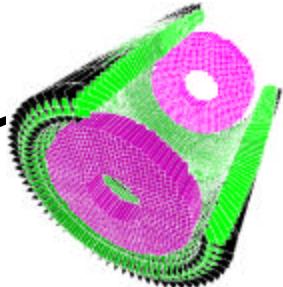
Search for $\eta_b(1^1S_0)$



Inclusive g spectrum
in multi-hadronic events

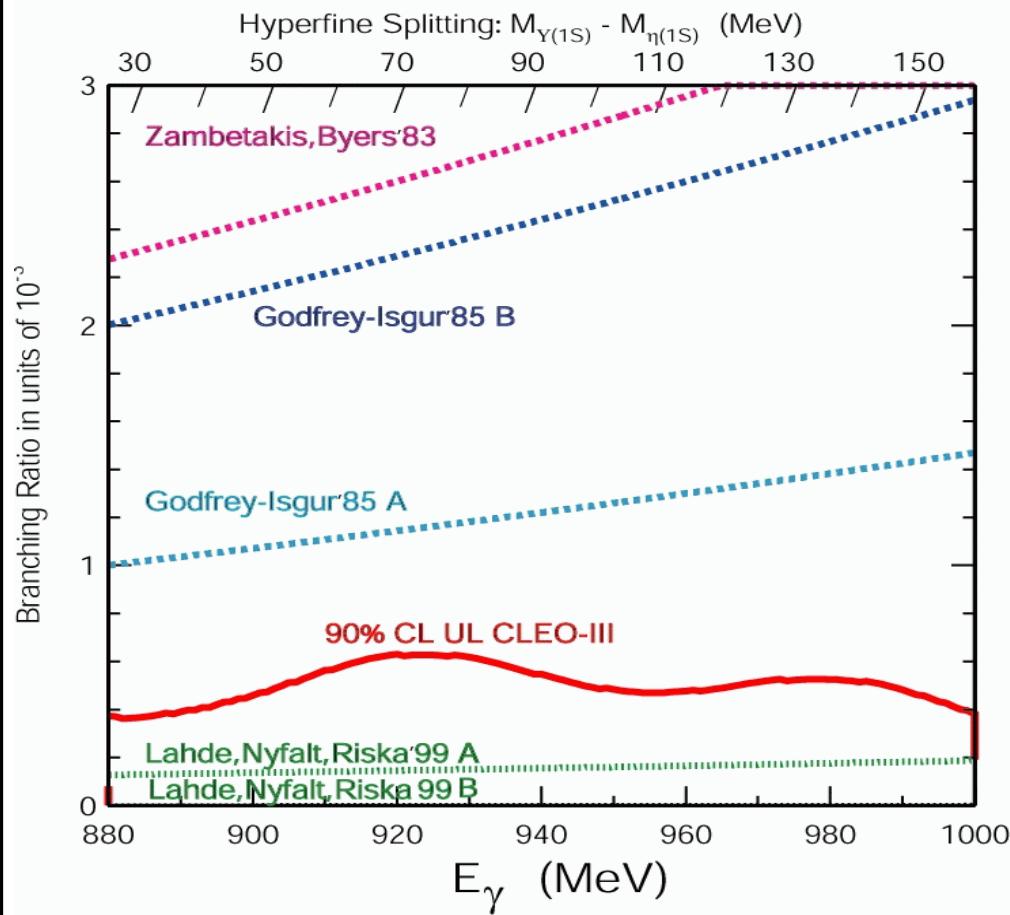


- No signal found



Search for $\eta_b(1^1S_0)$

- Test potential model predictions for Γ_{M1}



Models from the compilation by
Godfrey&Rosner PR D64, 074011
(2001) [scaled here by phase-space]

$$\Gamma_{M1} \propto \frac{e_b^2}{m_b^2} \left| \langle n_f L | n_i L \rangle \right|^2 E_g^3$$

DIRECT $n_i = n_f$

$\langle n_f L | n_i L \rangle = 1$ E_g^3 -tiny

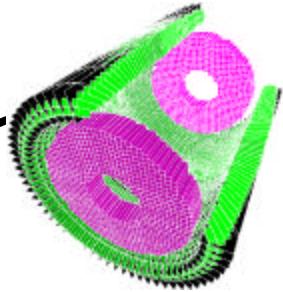
hopeless for $b\bar{b}$

HINDERED $n_i \neq n_f$

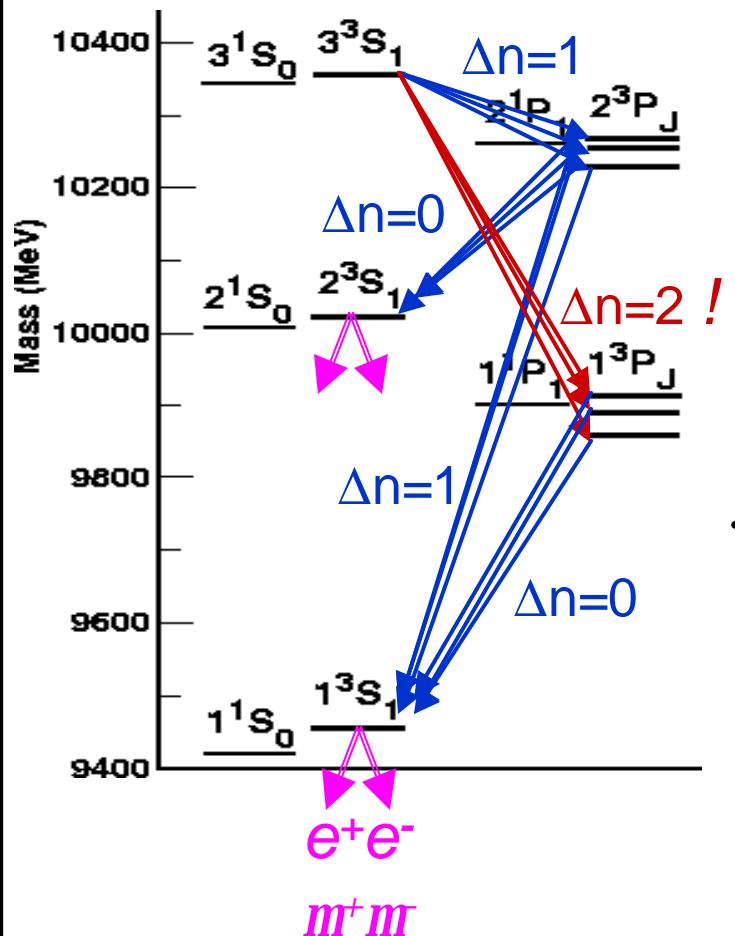
$\langle n_f L | n_i L \rangle \approx 0$ E_g^3 -large

difficult to predict

- Most of the calculations are ruled out!



Exclusive 2γ -cascades



- $\gamma\gamma l^+l^-$ final states
- No π^0 backgrounds from gluonic $b\bar{b}$ annihilation
- Low product branching ratio (a few 10^{-4})
- Sensitivity to hadronic widths of triplet P-states:

$$B(P \mapsto \gamma S) = \Gamma_{E1}/(\Gamma_{E1} + \Gamma_{had})$$
- $3S \mapsto \gamma 1P$ is a $\Delta n=2$ transition (rare)

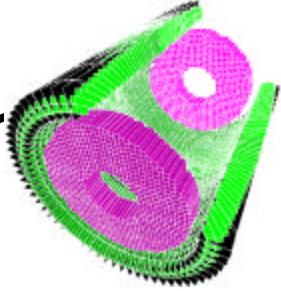
$$\Gamma_{E1} \propto e_b^2 \left| \langle n_f L_f | r | n_i L_i \rangle \right|^2 E_g^3$$

much larger than Γ_{M1} , since no suppression by $1/m_b^2$
 $L_f = L_i \pm 1$

as $\Delta n = n_i - n_f$ increases

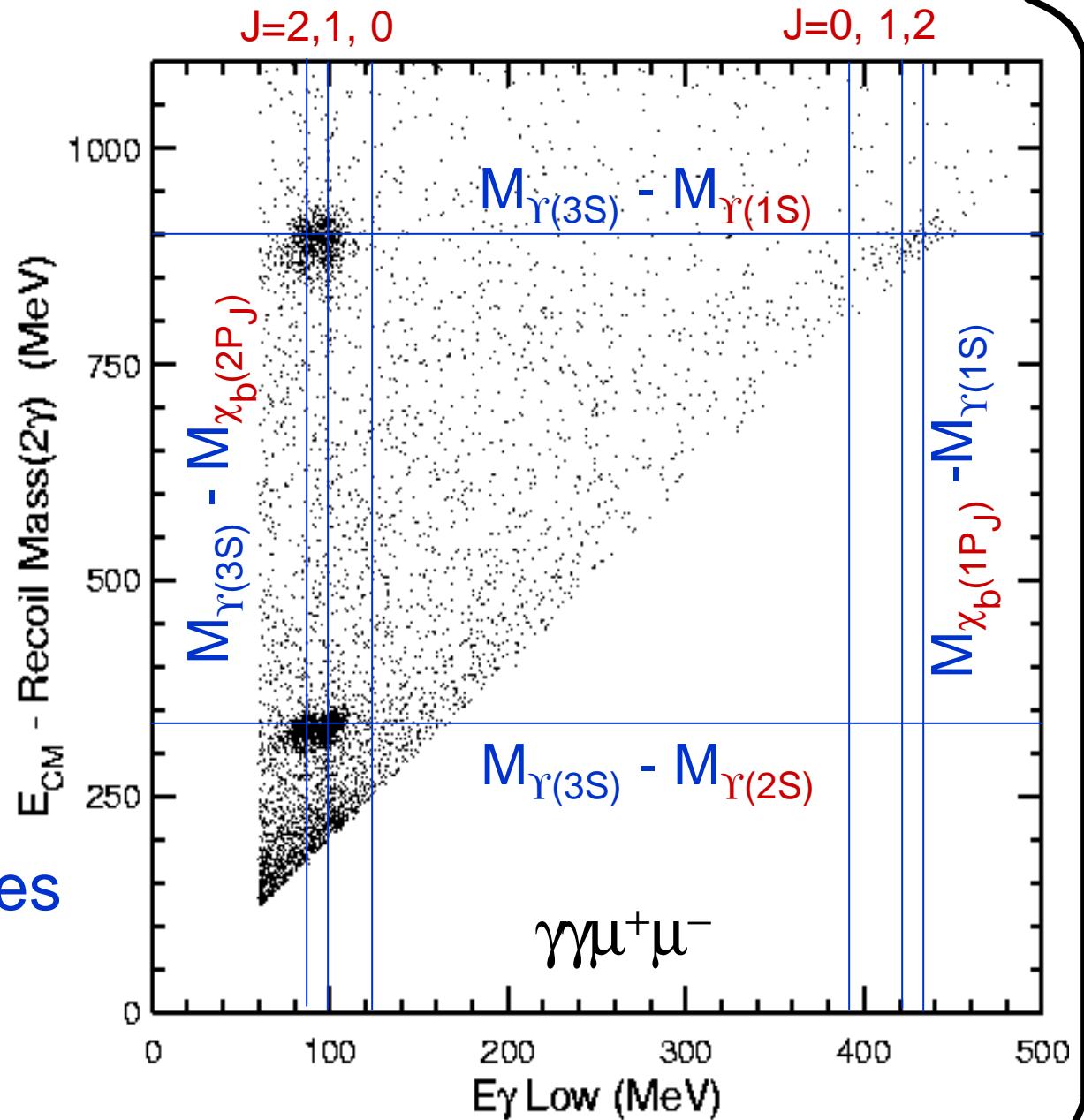
$$\langle n_f L_f | r | n_i L_i \rangle \text{ decreases}$$

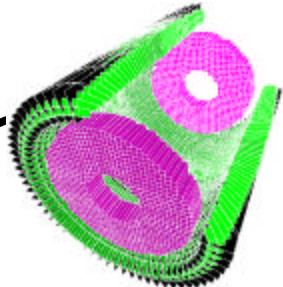
and Γ_{E1} becomes more difficult to predict



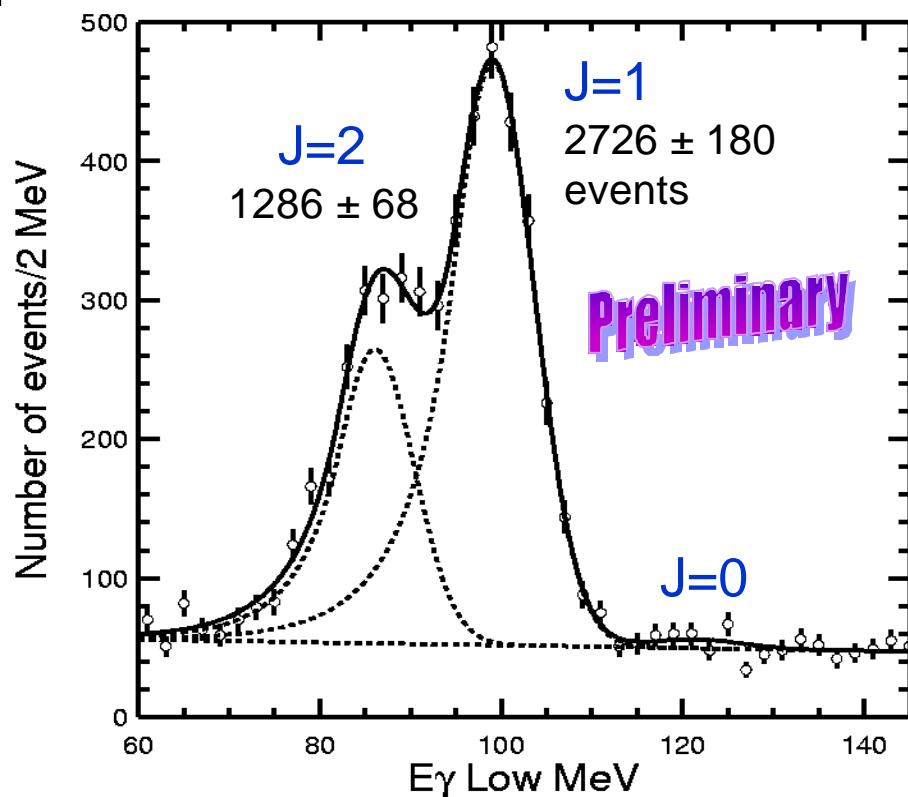
Exclusive 2γ -cascades

- Signal variables





$$\Upsilon(3S) \mapsto \gamma \chi_b(2P_J) \mapsto \gamma\gamma \ell^+ \ell^-$$

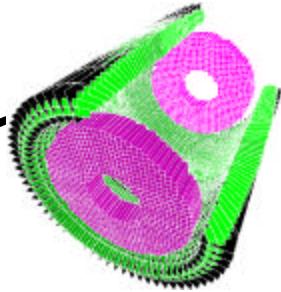


- Photon-line energies $\mapsto M(\chi_b(2P_J))$
 - $E_{J=2} = 86.09 \pm 0.30 \pm 0.29$ MeV
 - $E_{J=1} = 99.08 \pm 0.17 \pm 0.34$ MeV

Efficiency
 $(\mu\mu + ee)/2$
~32%

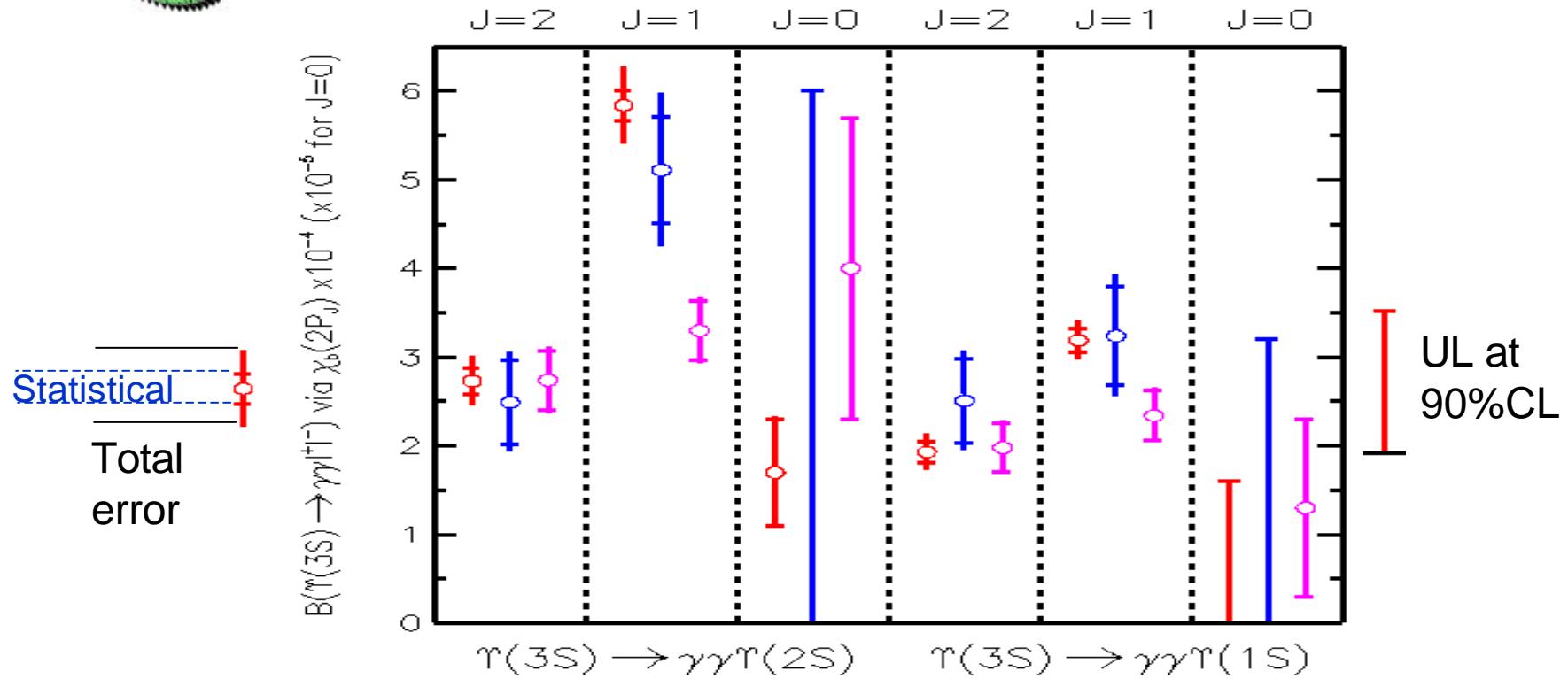
Energy resolution from the fit:
 4.6 ± 0.2 MeV @ 100 MeV

- Energy calibrated to $\pm 0.34\%$ with the photon-recoil mass and known $\Upsilon(nS)$ masses
 - More precise than previous measurements
 - Consistent with CUSB and CLEO-II results

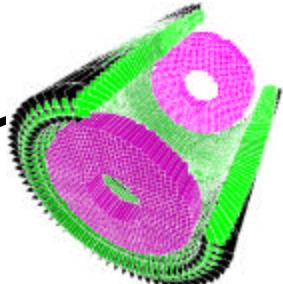


$$\Upsilon(3S) \mapsto \gamma \chi_b(2P_J) \mapsto \gamma \ell^+ \ell^-$$

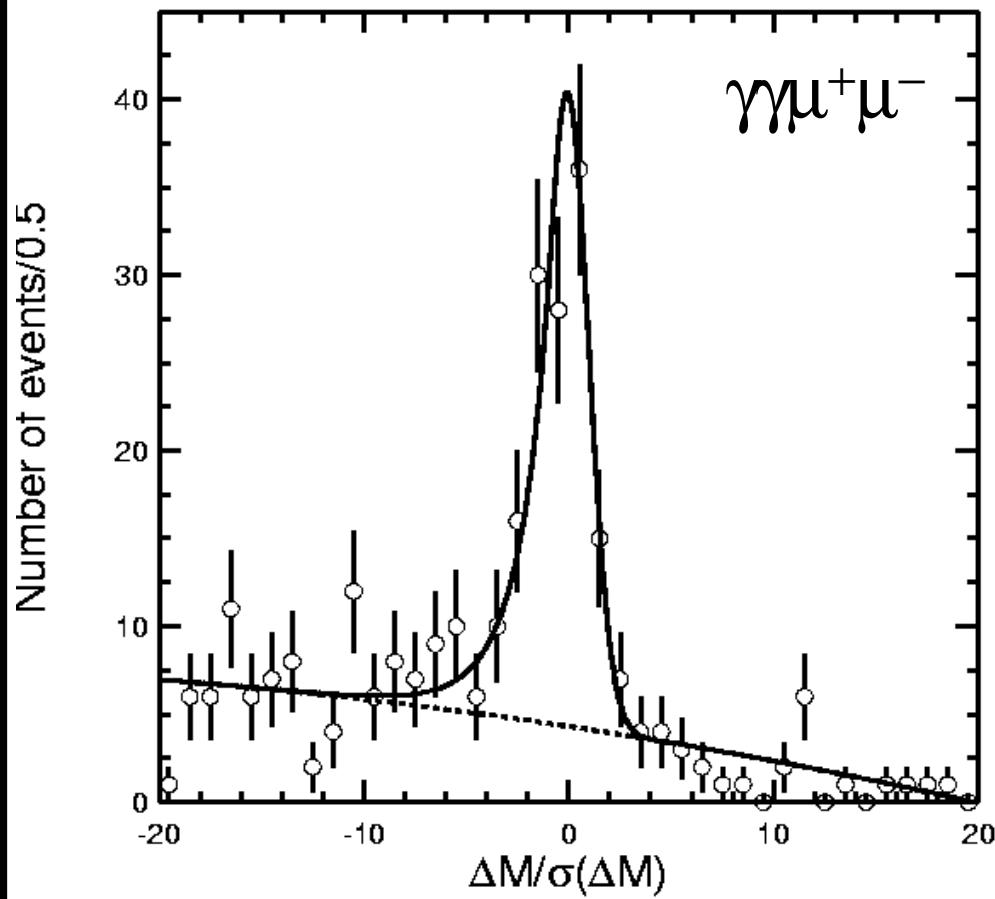
CLEO-III, CLEO-II, CUSB



- Throughout this talk: $B(\gamma \ell^+ \ell^-) = (B(\gamma \mu^+ \mu^-) + B(\gamma e^+ e^-))/2$
- Good agreement with CLEO-II
- CUSB measurements low for $J=1$
- Much improved errors in CLEO-III



$$\Upsilon(3S) \mapsto \gamma \chi_b(1P_J) \mapsto \gamma \ell^+ \ell^-$$



- Since cannot resolve $J=2,1,0$ states, fit the recoil mass distribution instead

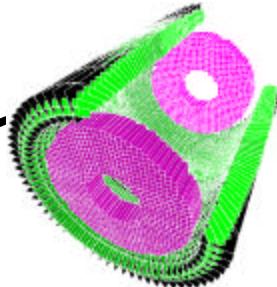
Events ($\mu\mu + ee$): 167 ± 19 events
Efficiency: $\sim 37\%$

$$B(\Upsilon(3S) \mapsto \gamma \chi_b(1P_{2,1}) \mapsto \gamma \gamma \Upsilon(1S) \mapsto \gamma \ell^+ \ell^-) = (5.2 \pm 0.5 \pm 0.5) 10^{-5}$$

$$B(\Upsilon(3S) \mapsto \gamma \chi_b(1P_{2,1}) \mapsto \gamma \gamma \Upsilon(1S)) = (2.1 \pm 0.2 \pm 0.2) 10^{-3}$$

CUSB

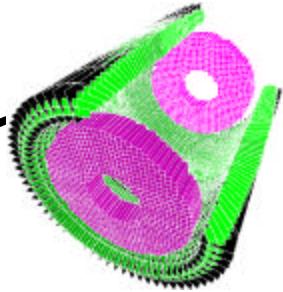
$(1.2 \pm 0.4 \pm 0.1) 10^{-3}$



Comparison of the measured E1 transition rates with the potential models

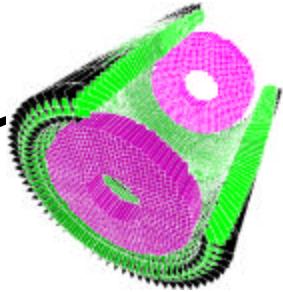
	$ < 2P r 3S > $	$ < 1P r 2S > $	$ < 1P r 3S > $	$ < 1S r 2P > $	$ < 2S r 2P > $	
	GeV^{-1}	GeV^{-1}	GeV^{-1}			
DATA	2.7 ± 0.2	1.9 ± 0.2	0.050 ± 0.006	0.096 ± 0.005		
	World Average		This measurement			
Model	NR	rel	NR	rel	NR	rel
Kwong,Rosner [13]	2.7		1.6		0.023	
Fulcher [14]	2.6		1.6		0.023	
Büchmuller et al.[15]	2.7		1.6		0.010	
Moxhay,Rosner [16]	2.7	2.7	1.6	1.6	0.024	0.044
Gupta et al.[17]	2.6		1.6		0.040	
Gupta et al.[18]	2.6		1.6		0.010	
Fulcher [19]	2.6		1.6		0.018	
Danghighian et al.[20]	2.8	2.5	1.7	1.3	0.024	0.037
McClary,Byers [21]	2.6	2.5	1.7	1.6		
Eichten et al.[22]	2.6		1.7		0.110	
Grotch et al.[23]	2.7	2.5	1.7	1.5	0.011	0.061

- Potential models:
 - easily reproduce the large E1 matrix elements
 - have trouble predicting small elements
(see $\gamma(3S) \mapsto \gamma \chi_b(1P_J)$ $\Delta n=2$)



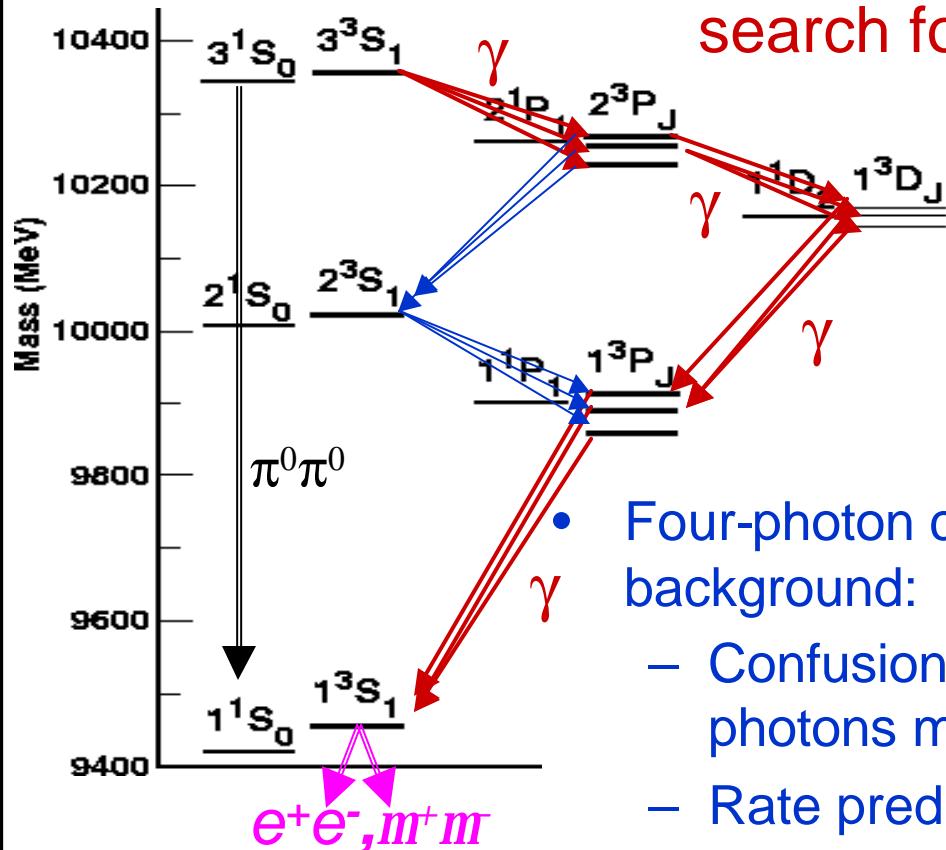
Searches for π^0, η transitions

- Also could contribute to $\gamma\gamma\ell^+\ell^-$ events
- Suppress photon transitions
- Look at $M\gamma\gamma$
- No signal found
- At 90% C.L.
 - $B(\Upsilon(3S) \rightarrow \pi^0 \Upsilon(1S)) < 0.17 \cdot 10^{-3}$
 - $B(\Upsilon(3S) \rightarrow \eta \Upsilon(1S)) < 0.9 \cdot 10^{-3}$
 - $B(\Upsilon(3S) \rightarrow \pi^0 \Upsilon(2S)) < 1.2 \cdot 10^{-3}$



Exclusive 4 γ -cascades

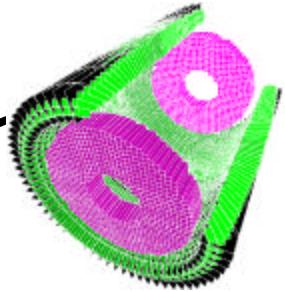
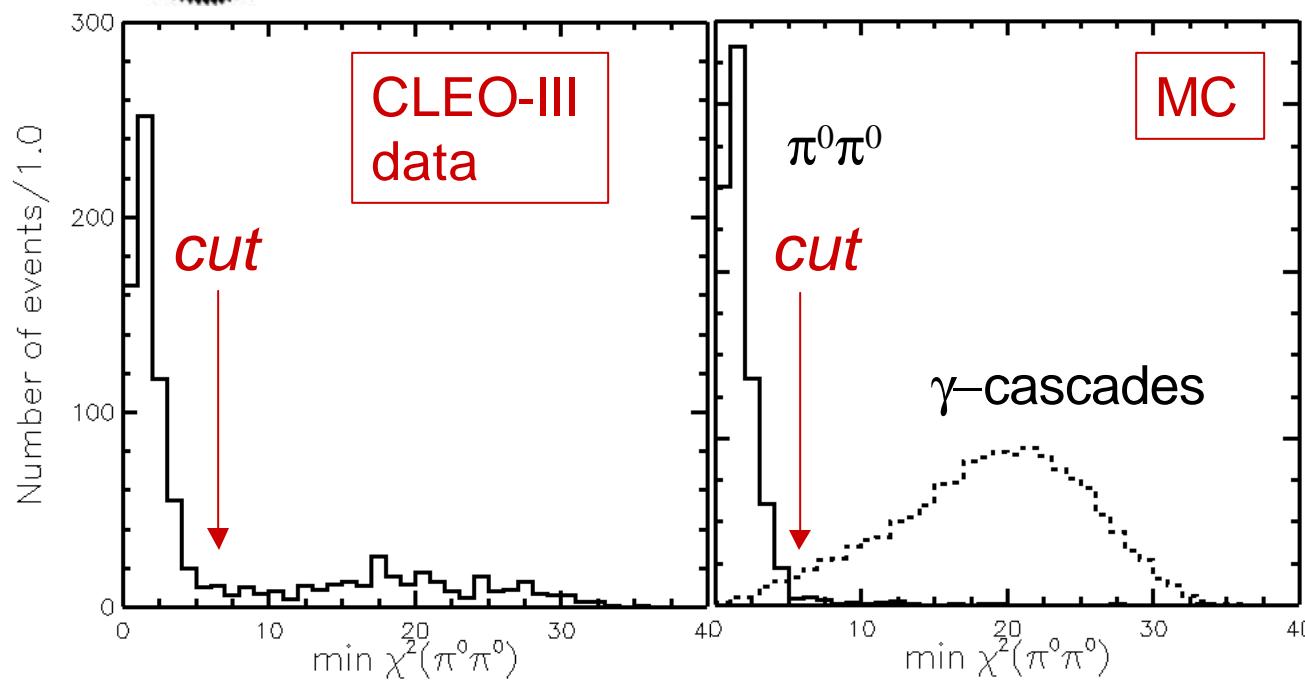
- Can use four-photon E1 cascade to search for $\Upsilon(1^3D_J)$!



Signal rate predicted by
Godfrey&Rosner: $3.76 \cdot 10^{-5}$
PR D64, 097501 (2001)

- Four-photon cascade via the $\Upsilon(2S)$ – the main background:
 - Confusion in ordering of the observed photons makes these two cascades similar
 - Rate predicted by Godfrey&Rosner: $3.84 \cdot 10^{-5}$

- Also $\Upsilon(3S) \rightarrow \pi^0\pi^0 \Upsilon(1S)$ is a potential background


 $\Upsilon(3S) \rightarrow \pi^0\pi^0 \Upsilon(1S)$


- Selection/rejection of $\Upsilon(3S) \rightarrow \pi^0\pi^0 \Upsilon(1S)$

Events ($\mu\mu + ee$):
 737 ± 28 events

Efficiency:
 $\sim 14\%$

$B(\Upsilon(3S) \rightarrow \pi^0\pi^0 \Upsilon(1S) \rightarrow \gamma\gamma \ell^+\ell^-) = (5.7 \pm 0.2 \pm 0.4) \cdot 10^{-4}$

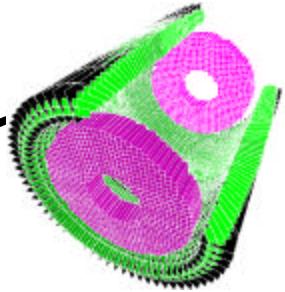
$B(\Upsilon(3S) \rightarrow \pi^0\pi^0 \Upsilon(1S)) = (2.33 \pm 0.09 \pm 0.16) \cdot 10^{-2}$

CLEO-II

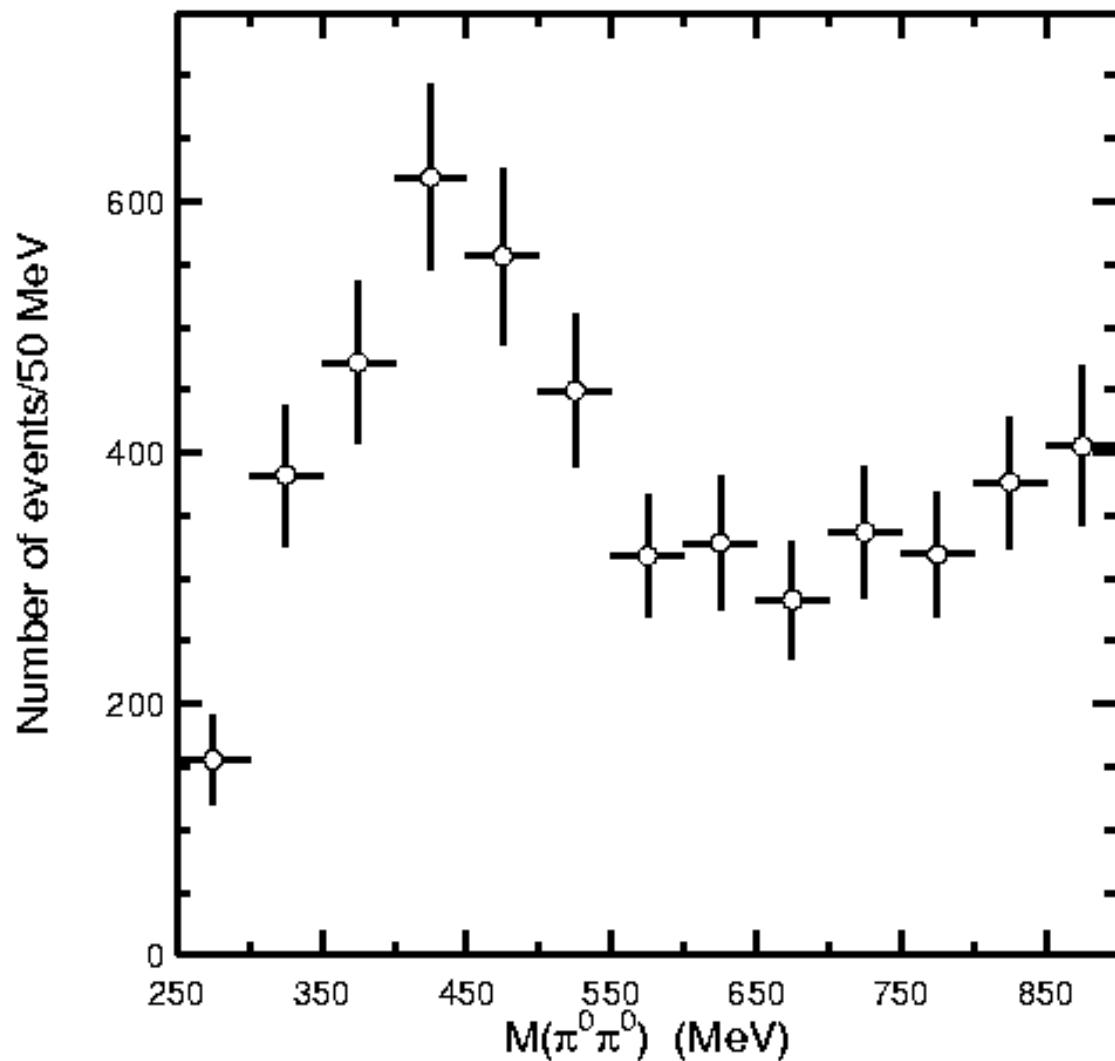
$(2.03 \pm 0.28 \pm 0.19) \cdot 10^{-2}$

CUSB

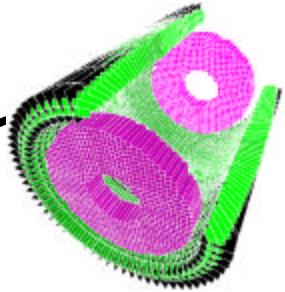
$(2.3 \pm 0.4 \pm 0.3) \cdot 10^{-2}$



$$\Upsilon(3S) \rightarrow \pi^0\pi^0 \Upsilon(1S)$$



- Double-peak structure in $M(\pi^0\pi^0)$ confirmed here with more data



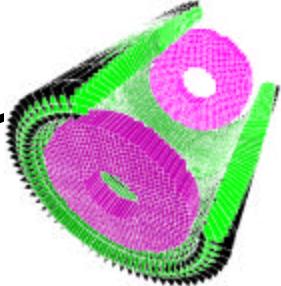
Selection of $\Upsilon(1D)$ events

$$\mathbf{c}_{1D}^2 = \min_{M_{1D}, J_{2P}, J_{1P}} \sum_{i=1}^4 \left(\frac{E_{g_i} - E_{g_i}^{\text{expected}}(M_{1D}, J_{2P}, J_{1P})}{s(E_{g_i})} \right)^2$$

- Implements constraints to the well known masses: M_{3S} , M_{2P_J} , M_{1P_J} , M_{1S}
- In addition to χ^2_{1D} value also obtain “most likely” mass of $\Upsilon(1D)$ for each event
- To suppress cascades through the $\Upsilon(2S)$ calculate:

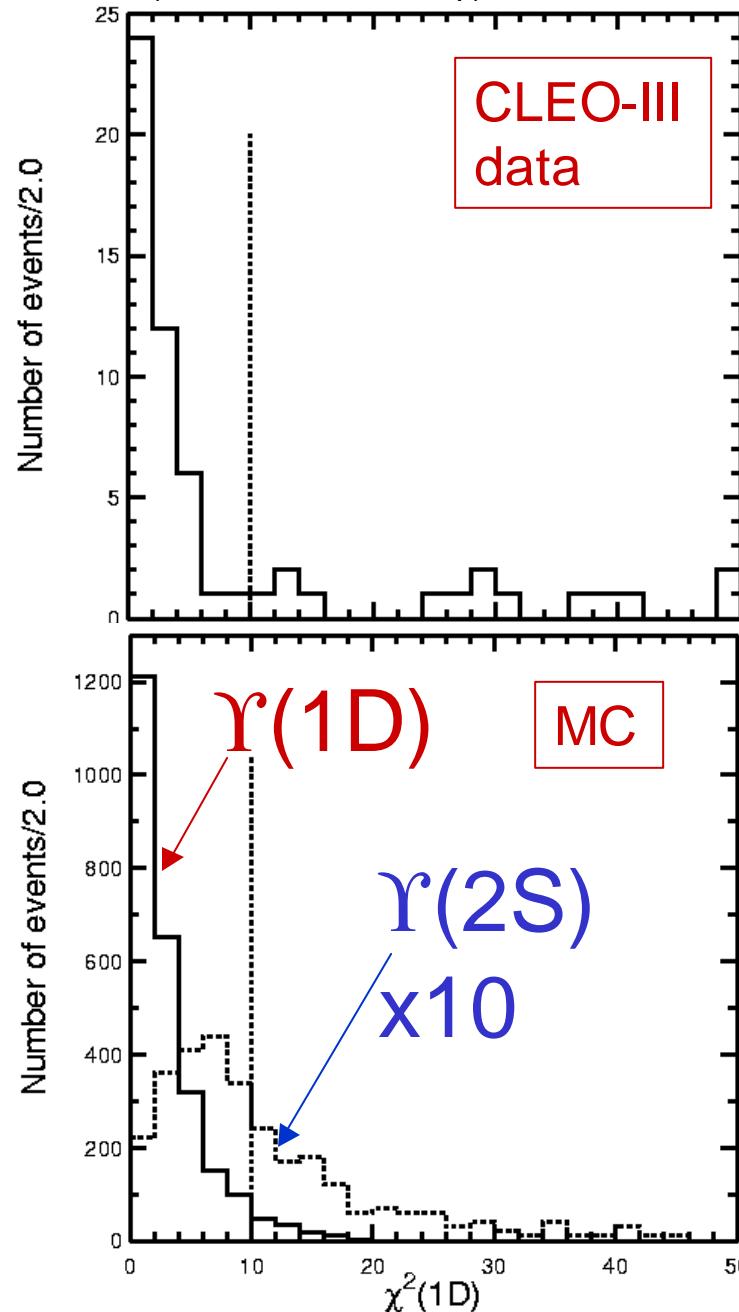
$$\mathbf{c}_{2S}^2 = \min_{J_{2P}, J_{1P}} \sum_{i=1}^4 \left(\frac{E_{g_i} - E_{g_i}^{\text{expected}}(M_{2S}, J_{2P}, J_{1P})}{s(E_{g_i})} \right)^2$$

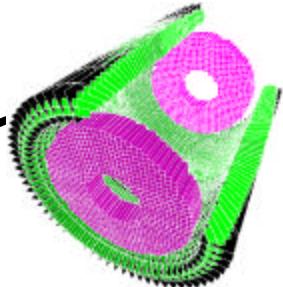
$$\mathbf{c}_{2S}^2 > 12$$



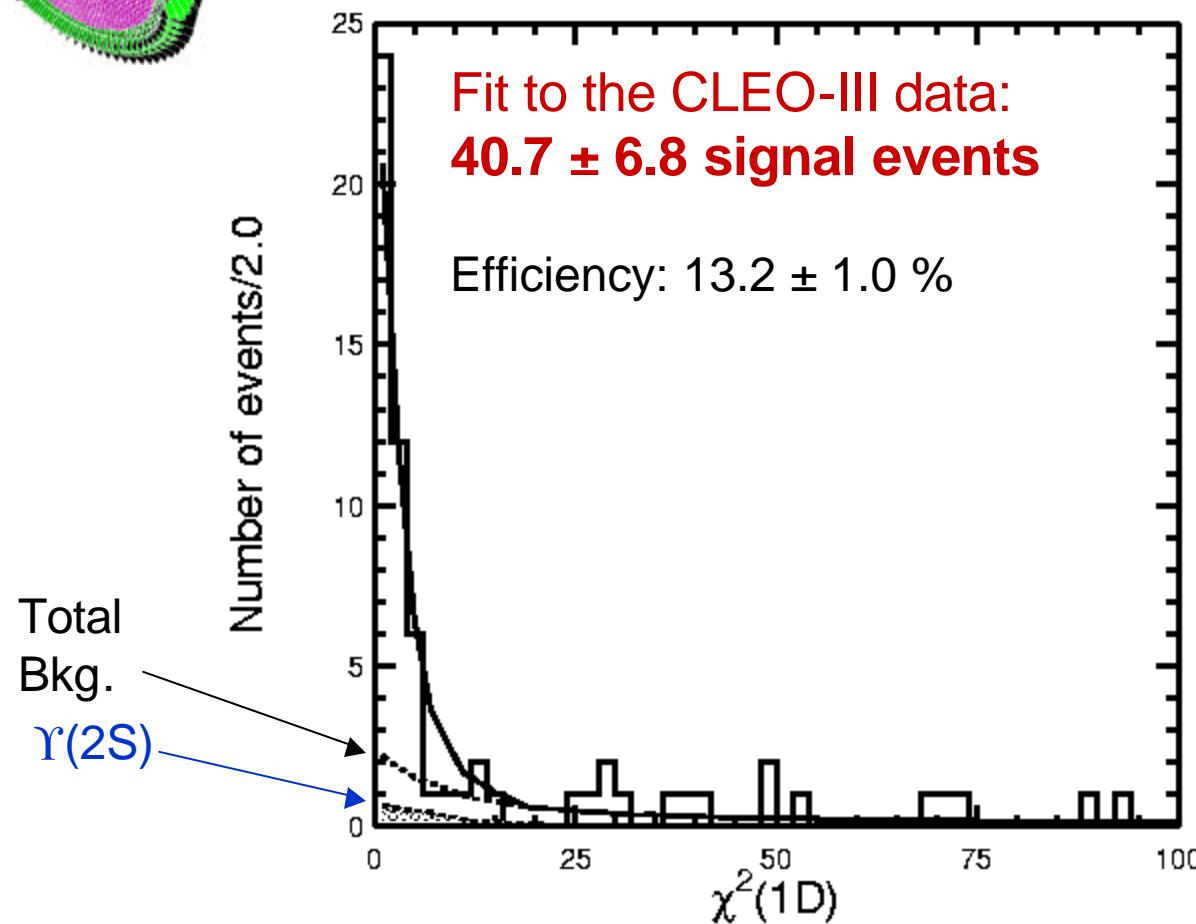
Inclusive $\Upsilon(1D)$ signal

- No background source can produce as narrow a peak as observed in the data
- For $\chi^2_{1D} < 10$:
 - 44 events in the data
 - 1.6-3.0 events due to $\Upsilon(2S)$
 - 0.8 events due to $\Upsilon(3S) \rightarrow \pi^0\pi^0\Upsilon(1S)$
 - 1.8-3.7 of other backgrounds (e.g. radiative Bhabhas and $\mu-$ pairs) estimated from the tail of the distribution
 - **Total background 10-14%**





Inclusive $\Upsilon(1D)$ signal



The signal significance estimated from the fits with the background contributions alone:

$\mu\mu+ee$: **9.7 σ**

$\mu\mu$: **8.2 σ**

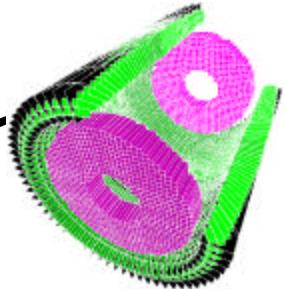
ee : **5.2 σ**

$$\mathcal{B}(\Upsilon(3S) \rightarrow gg\Upsilon(1D) \rightarrow gg\gamma\gamma \Upsilon(1S) \rightarrow gggg\ell^+\ell^-) = (3.3 \pm 0.6 \pm 0.5) 10^{-5}$$

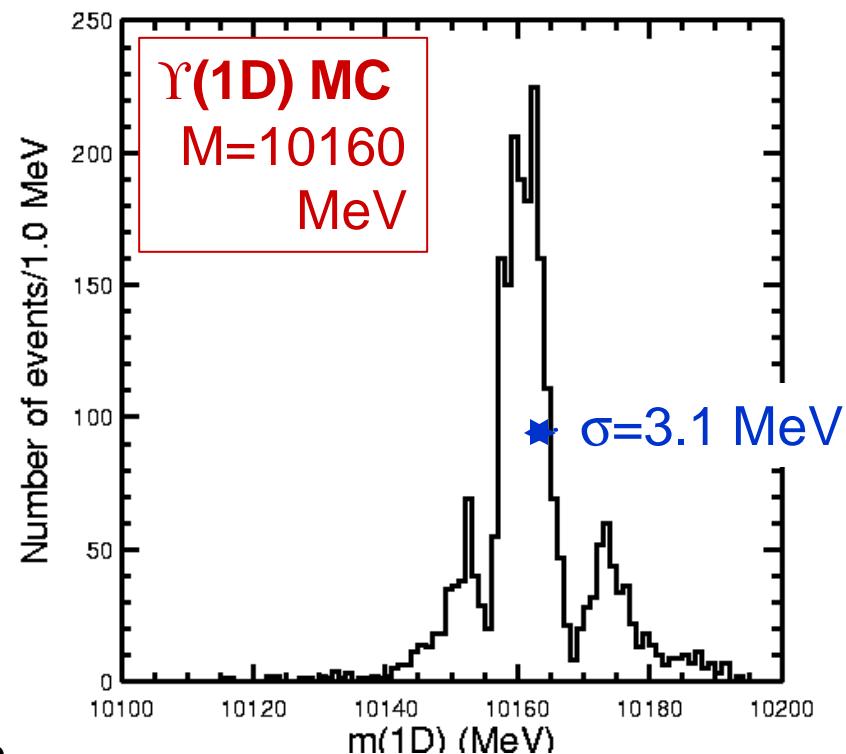
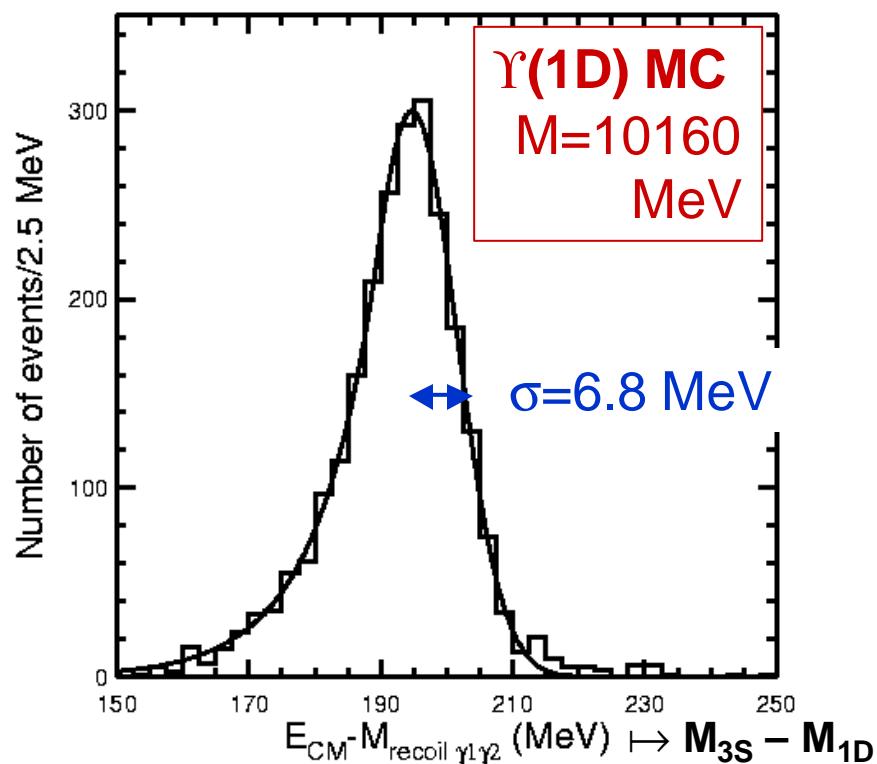
Godfrey&Rosner

3.8

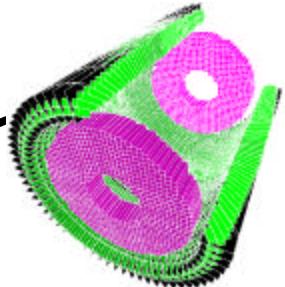
10^{-5}



$\gamma(1D)$ mass estimators

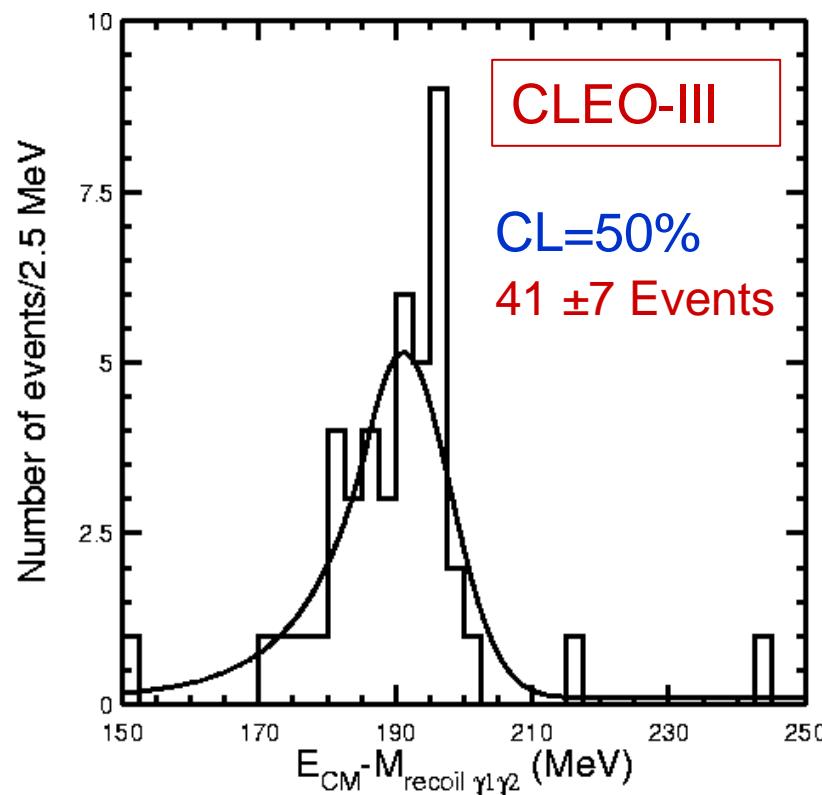


- Recoil mass against the two lowest energy photons:
 - Worse resolution
 - Simple shape
- Most likely mass (constrained to 2P,1P masses):
 - Better resolution
 - Satellite peaks due to wrong J_{2P}, J_{1P} minimizing the χ^2_{1D}

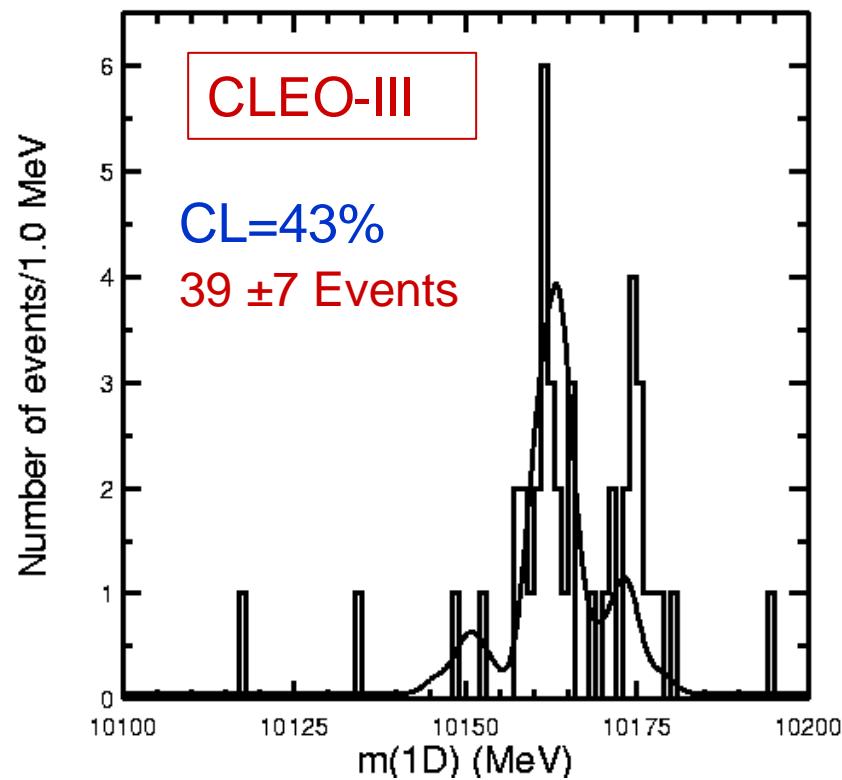


$\gamma(1D)$ mass analysis

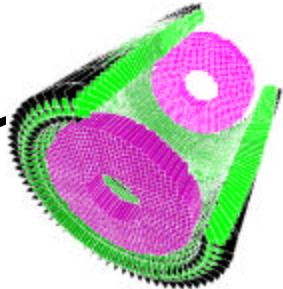
- Single-peak fits:



$$M = 10163.4 \pm 1.3 \text{ MeV}$$

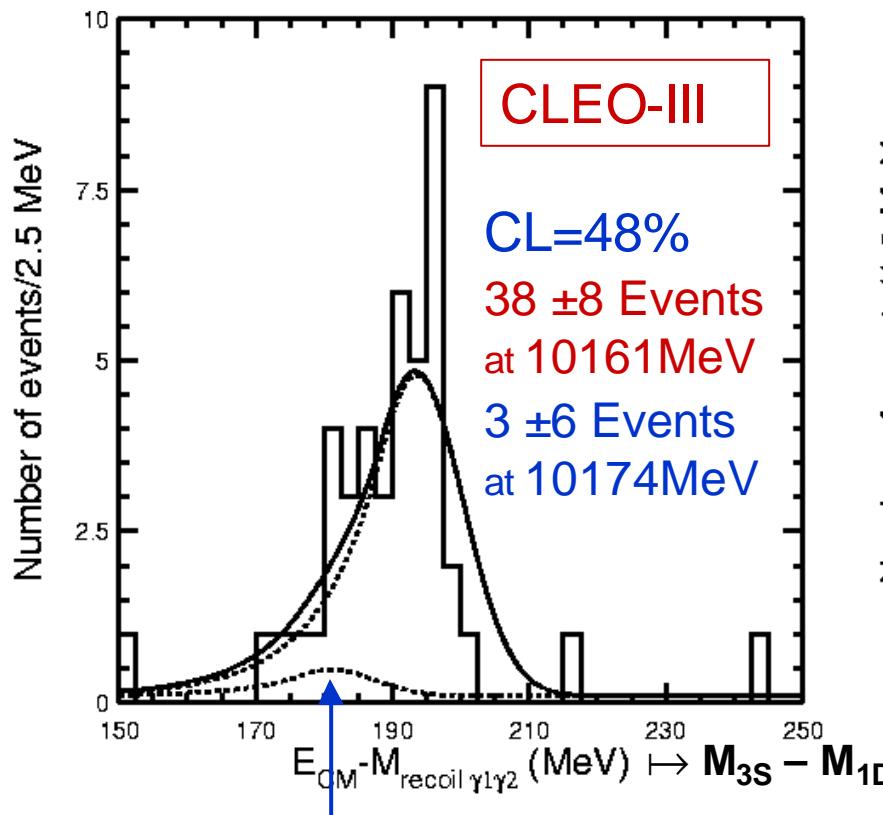


$$M = 10162.0 \pm 0.5 \text{ MeV}$$

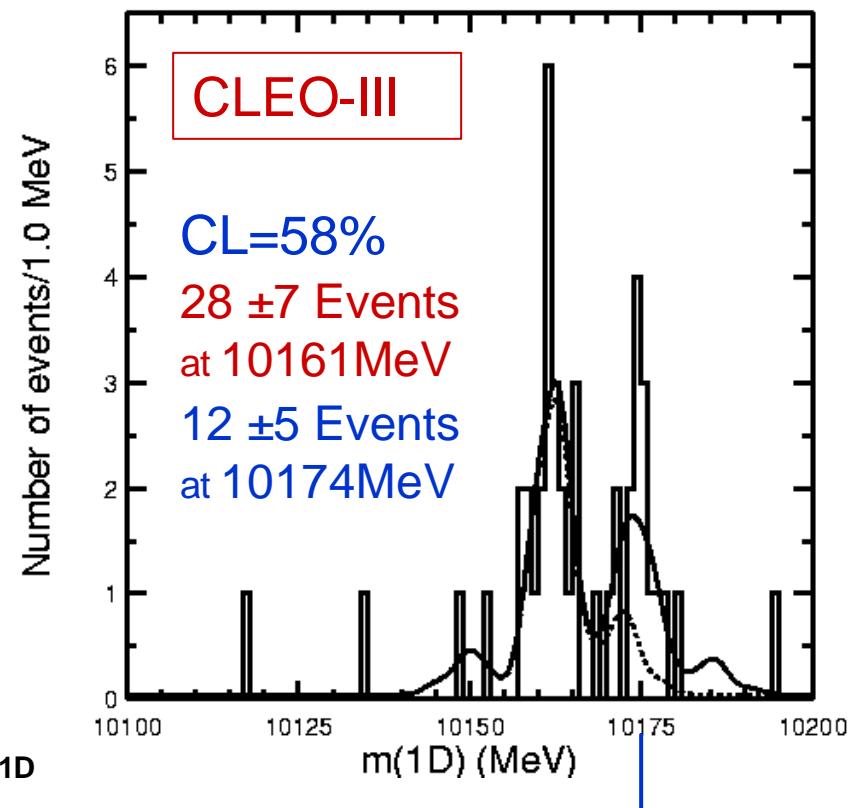


$\gamma(1D)$ mass analysis

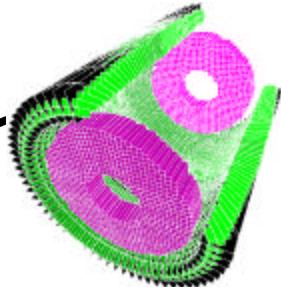
- Double-peak fits:



Masses fixed from the fit
shown on the right



$M=10161.2 \pm 0.7$ MeV



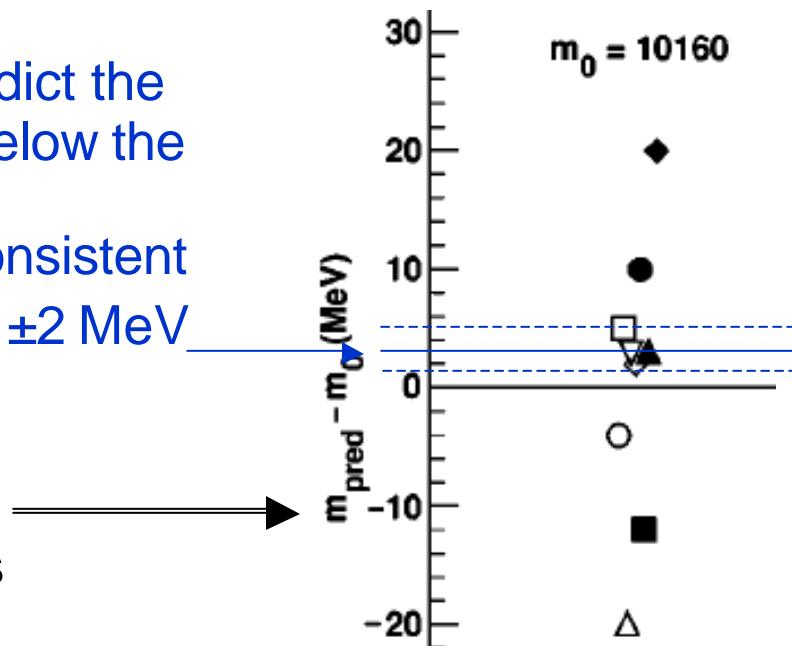
$\Upsilon(1D)$ mass analysis

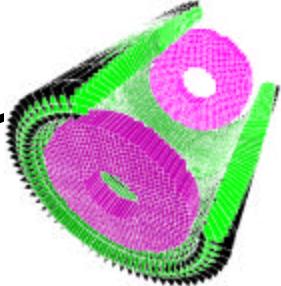
- No compelling evidence for more than one state
- Significance of the peak at 10162: **6.8s**
- Mass averaged over different fits: **10162.2 ± 1.6 MeV**
- Inconsistent with the $\Upsilon(1D_3)$
- Could be the $\Upsilon(1D_2)$ or $\Upsilon(1D_1)$
- The theory predicts the rate ratio: $\Upsilon(1D_2)/\Upsilon(1D_1)=6$
- **Thus, the $\Upsilon(1D_2)$ is the most likely interpretation**

All calculations of the fine splitting predict the $\Upsilon(1D_2)$ mass from -0.5 to -1.0 MeV below the center-of-gravity of the triplet

→ Our mass measurement is consistent with the c.o.g. $\sim 10163 \pm 2$ MeV

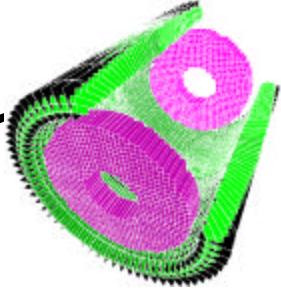
Spread in the predictions of the center-of-gravity of the triplet 1D states by various potential models
(from Godfrey&Rosner)





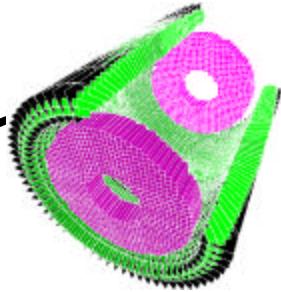
Summary

- No evidence for hindered M1 transitions $\Upsilon(3S) \rightarrow g h_b(1S)$ found in contradiction with many theoretical estimates of the transition width
- Much improved results for:
 - $B(\Upsilon(3S) \rightarrow g c_b(1P_{2,1}) \rightarrow gg\Upsilon(1S))$
 - $B(\Upsilon(3S) \rightarrow g c_b(2P_{2,1,0}) \rightarrow gg\Upsilon(2,1S))$
 - $B(\Upsilon(3S) \rightarrow p^0 p^0 \Upsilon(1S))$
 - Upper limits on:
 - $B(\Upsilon(3S) \rightarrow p^0 \Upsilon(1S))$
 - $B(\Upsilon(3S) \rightarrow h\Upsilon(1S))$
 - $B(\Upsilon(3S) \rightarrow p^0 \Upsilon(2S))$



Summary

- First observation of $\Upsilon(1D)$:
 - Signal is **9.7s significant**
 - Inclusive (i.e. sum over all J) product branching ratio for production in $gggg\ell^+\ell^-$ **($3.3 \pm 0.6 \pm 0.5$) 10^{-5}**
 - In agreement with the prediction by Godfrey&Rosner ($3.8 \cdot 10^{-5}$)
 - Evidence for a state at **10162.2 ± 1.6 MeV**
 - Likely interpretation: $\Upsilon(1D_2)$
 - The mass is consistent with the predictions of some of the potential models
 - First new narrow bb state observed in 19 years
 - The only long-lived $L=2$ meson we know



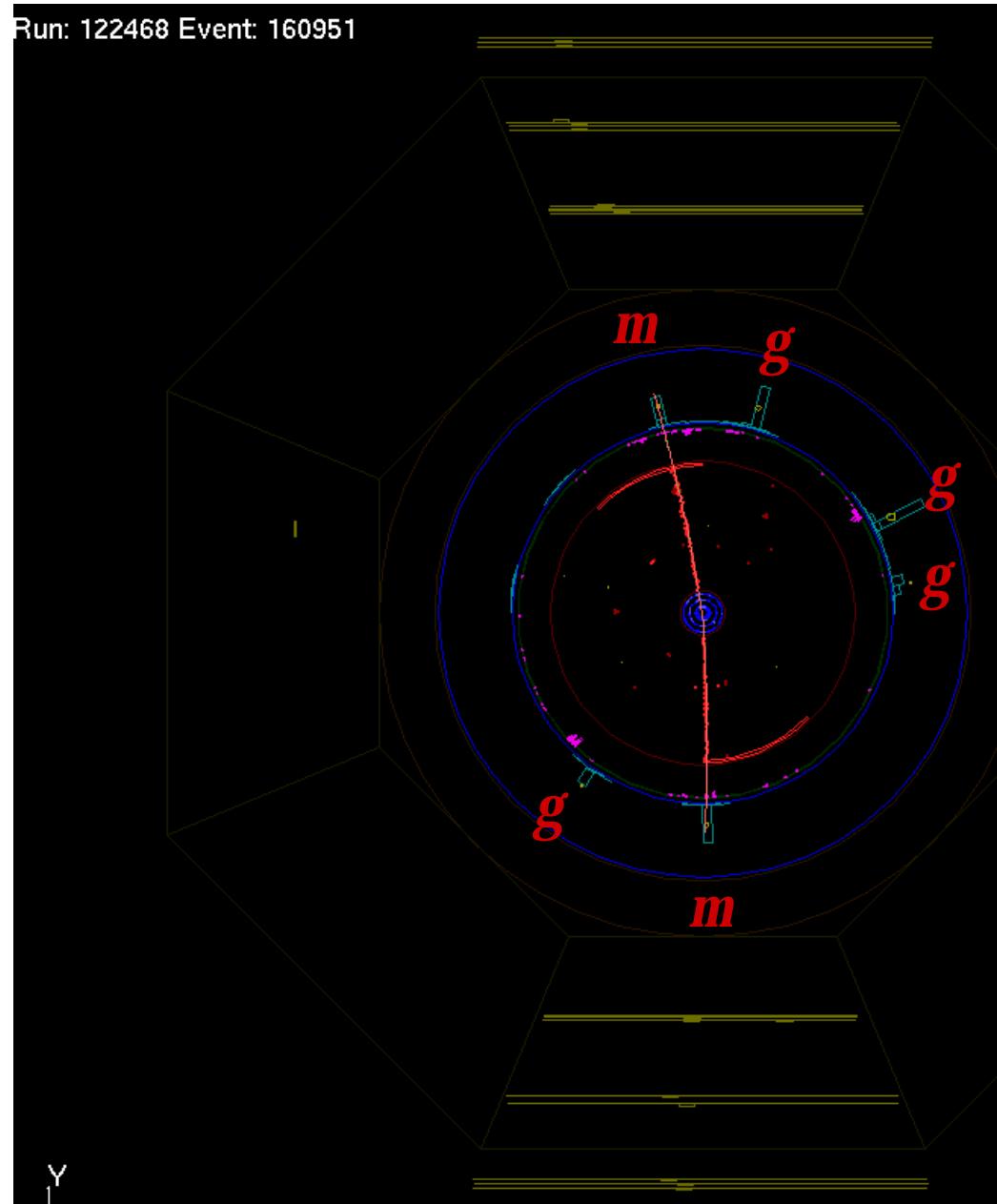
Tomasz Skwarnicki, Syracuse U.

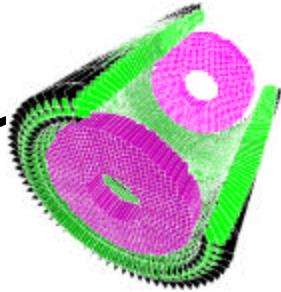
ICHEP, Amsterdam July,2002

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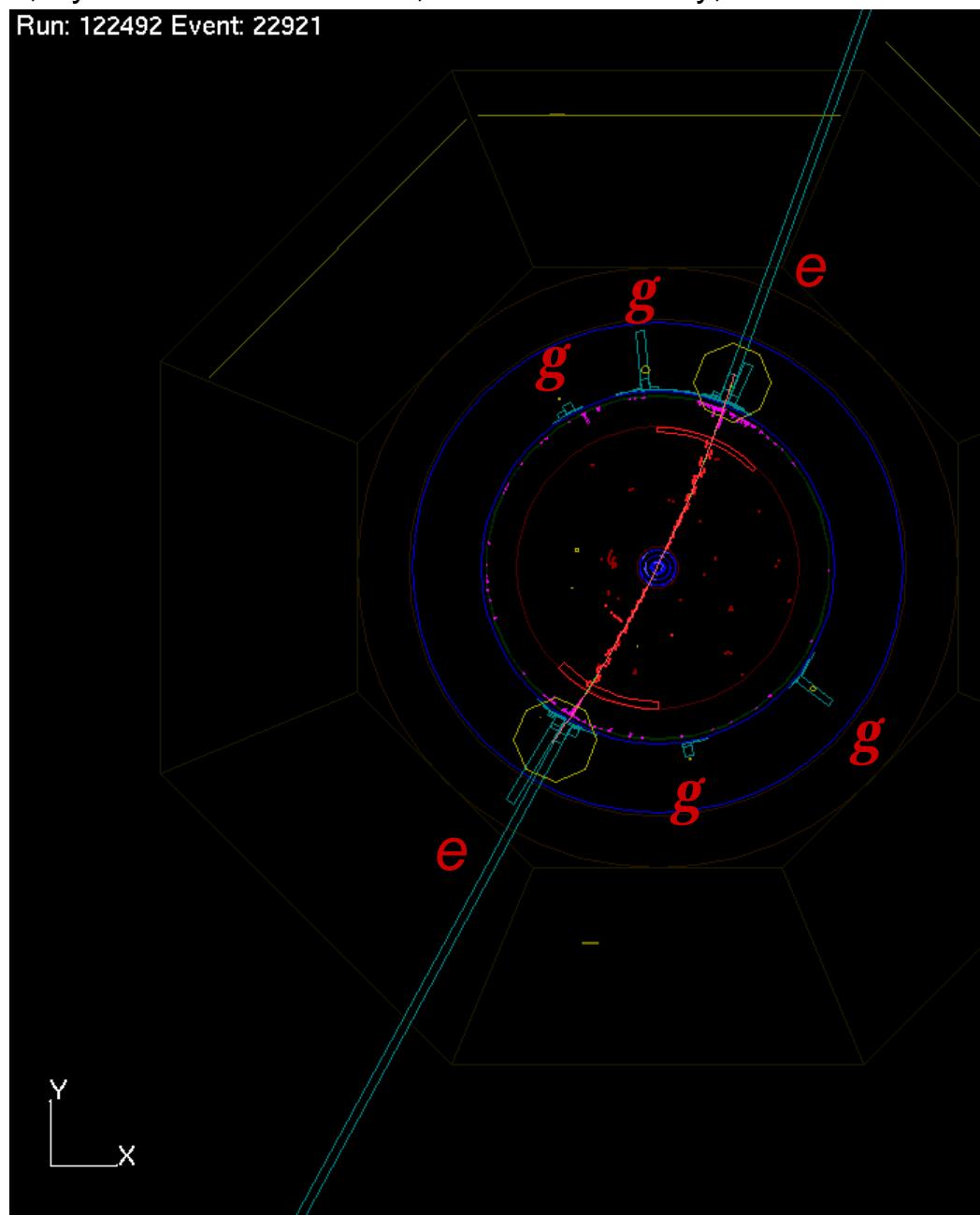
$\gamma\gamma\mu\mu$
1D candidate

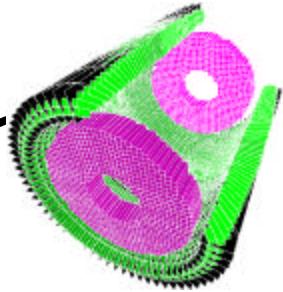




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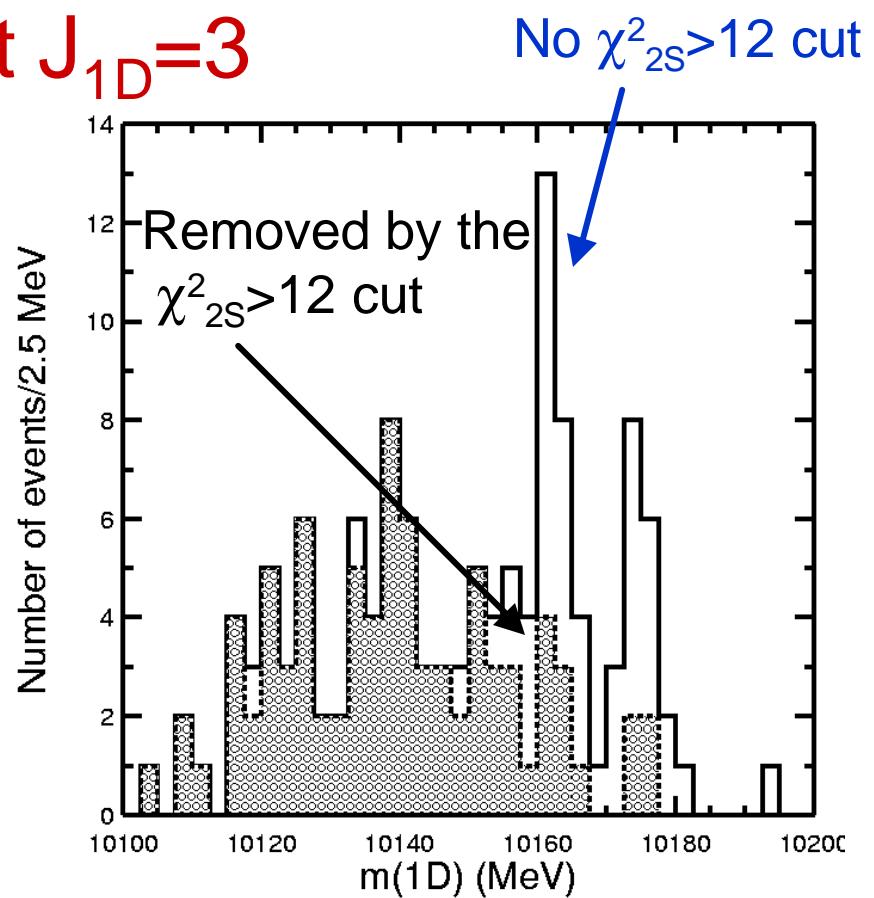
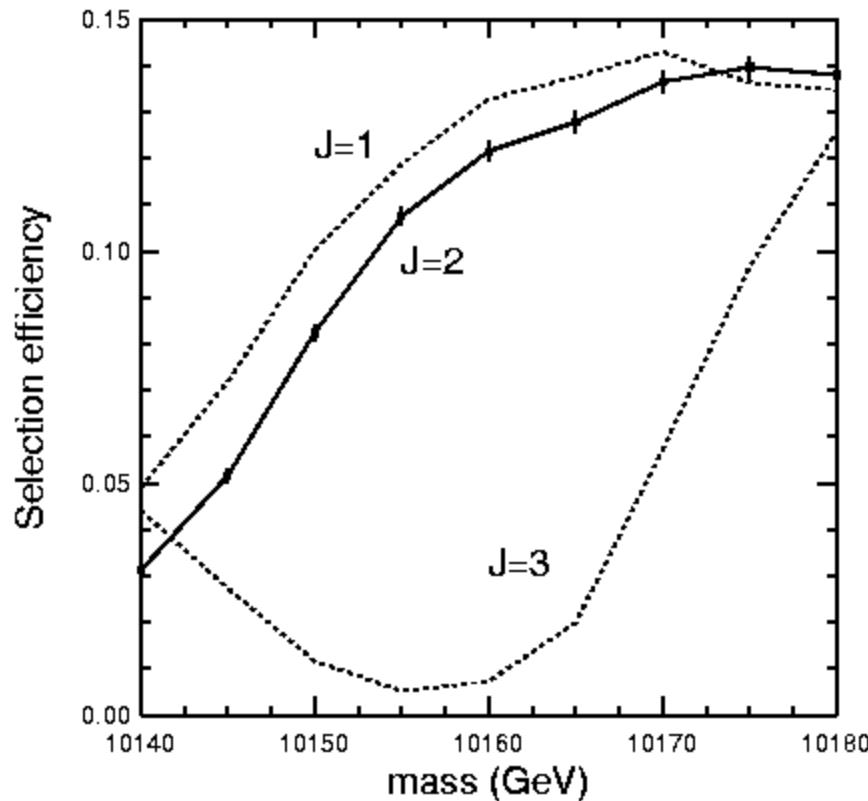
$\gamma\gamma\gamma ee$ 1D
candidate



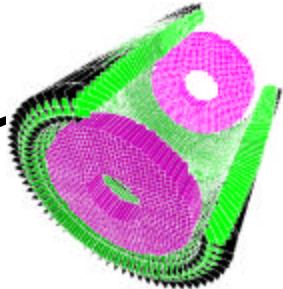


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Ruling out $J_{1D}=3$



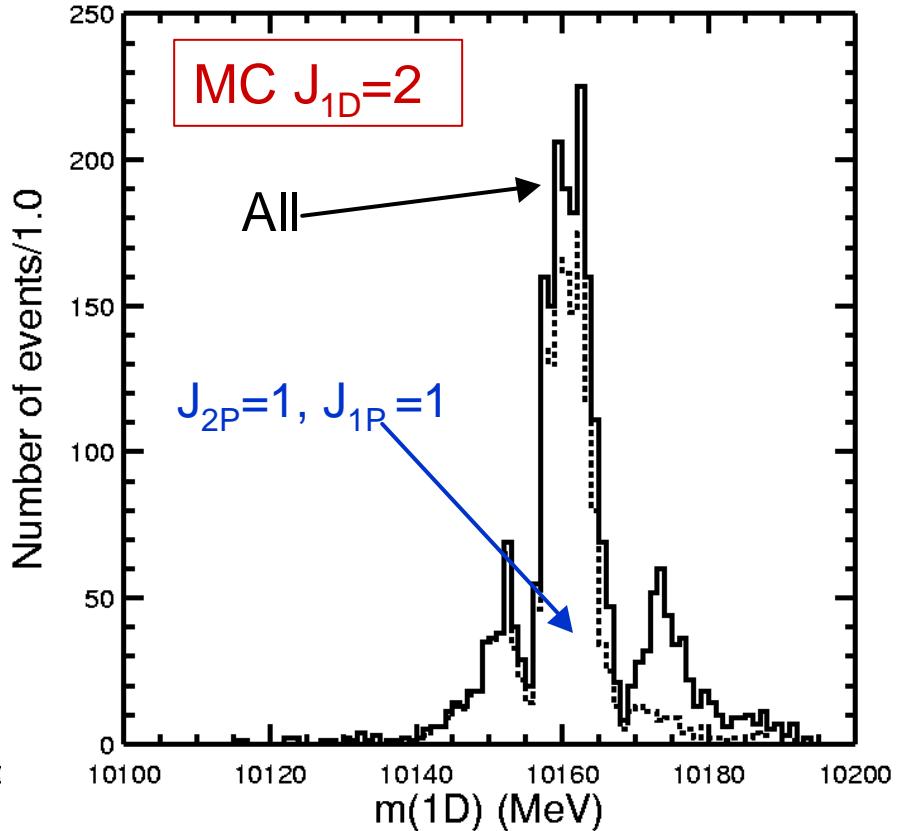
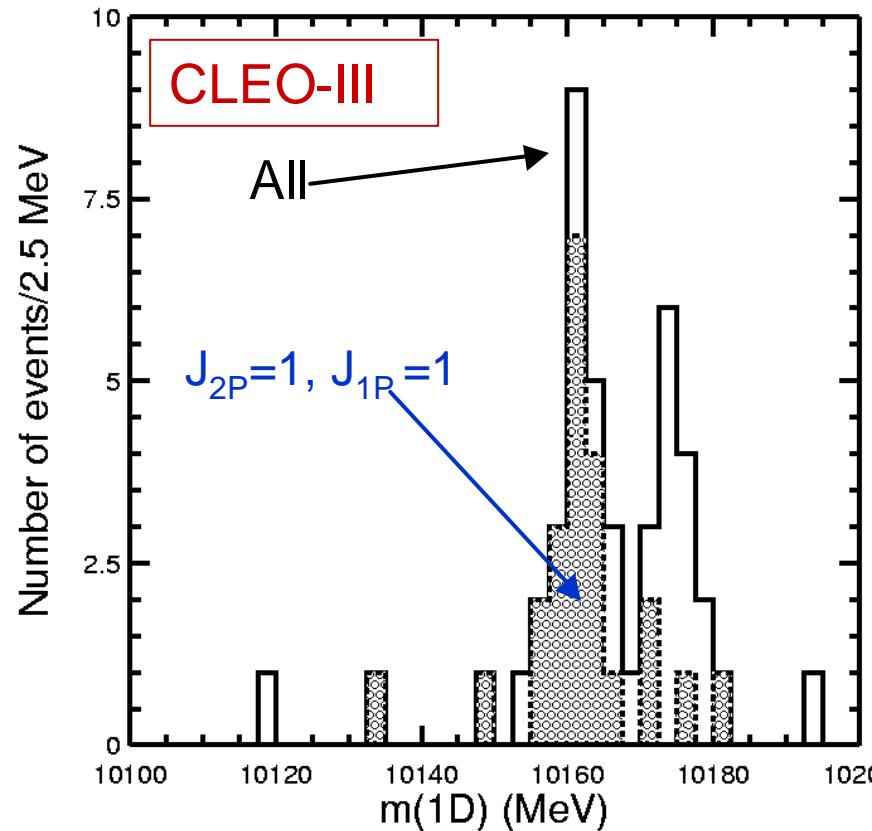
- No efficiency for the $\Upsilon(1D_3)$ at 10162 MeV because of the $\chi^2_{2S} > 12$ cut
- However the $\chi^2_{2S} > 12$ cut does not change the 10162 peak amplitude much

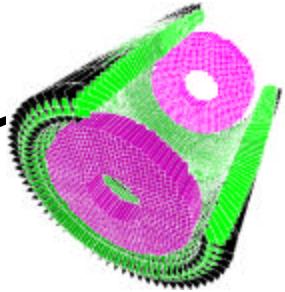


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Probing J_{1D} via J_{2P}, J_{1P}

- The peak at 10162 MeV has a large fraction of $J_{2P}=1$, $J_{1P}=1$ events, as expected for $J_{1D}=1$ or 2.

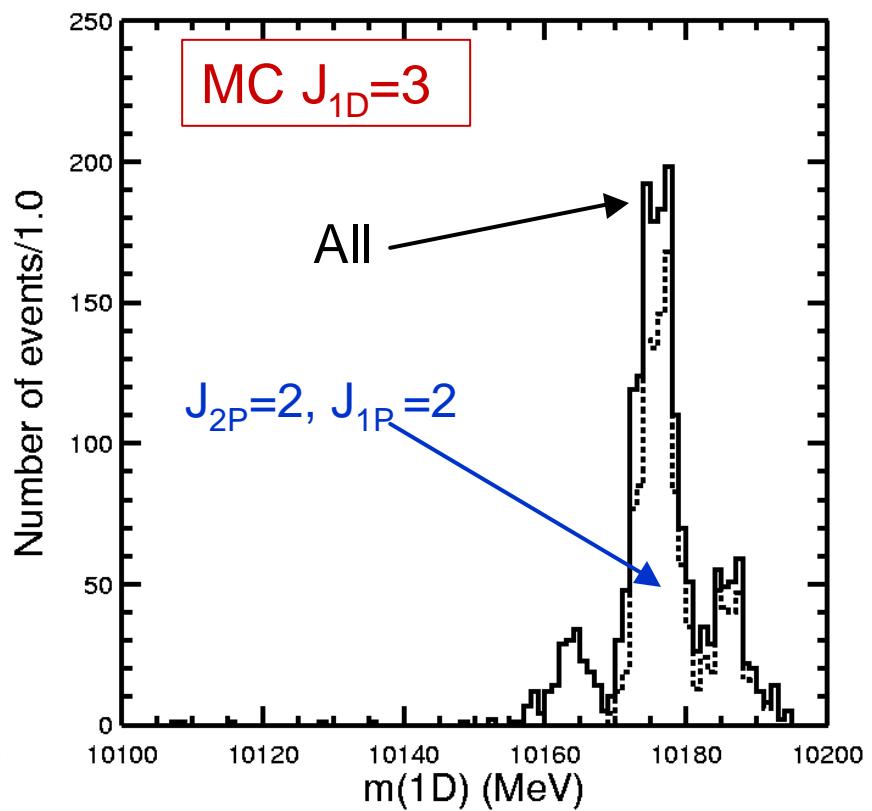
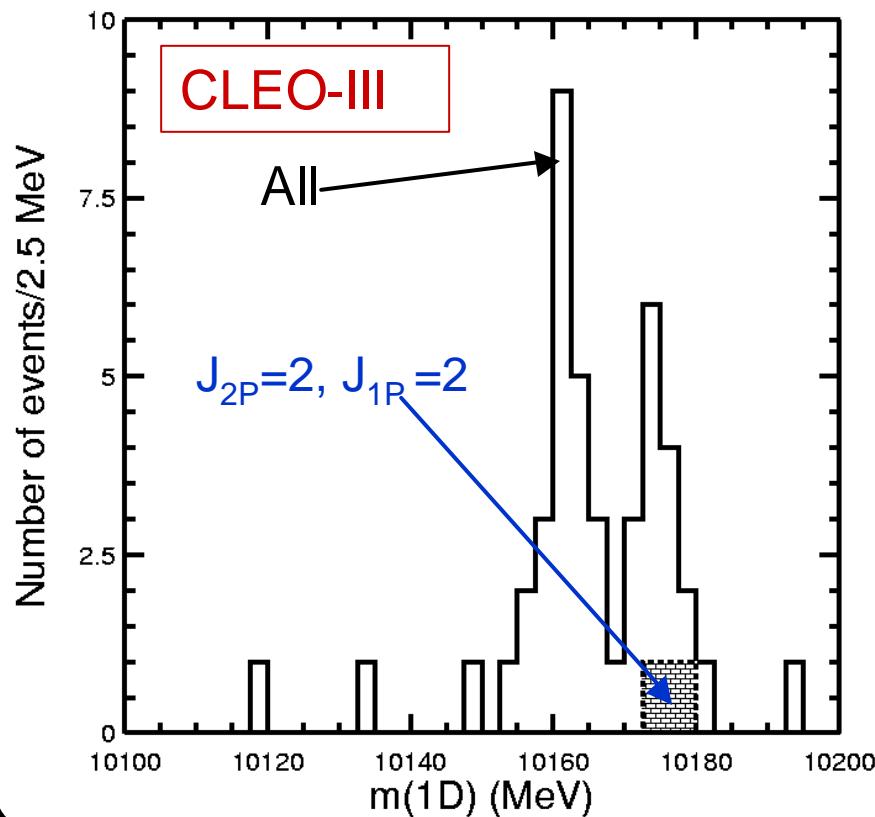


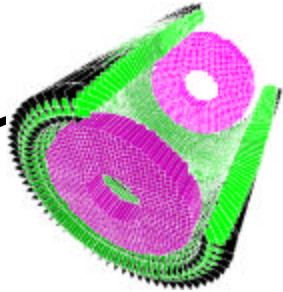


(Extra Slide)

Probing J_{1D} via J_{2P} , J_{1P}

- The peak at 10174 MeV has a too small a fraction of $J_{2P}=2$, $J_{1P}=2$ events, to be $J_{1D}=3$ (in fact a few such events expected due to the satellite from $J_{1D}=2$ at 10162 MeV)





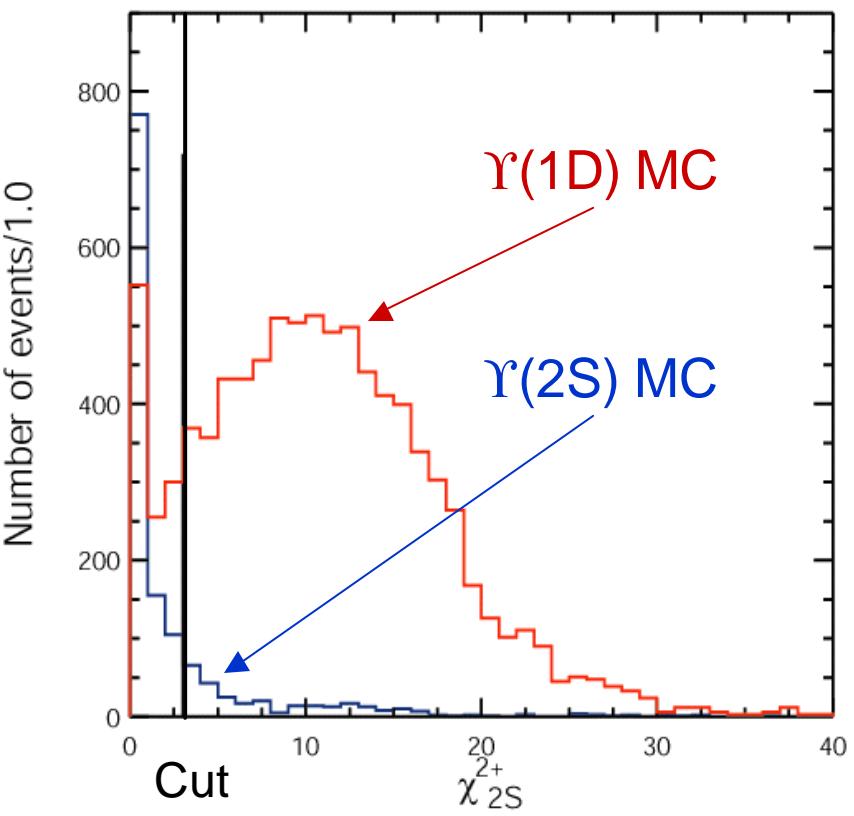
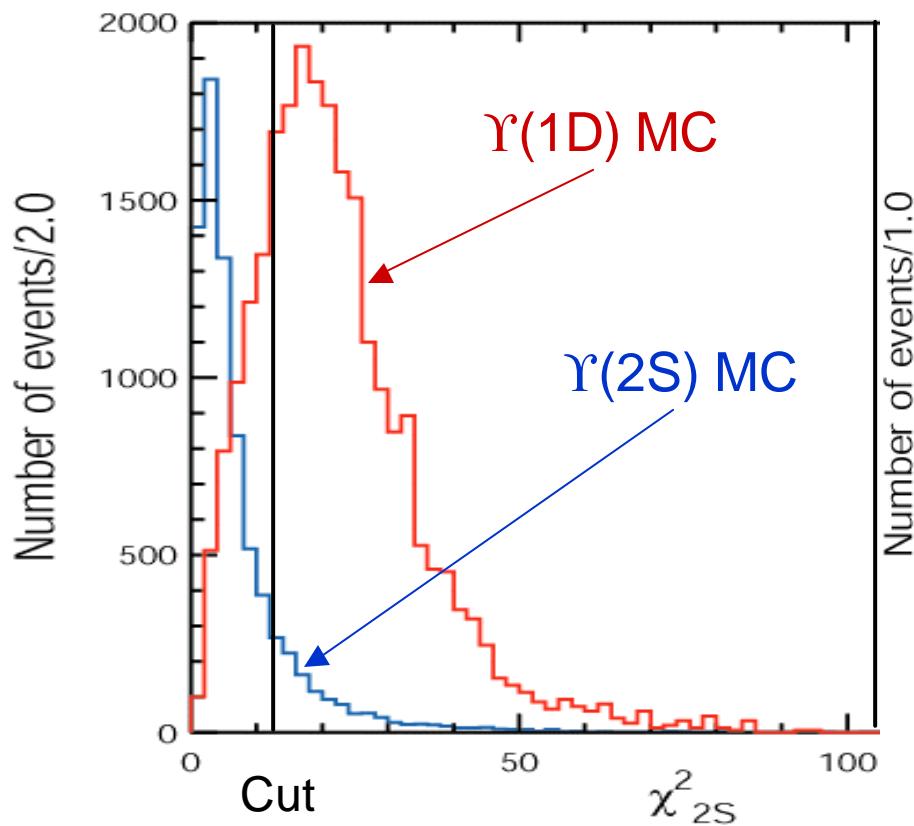
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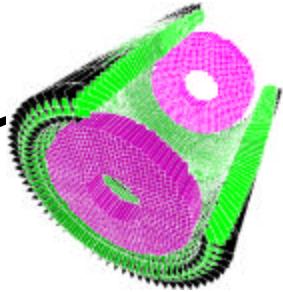
Suppression of the $\Upsilon(2S)$ 4 γ -cascades

$$c_{2S}^2 = \min_{J_{2P}, J_{1P}} \sum_{i=1}^4 \left(\frac{E_{g_i} - E_{g_i}^{\text{expected}}(M_{2S}, J_{2P}, J_{1P})}{s(E_{g_i})} \right)^2 \quad c_{2S}^{2+} = \min_{J_{2P}, J_{1P}} \sum_{i=1}^4 \left(\frac{\max\{0, E_{g_i} - E_{g_i}^{\text{expected}}(M_{2S}, J_{2P}, J_{1P})\}}{s(E_{g_i})} \right)^2$$

$c_{2S}^2 > 12$

$c_{2S}^{2+} > 3$





(Extra Slide)

Signal for $\Upsilon(3S) \rightarrow \gamma \chi_b(2P_0) \rightarrow \gamma \gamma \Upsilon(2S)$

- Projection of 2D-fit to $\delta = (\text{Recoil Mass}(2\gamma) - M_{\Upsilon(2S)})/\sigma$ vs. E_γ^{Low} onto E_γ^{Low} with the $\exp(-\delta^2/2)$ as weight

