

Recent results from the K2K experiment

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for the K2K collaboration

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- Introduction
- Summary of the results in 2001
- Overview of the new analysis
(Number of events + Spectrum)
- Flux measurements at KEK
- Oscillation analysis and results
- Summary

K2K Collaboration

**High Energy Accelerator Research Organization(KEK)
Institute for Cosmic Ray Research(ICRR), University of Tokyo
Kobe University
Kyoto University
Niigata University
Okayama University
Tokyo University of Science
Tohoku University**

**Chonnam National University
Dongshin University
Korea University
Seoul National University**

**Boston University
University of California, Irvine
University of Hawaii, Manoa
Massachusetts Institute of Technology
State University of New York at Stony Brook
University of Washington at Seattle**

**Warsaw University
Solton Institute**

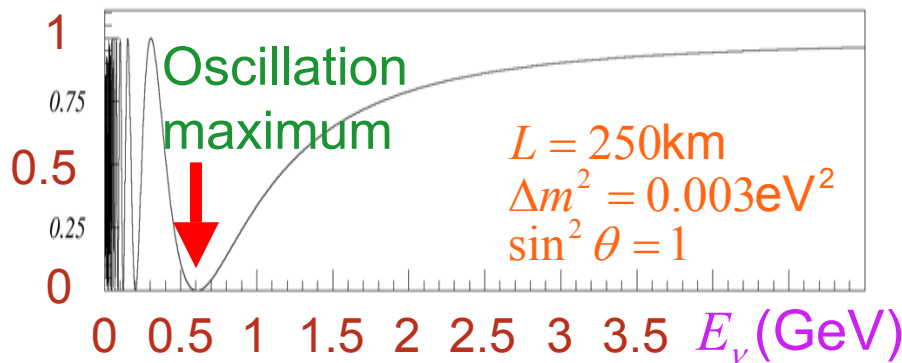
Introduction

Principle of the long baseline experiment

$$\text{Oscillation Probability} = \sin^2 2\theta \cdot \sin^2 \left(\frac{1.27 \Delta m^2 L}{E_\nu} \right)$$

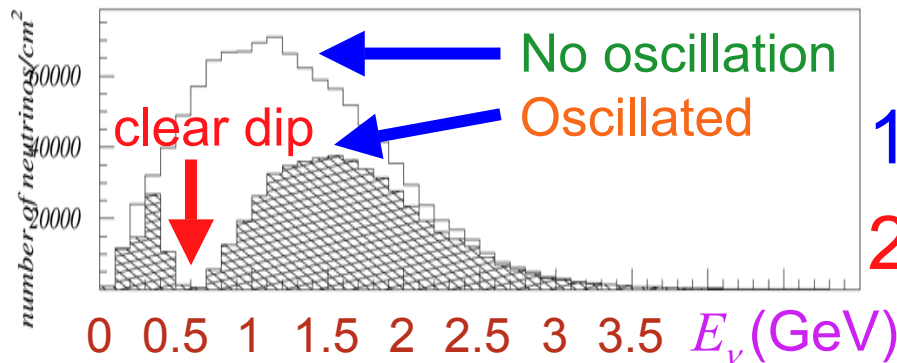
(2 flavor)

Probability



long baseline experiment
→ **Fixed distance**

of interactions



Compared to
the NULL oscillation case,

- 1) Reduced number of events
- 2) Distorted energy spectrum
dip around 0.6 ~ 0.7 GeV

The K2K experiment

Super Kamiokande
(far detector)
50kt water Cherenkov detector



12GeV PS@KEK
 ν beamline
beam monitors
near detectors

Neutrino beam

$$\langle E_\nu \rangle \sim 1.3 \text{ GeV}$$

almost pure ν_μ ($\sim 98\%$)

Beam monitors & near detectors

- beam direction
- π monitor
- ν_μ flux & spectrum

Far detector

- ν_μ flux & spectrum

ν_μ disappearance
spectrum distortion

Neutrino beamline

Front (Near) Detector

direction (ν)

spectrum, rate

μ -monitor

direction ($p \rightarrow \mu$)

Front detector

μ -monitor



North
Counter
Hall

12 GeV PS

Target
station

Al target

Decay
section
($\pi \rightarrow \mu \nu_\mu$)

200m

p

Primary beam line

12 GeV PS

fast extraction

every 2.2sec

beam spill 1.1 μ s

$\sim 6 \times 10^{12}$ protons/spill

Double Horn(250kA) $\sim 20 \times$ flux

Pion monitor

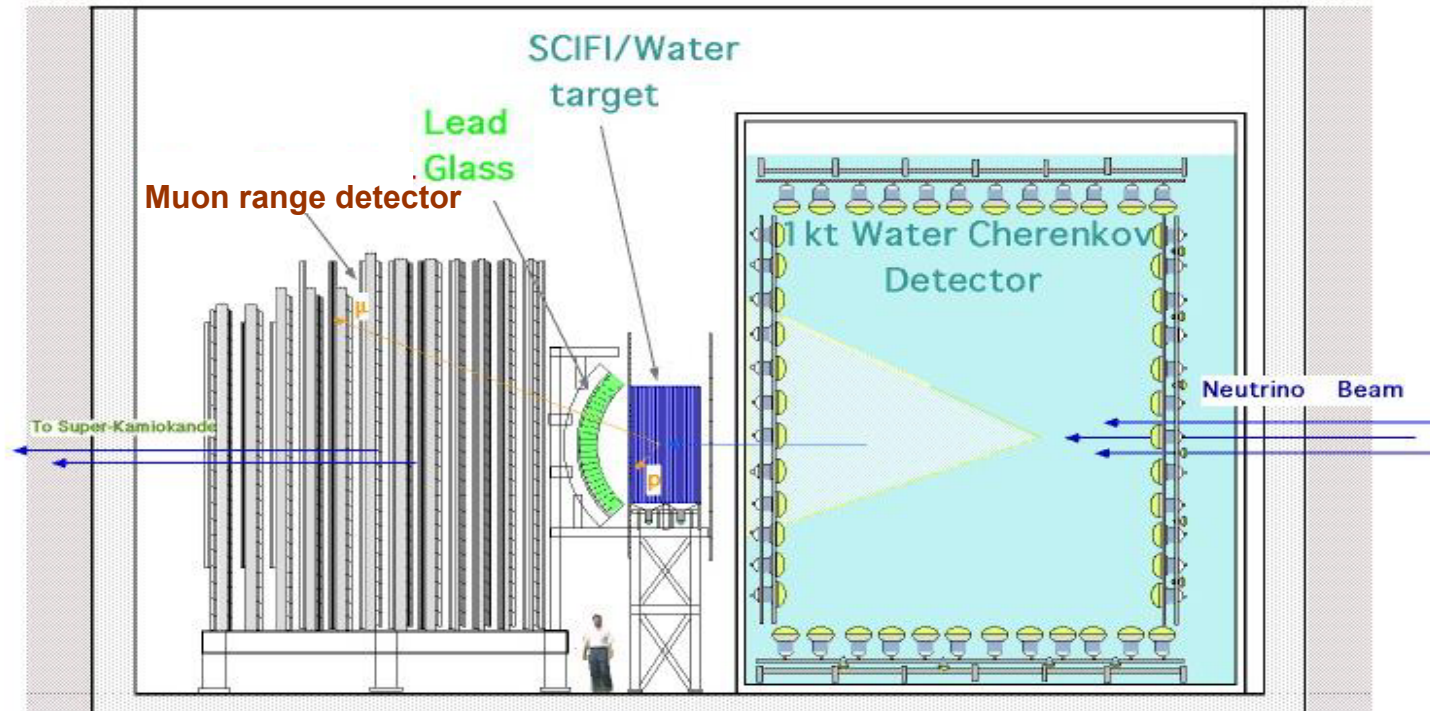
(P_π, θ_π after Horn)

Near to Far flux ratio

R_{FN}



Near neutrino detectors



1 kt water Cherenkov detector (1 kt) water target (25t fid. vol.)
same type as SK

Fine grained detectors

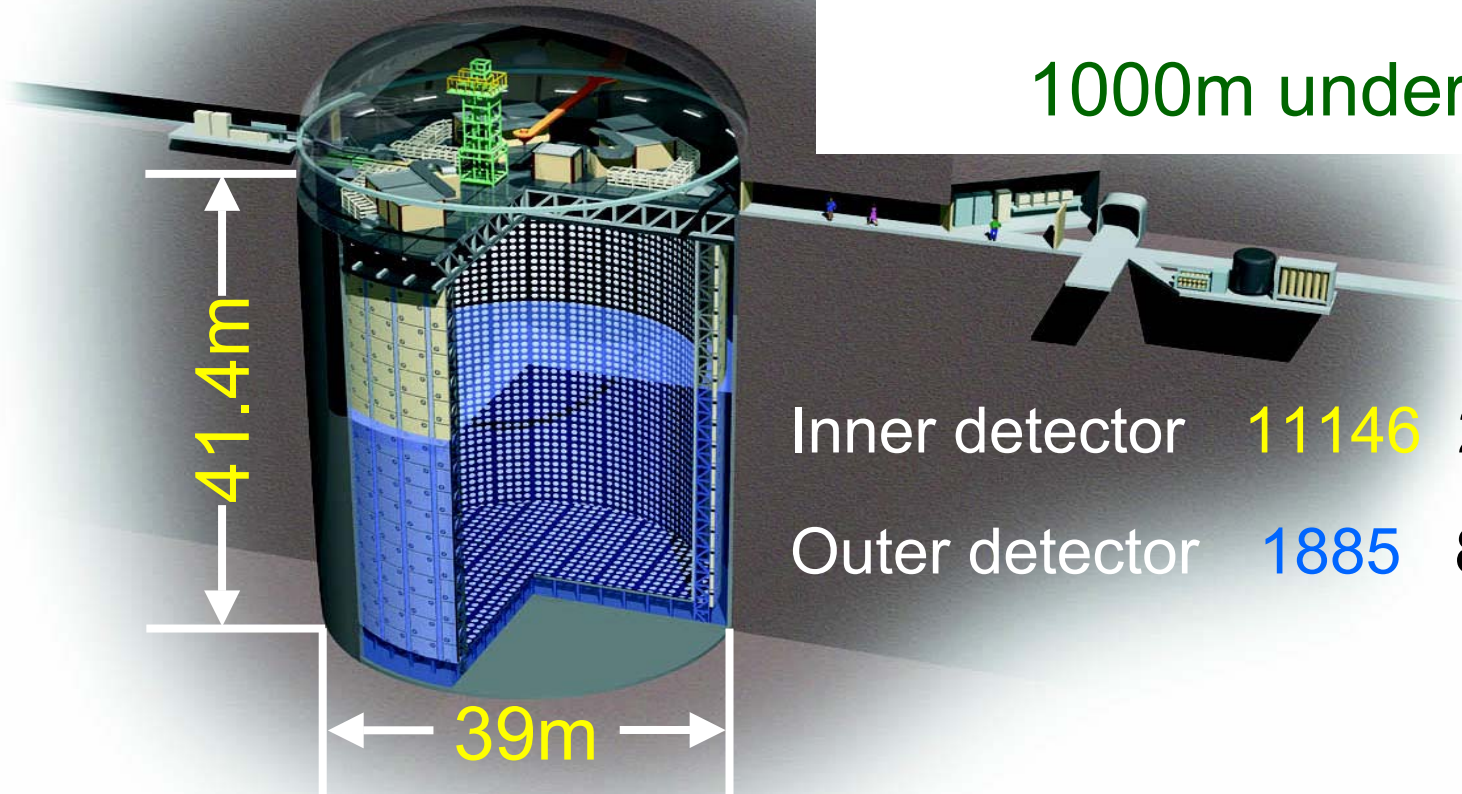
- **Scintillating fiber tracker** (SciFi) water target (6t fid. vol.)
CCQE identification
- **Muon range detector** (MRD) Iron target (330t fid. vol.)
 ν beam monitor (mom. & dir.)

Super-Kamiokande (Far detector)

50kt water Cherenkov detector

Fiducial volume 22.5kt

1000m under the ground



Inner detector 11146 20" PMTs

Outer detector 1885 8" PMTs

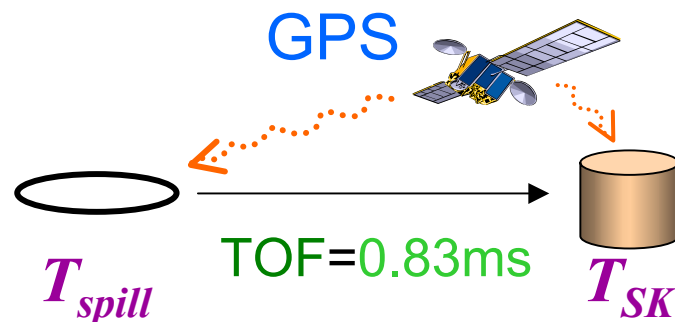
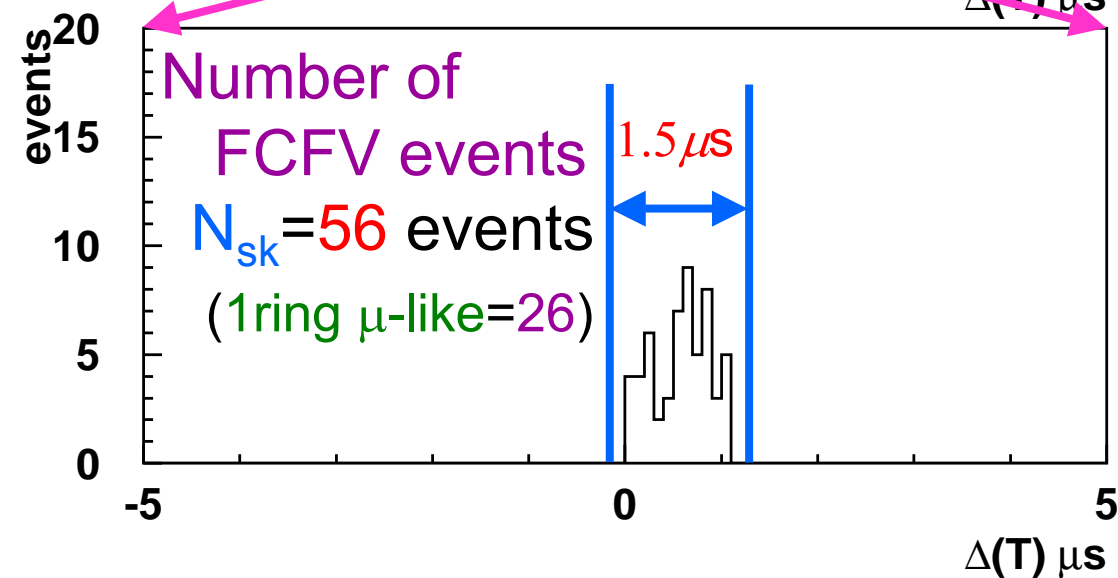
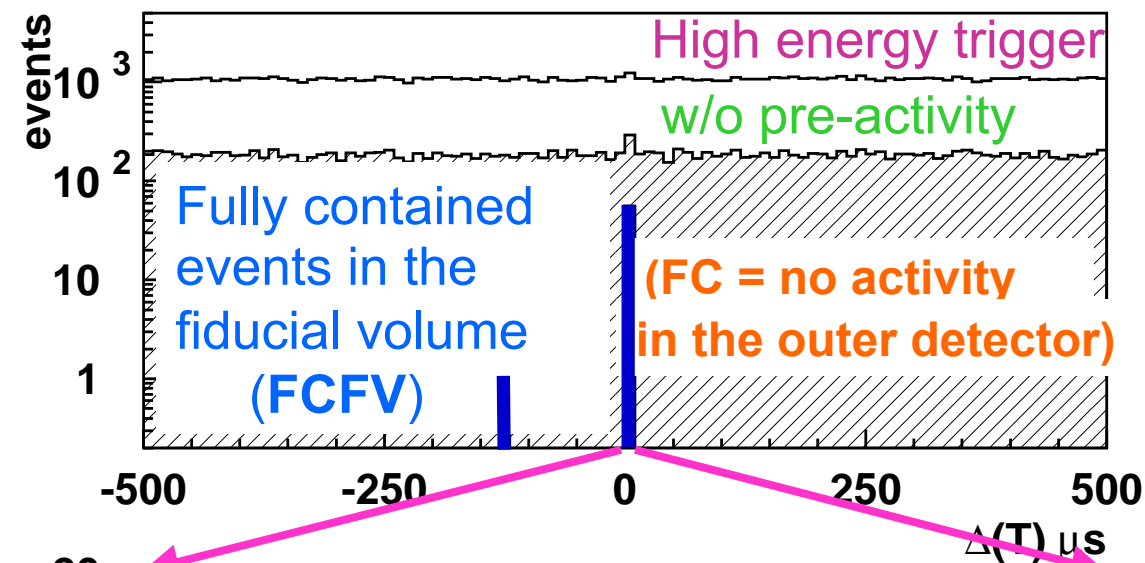
Atmospheric ν ~ 8 events/day (FCFV)

→ accidental coincidence $\sim 10^{-5}$ events/day

Event selection at Super-Kamiokande

From June '99 to July '01 (4.8×10^{19} protons on target)

FC



$$\Delta T \equiv T_{SK} - T_{spill} - TOF$$

requiring

$$-0.2 \leq \Delta T \leq 1.3 \mu s$$

Expected number of atmospheric ν BG

$$< 10^{-3} \text{ events}$$

Summary of K2K results in 2001

From June '99 to July '01

accumulated number of protons

4.8×10^{19} POT for the analysis

Neutrino beam was very stable

- **direction of the beam** : controlled less than **1mrad**.

confirmed by μ profile monitor ($\pi \rightarrow \mu$ decay)

muon range detector (MRD)

(ν interaction vertex)

- **energy spectrum of ν**

confirmed by 1kt water Cherenkov detector (1kt) & MRD

Summary of K2K results in 2001

Neutrino flux @ SK

estimated by Monte-Carlo

confirmed by π monitor

normalization was done by 1kt water Cherenkov detector

[number of protons was measured by Current Transformer (CT)]

of FCFV events in Super-K

Observed : 56 \longleftrightarrow Expected : $80^{+7.3}_{-8.0}$

\longrightarrow Probability of null oscillation $< 3\%$

\longrightarrow Next step

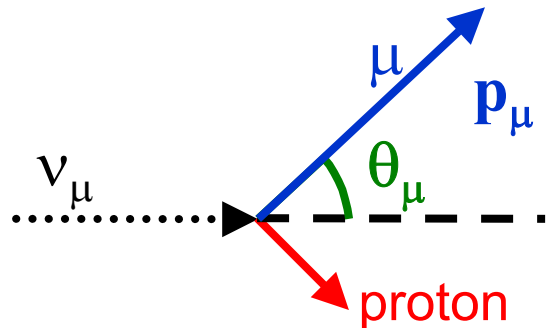
Full and Improved error estimations

and spectrum shape analysis

Neutrino interactions around 1GeV

Charged current quasi-elastic scattering (QE)

E_ν can be reconstructed from \mathbf{P}_μ and θ_μ .



\mathbf{p}_μ : muon momentum

θ_μ : muon angle

1kt : **single ring μ -like FCFV** events

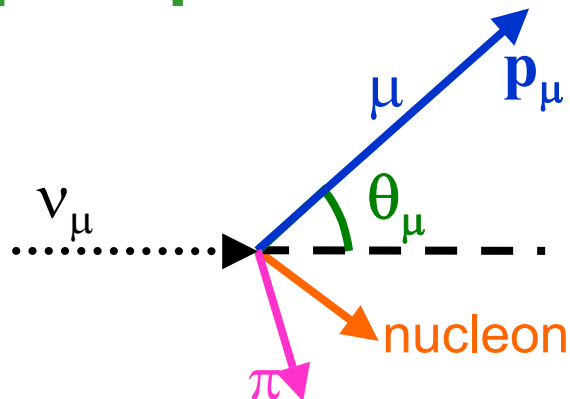
SciFi: **QE-like 2track** events

(when protons are identified)

or single track events

(when protons are not observed)

pion productions



1kt : **FCFV** events

(single-ring or multi-rings)

SciFi: **non-QE-like** events

Flow of the oscillation analysis

Observed quantities at the near detectors

$(\mathbf{P}_\mu, \theta_\mu)$ for each event category

Neutrino interaction models

Obtain neutrino spectrum at near detectors

Near to Far extrapolation ($\mathbf{R}_{\text{FN}}(\mathbf{E}_\nu)$)

Predict neutrino spectrum without oscillation ($\phi_{\text{SK}}(\mathbf{E}_\nu)$)

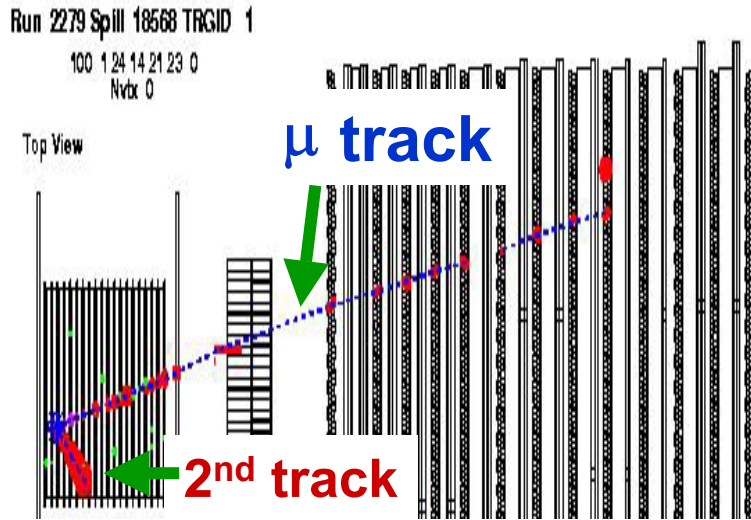
Observables at SK

- number of events (\mathbf{N}_{SK})
- Reconstructed energy of ν ($\mathbf{E}_\nu^{\text{rec}}$)

→ Fit the observed results at SK with ($\sin^2 2\theta, \Delta m^2$)

Use maximum likelihood fit

QE and non-QE events in SciFi 2track events

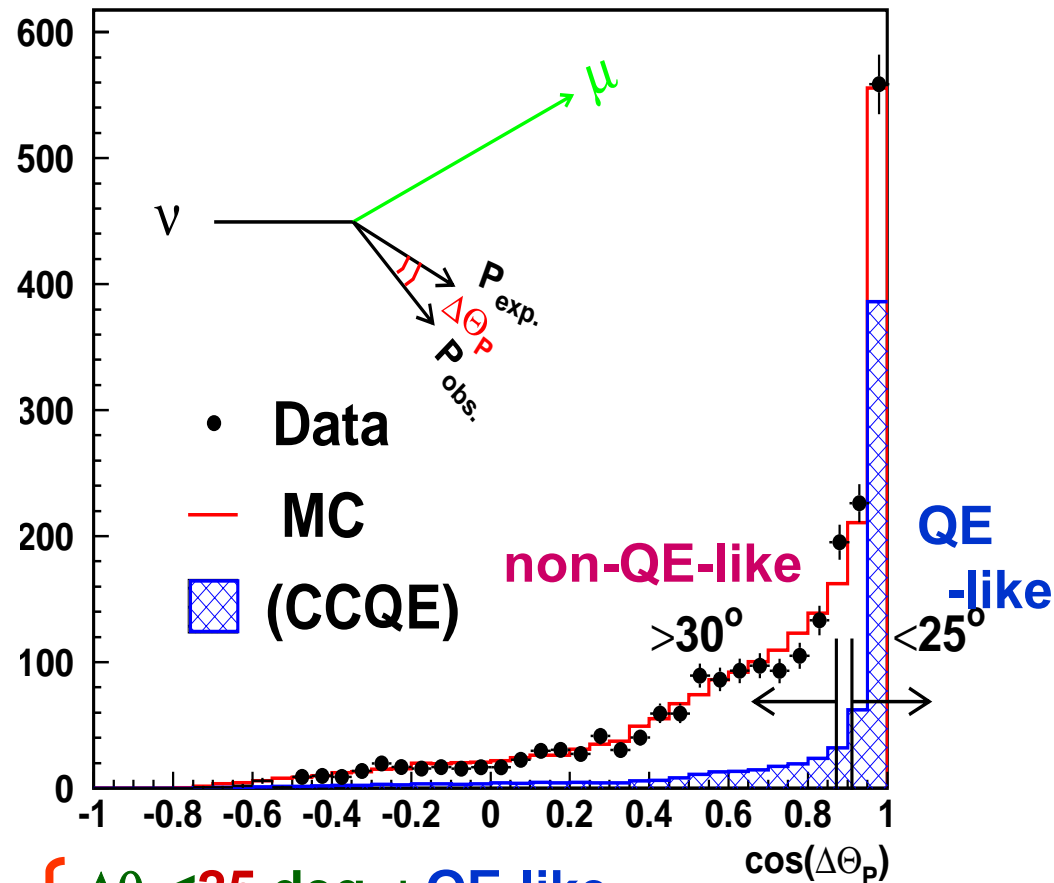


QE/non-QE selection

1. Select **2track** events
2. Select **muon track**
3. Calculate expected direction of proton θ_{exp} (assuming **QE** interaction)
4. **Compare** with the observed direction of 2nd track θ_{obs}

$$\Delta\theta_p = |\theta_{\text{exp}} - \theta_{\text{obs}}|$$

SciFi 2 track $\cos(\Delta\theta_p)$ distribution



$$\left\{ \begin{array}{l} \Delta\theta_p < 25 \text{ deg. : QE-like} \\ \Delta\theta_p > 30 \text{ deg. : non-QE-like} \end{array} \right.$$

Used data for $\phi_{\text{near}}(E_\nu)$

KT

**Fully Contained Fiducial
Volume (FCFV) events**

(0) No. of events

(Evis > 100 MeV)

(1) **Single μ -like events**

→ **4 sets of (p_μ, θ_μ) distributions**

SciFi

(2) **1-track μ events**

(3) **2-track QE-like events**

(4) **2-track non QE-like events**

Pion monitor & Beam simulation

π distribution in (p_π, θ_π)

→ **flux estimation $\phi_{\text{near}}(E_\nu)$ w. error**

ν flux $\phi_{\text{near}}(E_\nu)$ (8 bins)

ν interaction model

(parameterized as **QE/non-QE** ratio)

Fitting method (ν flux at KEK)

Prepare MC templates

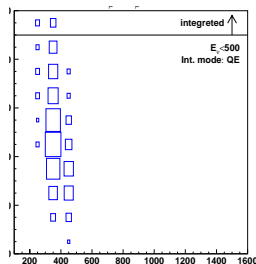
neutrino spectrum $\phi(E_\nu)$

and ν interaction model

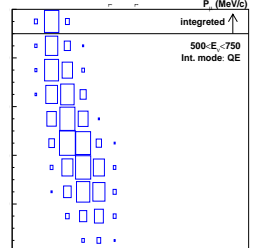
(QE/non-QE ratio).

E_ν QE(MC) non-QE(MC)

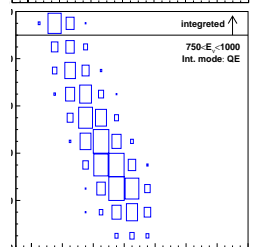
0
~0.5 GeV



0.5
~0.75 GeV

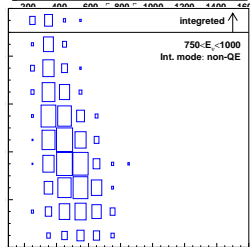
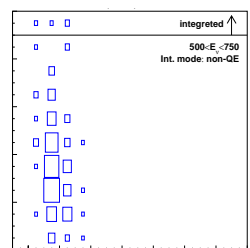


0.75
~1.0 GeV



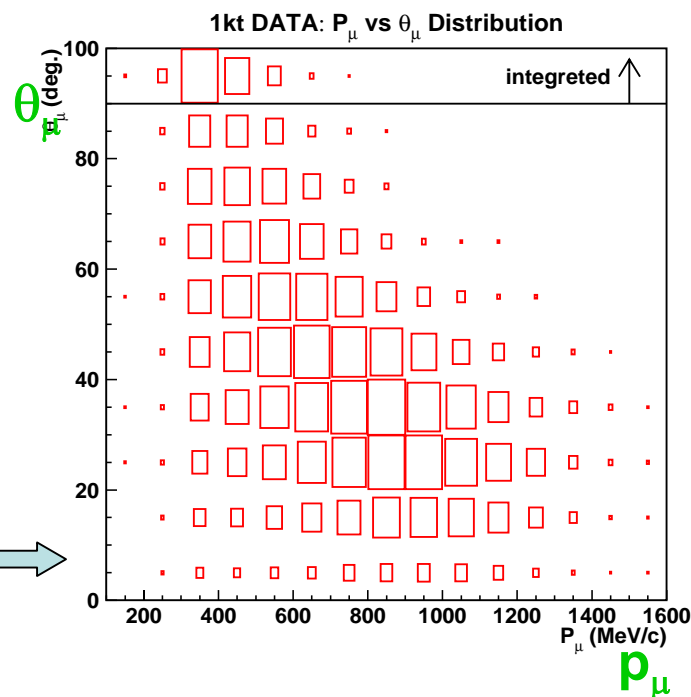
8bins

θ_μ
 p_μ



7bins

Measured (p_μ, θ_μ)



Fit the parameters.

$\phi(E_\nu)$, QE/non-QE ratio ...

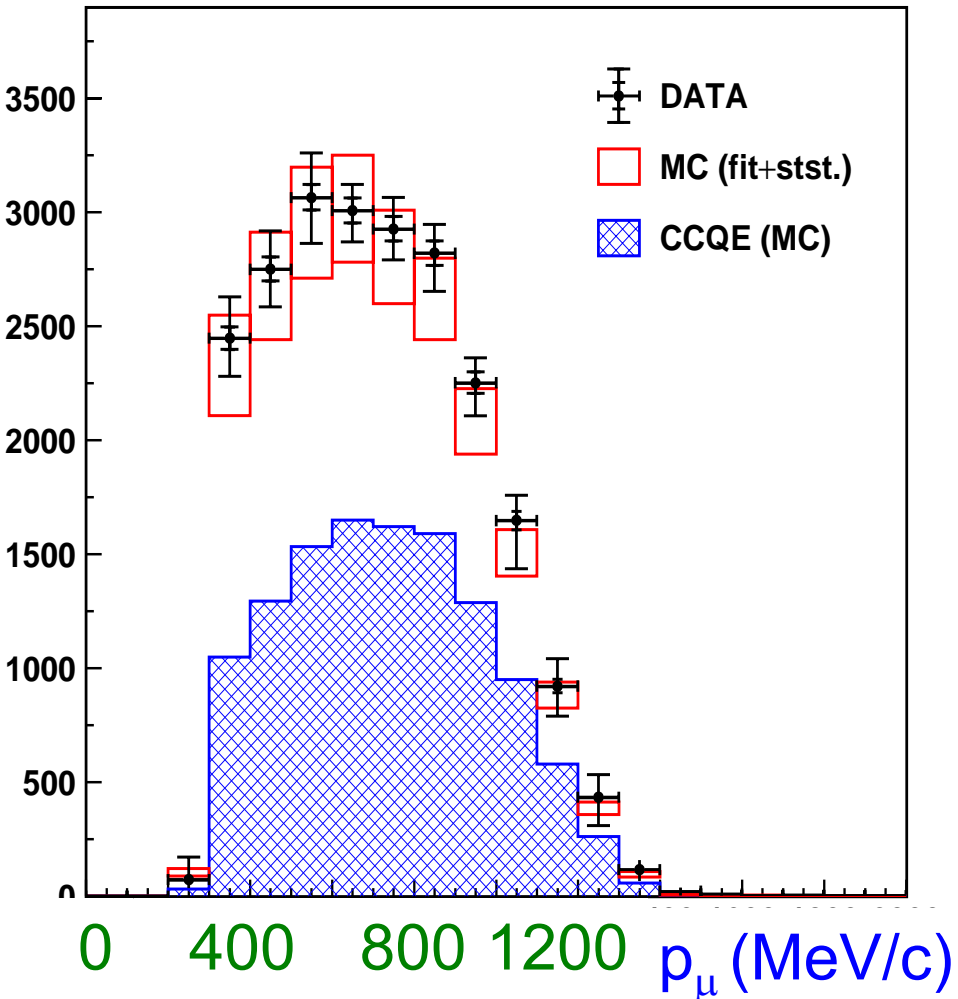
→ $\chi^2=227$ for 197 d.o.f.
(90 from 1kt, 137 from FGD)
for fitted (p_μ, θ_μ) dist. of 1KT
and FGD (124 data points)

Fitted results (1kt)

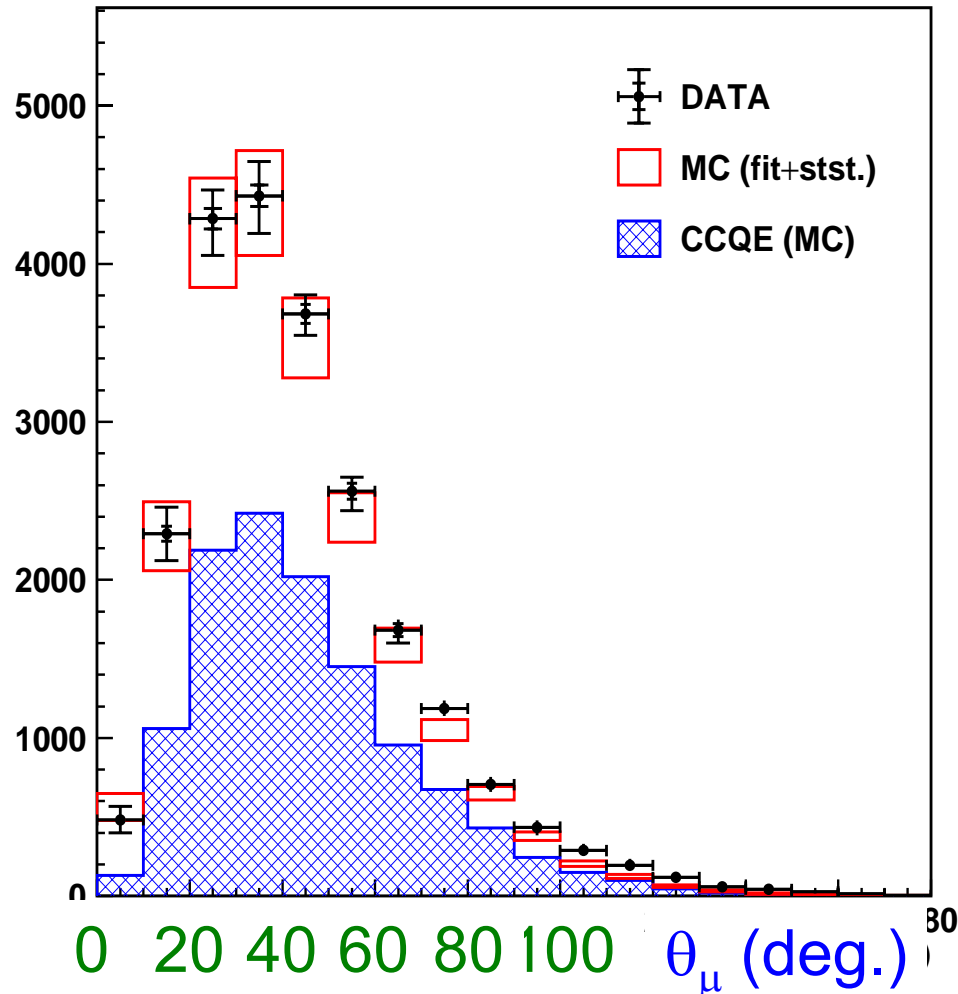
p_μ and θ_μ distributions of 1kt 1ring μ -like FCFV events

Both distributions **agree well** with the fitted MC.

1kt: μ -momentum Distribution (Fid.25t FC 1-Ring μ -like)

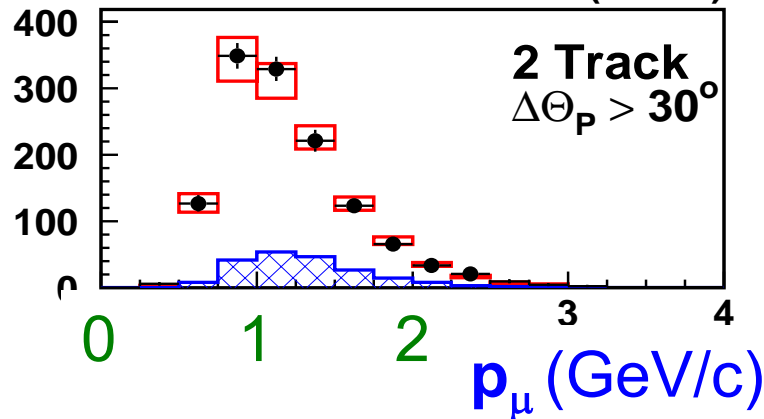
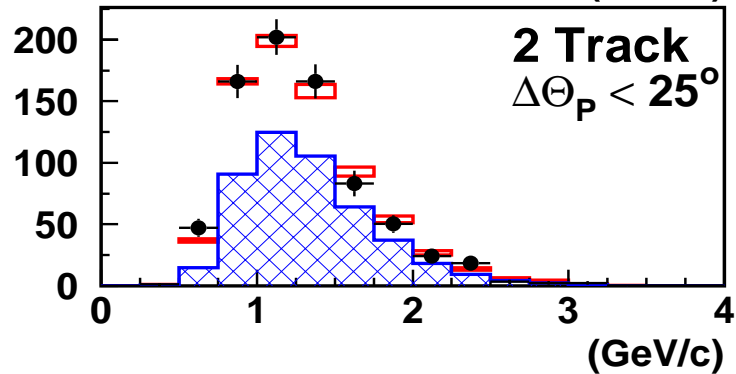
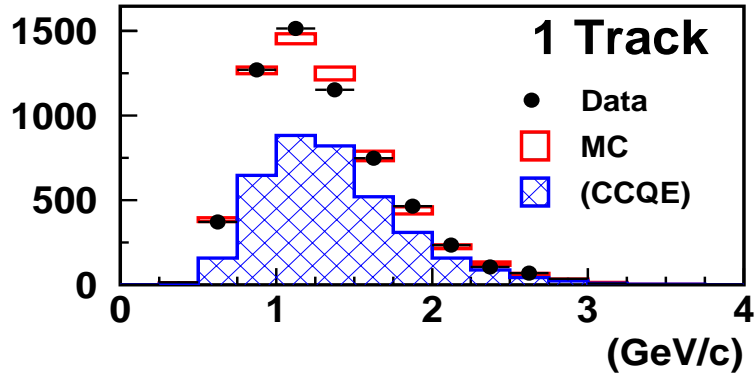


1kt: μ -angular Distribution (Fid.25t FC 1-Ring μ -like)

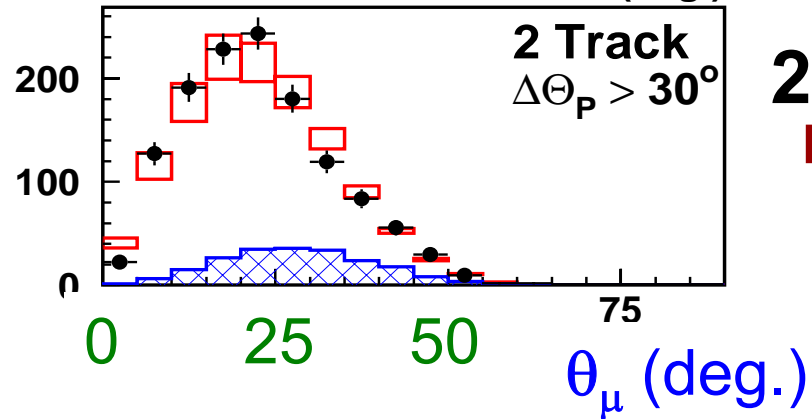
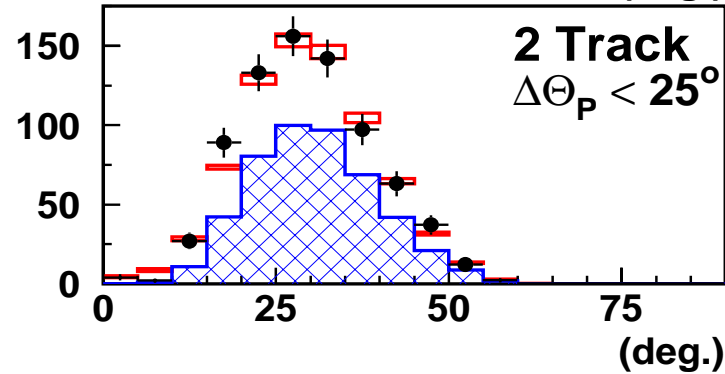
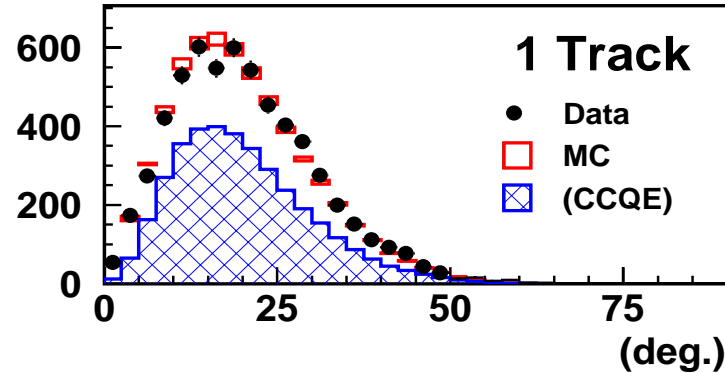


Fitted results (SciFi) Very good agreements

SciFi P_μ distributions



SciFi Θ_μ distributions



1track

2track
QE like

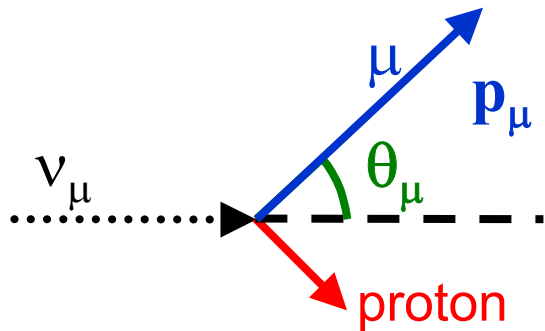
2track
non-QE
like

Reconstruction of neutrino energy at SK

Use single-ring μ -like FCFV ($1R_\mu$) events

Assuming QE interaction and reconstruct E_ν

($\sim 50\%$ of K2K $1R_\mu$ events are CCQE)



$$E_\nu^{\text{rec}} = \frac{m_n E_\mu - m_\mu^2 / 2}{m_n - E_\mu + P_\mu \cos \theta_\mu}$$

E_μ : muon energy

p_μ : muon momentum

θ_μ : muon angle

Oscillation analysis

1) Used data sets

1. Number of events

June '99 - July '01

FCFV events
(56 events)

2. Spectrum shape


Nov. '99 - July '01

1R μ events
(29 events)

2) Analysis methods

1. Maximum Likelihood method

$$L_{tot} = L_{norm}(f) L_{shape}(f) L_{syst}(f)$$


Normalization term ($N_{exp} = 80.1^{+6.2}_{-5.4}$)

Shape term (for 1R μ)

Constraint term for
systematic
parameters. (error
matrices)

2. different treatment of systematic term.

Generate many MC samples.

(changing systematic parameters within the error.)

$L_{tot} = \text{weighted mean of } L \text{ for the MC samples}$

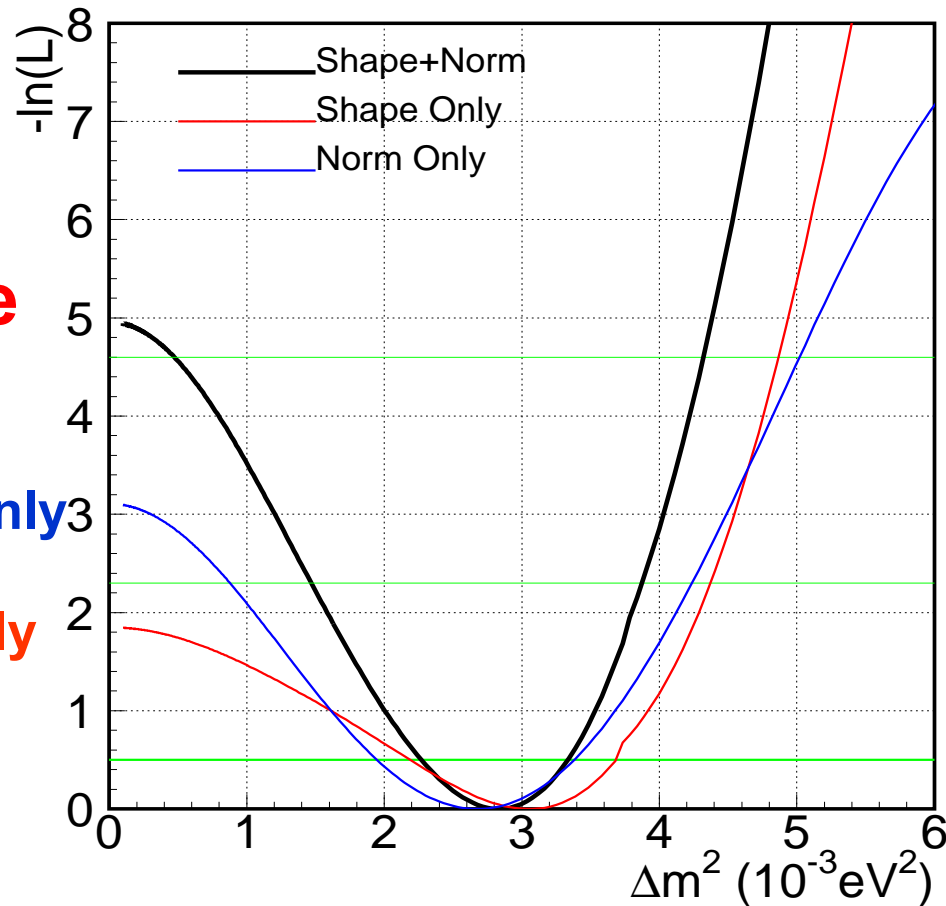
Allowed Δm^2

by N_{SK} and shape analysis

N_{SK} + Shape

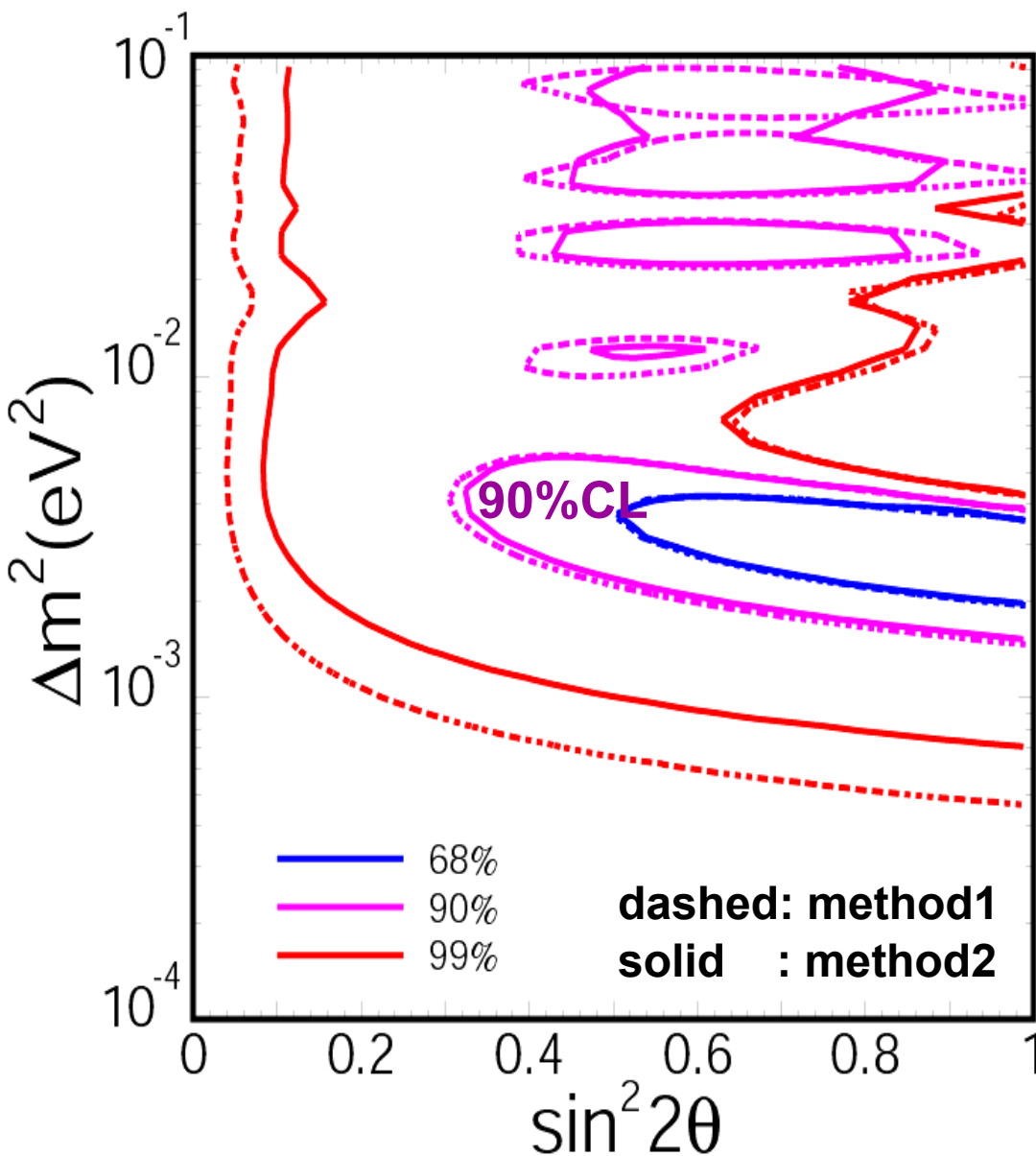
Number of Events only

Spectrum Shape only



**N_{SK} and shape analysis indicate
the same Δm^2 region for $\sin^2 2\theta = 1$**

Allowed regions



$$\Delta m^2 = 1.5 \sim 3.9 \times 10^{-3} \text{ eV}^2$$

$$@ \sin^2 2\theta = 1 \text{ (90\%CL)}$$

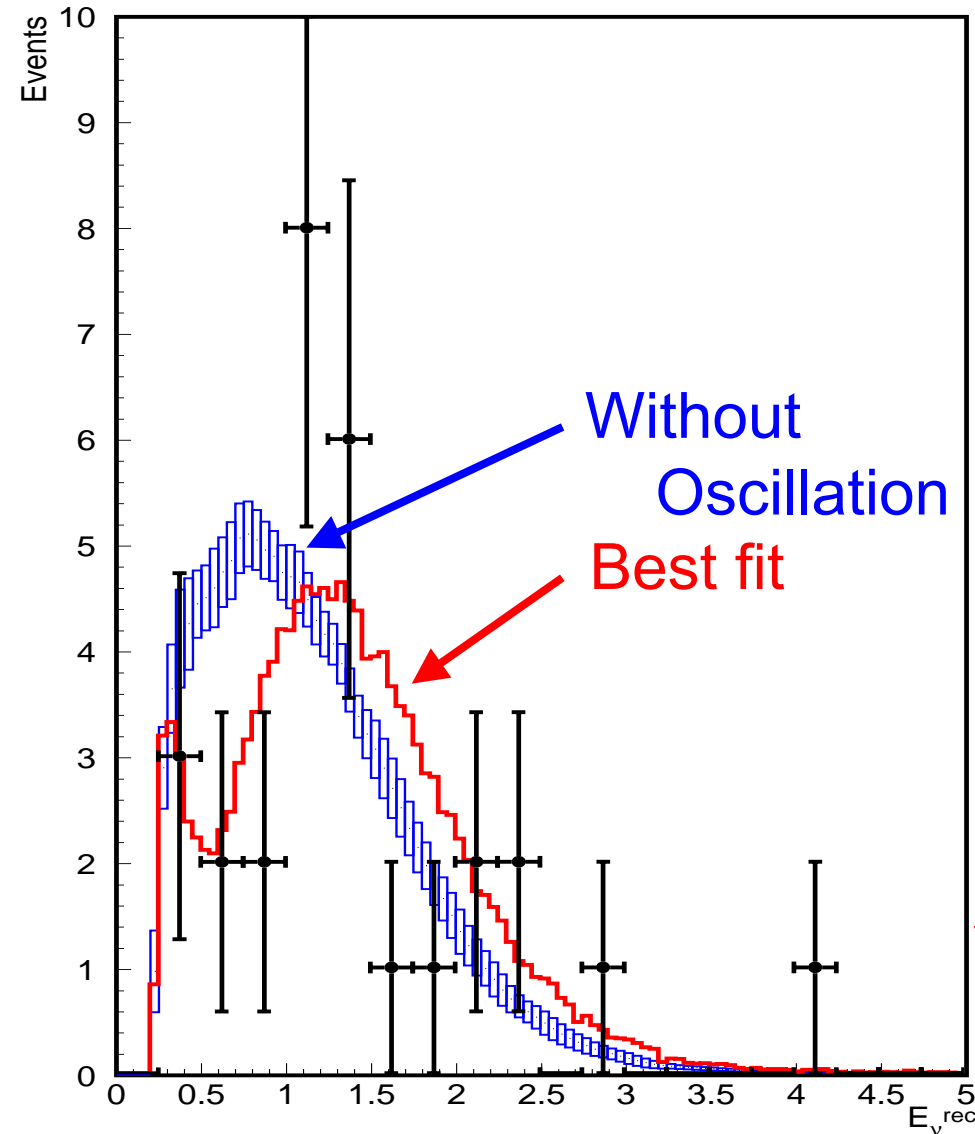
Best fit parameters

$$(\sin^2 2\theta, \Delta m^2)$$

$$\left\{ \begin{array}{l} = (1.0, 2.8 \times 10^{-3} \text{ eV}^2) \\ \quad \text{[Method-1]} \\ = (1.0, 2.7 \times 10^{-3} \text{ eV}^2) \\ \quad \text{[Method-2]} \end{array} \right.$$

Best fit results

Reconstructed energy of neutrino for $1R_\mu$ and N_{SK}



Best fit parameters
($\sin^2 2\theta, \Delta m^2$)

$$=(1.0, 2.8 \times 10^{-3} \text{ eV}^2)$$

- N_{SK}

Expected (W/osc.) = 54

Observed = 56

- Shape

KS test 79%

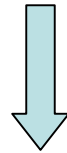
Very good agreements

Both shape and N_{SK}

Null oscillation probability

Use $\Delta\log(\text{likelihood})$ from best fit point
in the physical region

	method-1	method-2
N_{SK} only	1.3%	0.7%
Shape only	15.7%	14.3%
N_{SK} + Shape	0.7%	0.4%



Probability of null oscillation
is **less than 1%.**

Summary

K2K Oscillation analysis on June'99 ~ July '01 data

Full and Improved error estimations

and spectrum shape analysis

Use both Number of events + Spectrum shape
(June '99 – July '01) (Nov. '99 – July '01)

- Null oscillation probability is less than 1%.
- Spectrum shape distortion
and observed # of events @SK
indicates consistent oscillation parameter regions.
- Oscillation parameters ($\sin^2 2\theta$ and Δm^2) are
consistent with the atmospheric ν results.
 $\Delta m^2 = 1.5 \sim 3.9 \times 10^{-3} \text{ eV}^2$ @ $\sin^2 2\theta = 1$ (90%CL)
(c.f. ATM ν : $\Delta m^2 = 1.6 \sim 3.9 \times 10^{-3} \text{ eV}^2$ @ $\sin^2 2\theta = 1$ (90%CL))