

JHF-Kamioka Neutrino Oscillation Experiment using JHF 50 GeV PS

- Introduction
- Facility
- Physics goal
- Summary

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Where we go ?

Cosmic ray
Discovery Phase


Atmospheric ν

$$\Delta m_{23}^2 \sim 1.6 - 3.9 \times 10^{-3} \text{ eV}^2$$

$$\sin^2 2\theta_{23} > 0.92$$

Discovery of ν_μ oscillations

Solar ν

$$\Delta m_{12}^2 \sim 3 - 20 \times 10^{-5} \text{ eV}^2$$

$$\sin^2 2\theta_{12} \sim 0.55 - 0.95$$

Large mixing in ν_e oscillations

K2K

The 1st long baseline ν experiment



Precision study of 3-flavor ν mixing matrix

3 Flavor Mixing

- If neutrinos are massive particles, then it is possible that the **mass eigenstates** and the **weak eigenstates** are not the same:

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

↑ *3 independent parameters + 1 complex phase*

Weak eigenstates
 „flavor eigenstates“

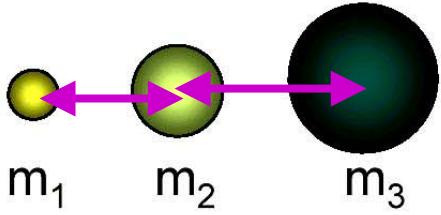
$\Delta m^2_{12}, \Delta m^2_{23}$

Mass eigenstates



$\nu_e \quad \nu_\mu \quad \nu_\tau$

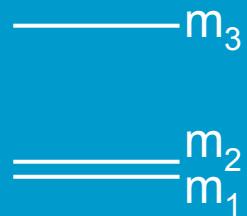
$\theta_{12}, \theta_{23}, \theta_{13}$
+ δ



$m_1 \quad m_2 \quad m_3$

MNS (Maki-Nakagawa-Sakata) matrix

3-flavor Oscillation



Oscillation Probabilities when $\Delta m_{12}^2 \ll \Delta m_{23}^2 \approx \Delta m_{13}^2$

➤ θ_{23} : ν_μ disappearance

$$P_{\mu \rightarrow x} \approx 1 - \cos^4 \theta_{13} \cdot \sin^2 2\theta_{23} \cdot \sin^2 \left(1.27 \frac{\Delta m_{23}^2 L}{E_\nu} \right)$$

➤ θ_{13} : ν_e appearance

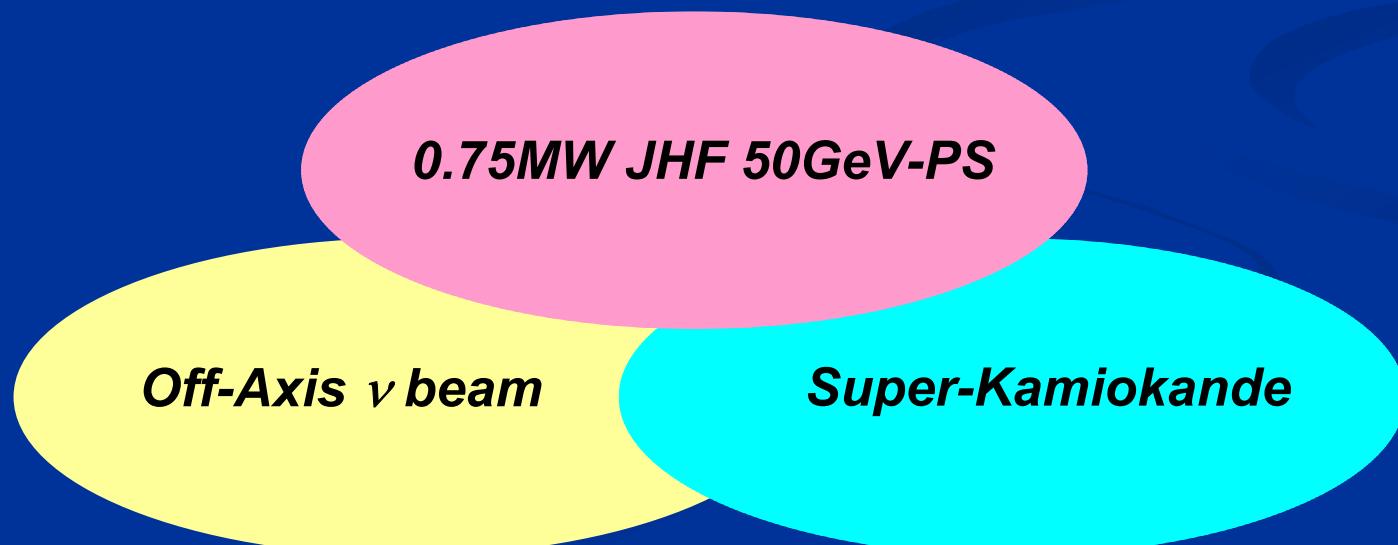
$$P_{\mu \rightarrow e} \approx \sin^2 \theta_{23} \cdot \sin^2 2\theta_{13} \cdot \sin^2 \left(1.27 \frac{\Delta m_{13}^2 L}{E_\nu} \right)$$

➤ δ : ~~CP~~ in ν_e appearance

$$A_{CP} = \frac{P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}{P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)} \approx \frac{\Delta m_{12}^2 L}{4E_\nu} \bullet \frac{\sin 2\theta_{12}}{\sin \theta_{13}} \bullet \sin \delta$$

Strategy

- High statistics by high intensity ν beam
- Tune $E\nu$ at oscillation maximum
- Narrow band beam to reduce BG
- Sub-GeV ν beam for Water Cherenkov



JHF-Kamioka Neutrino Experiment

(hep-ex/0106019)



JHF 0.75MW + Super-Kamiokande

Future Super-JHF 4MW + Hyper-K ~ JHF+SK × 200

Organization

Experiment working group

- Dec.1999 : formed (ICRR/KEK/Kyoto/Kobe/Tohoku/TRIUMF)
- Jun.2001 : Letter of Intent (hep-ex/0106019)
- Mar.2002 : Meeting to discuss possibility to organize international collaboration

Facility Construction Group

- Officially formed in KEK on **April, 2001**
- The 3rd physics division, IPNS, Cryogenic facility group, Cryogenic Science Center, w/ strong support from KEK-PS beam channel group

JHF-SK Neutrino Working Group

ICRR/Tokyo-KEK-Kobe-Kyoto-Tohoku-TRIUMF

**Y. Itow, T. Kajita, K. Kaneyuki, Y. Obayashi,
C. Saji, M. Shiozawa, Y. Totsuka (ICRR/Tokyo)**

**Y. Hayato, A. Ichikawa, T. Ishida, T. Ishii, T. Kobayashi,
T. Maruyama, K. Nakamura, Y. Oyama, M. Sakuda (KEK)**

S. Aoki, T. Hara, A. Suzuki (Kobe)

T. Nakaya, K. Nishikawa (Kyoto)

T. Hasegawa, K. Ishihara (Tohoku)

A. Konaka (TRIUMF, CANADA)

In addition to the above user group, the neutrino facility construction group is OFFICIALLY formed at KEK.



JHF Facility

Pacific Ocean

Construction
2001~2006 (approved)

✓ beam-line
budget request submitted

3GeV PS

400MeV LINAC

- 8 bunches/ $\sim 5\mu s$
- 3.3×10^{14} proton/pulse
- 3.94 (3.64) sec cycle
- $1 \text{ year} \equiv 10^{21} \text{ POT}$ (130 days)

Transport line
(Super-cond. Mag.)

50GeV PS

Neutrino facility

Target station

Decay volume

Near detectors (280m, 2km)

Off Axis Beam

(ref.: BNL-E889 Proposal)

- ◆ Quasi Monochromatic Beam
- ◆ $\times 2 \sim 3$ intense than NBB

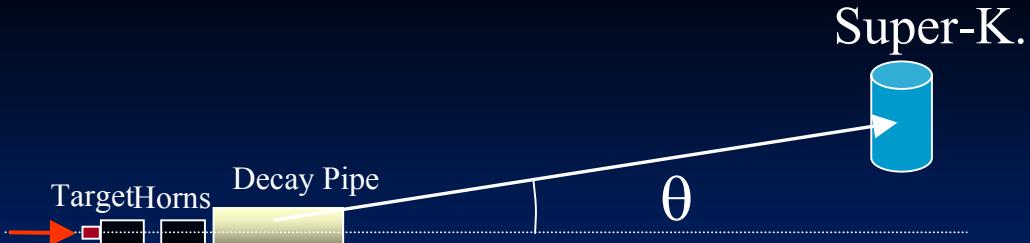
Statistics at SK

(OAB2deg, 1yr, 22.5kt)

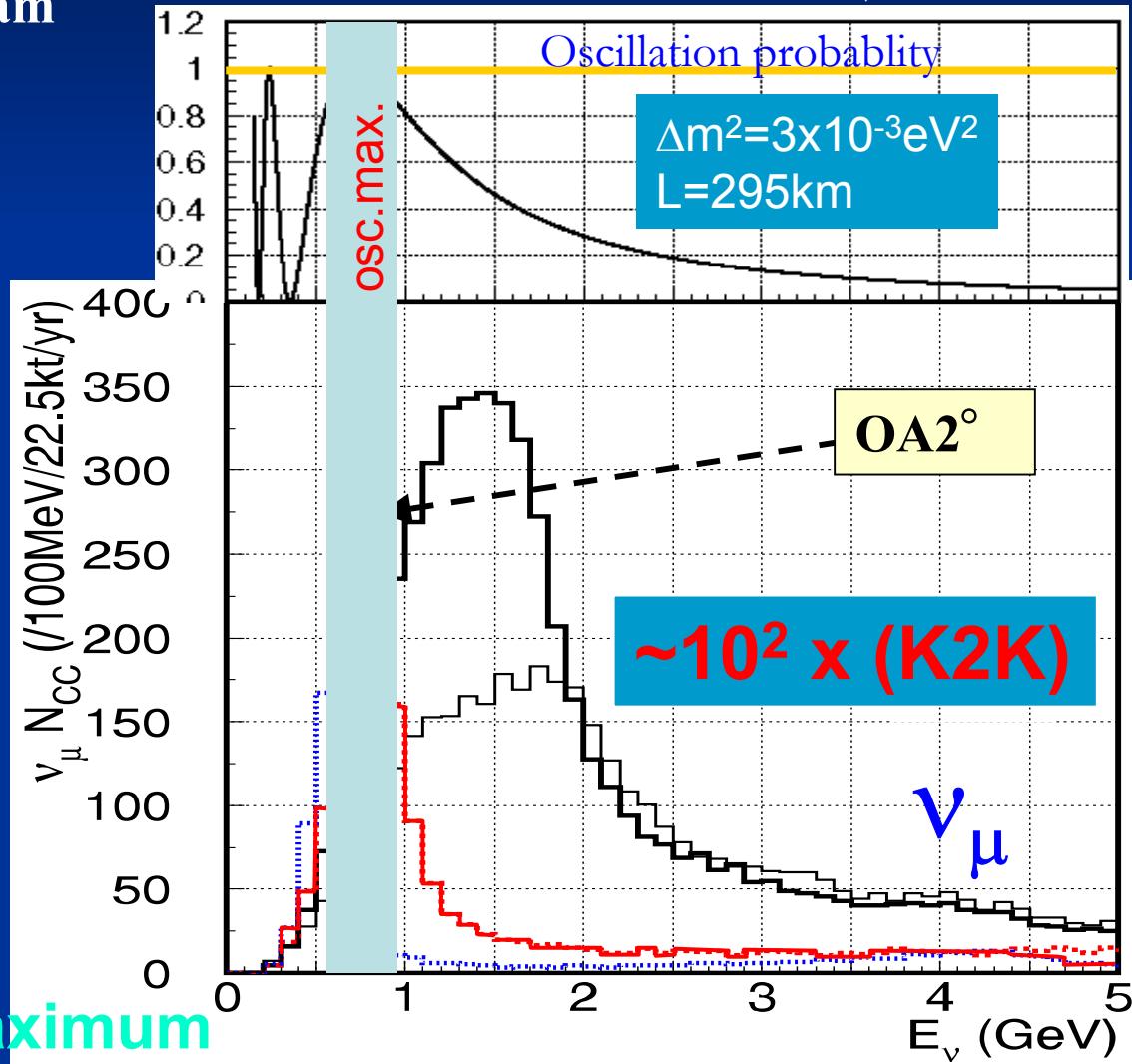
~4500 ν_μ tot

~3000 ν_μ CC

ν_e ~0.2% at ν_μ peak

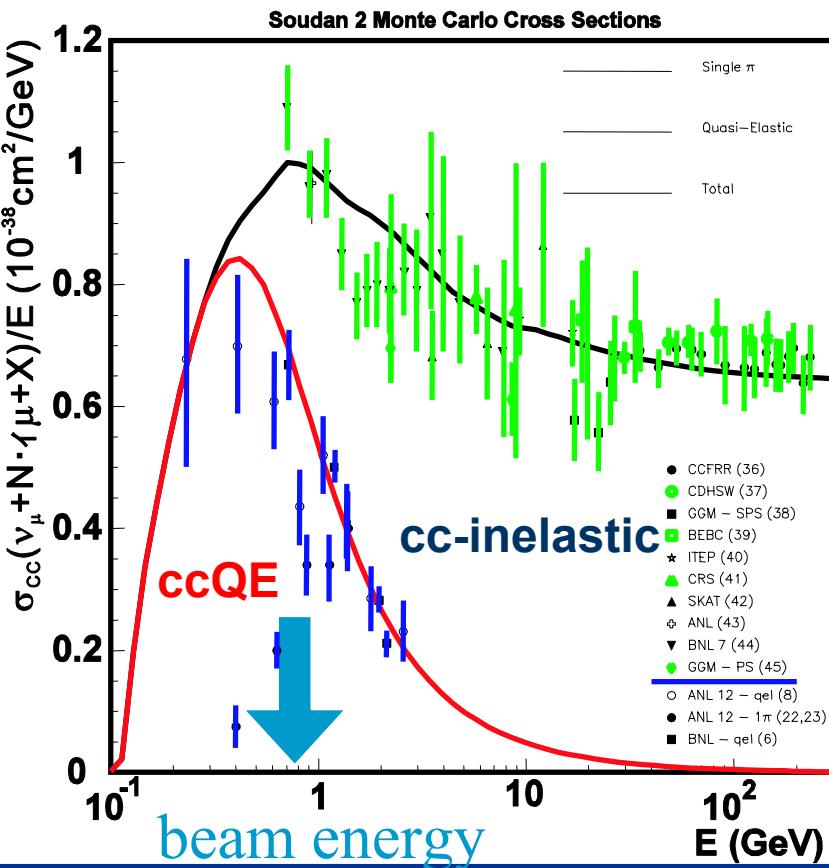
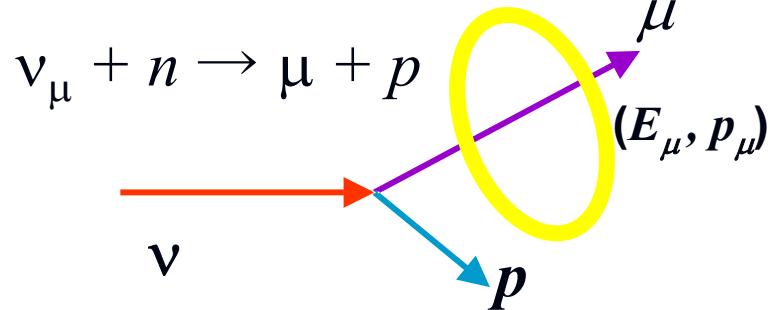


$$\text{Osc. Prob.} = \sin^2(1.27\Delta m^2 L/E_\nu)$$

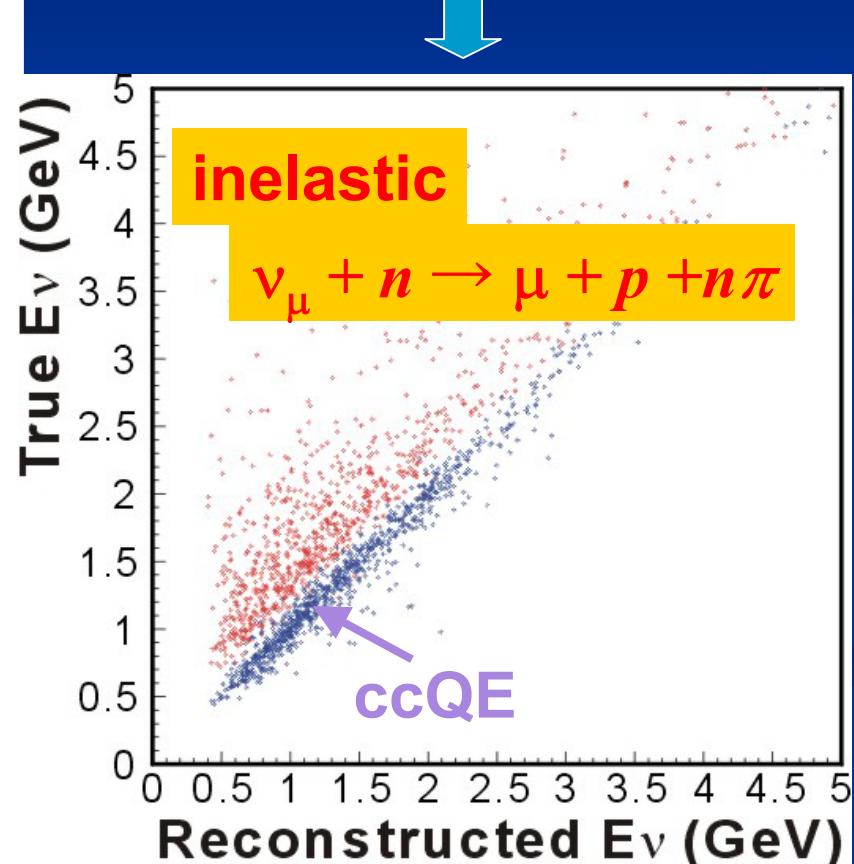


Ev reconstruction in water Cherenkov

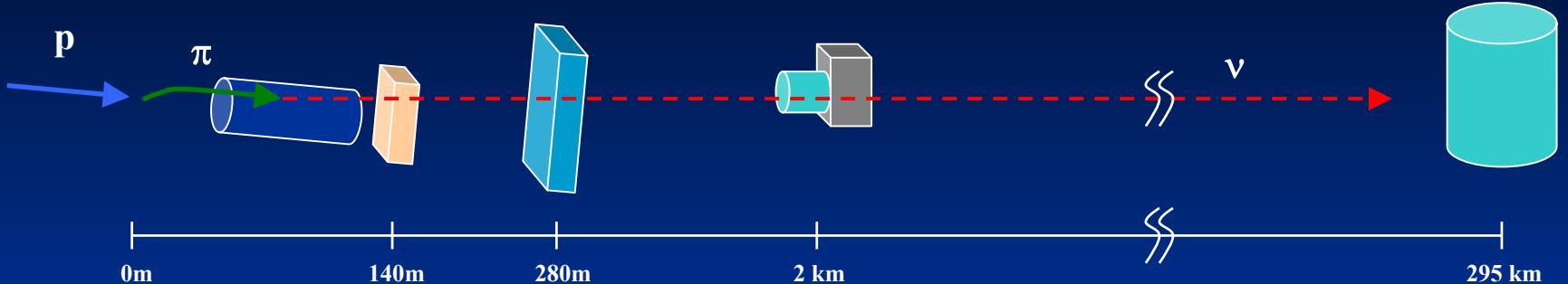
Assume CC Quasi Elastic (QE) reaction



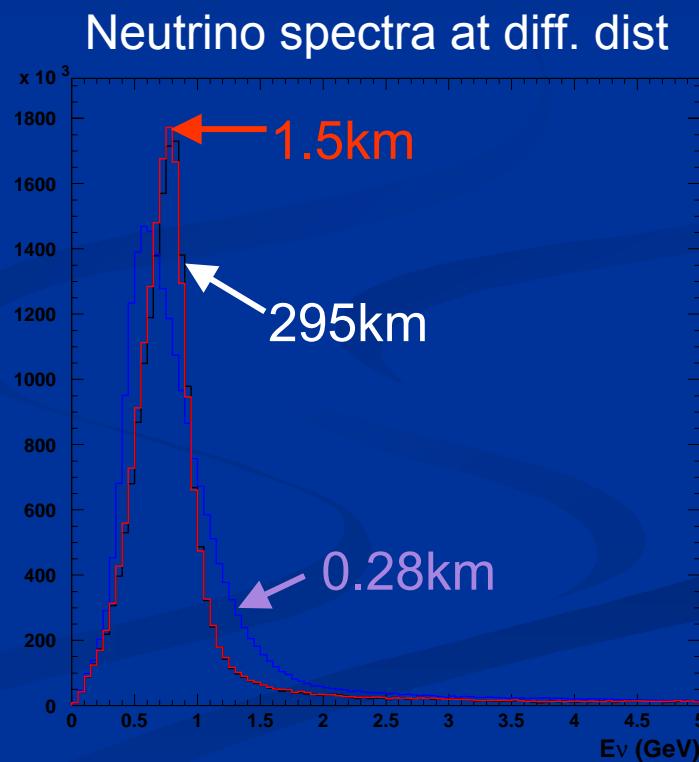
$$E_\nu = \frac{m_N E_\mu - m_\mu^2/2}{m_N - E_\mu + p_\mu \cos\theta_\mu}$$



Detectors



- Muon monitors @ $\sim 140\text{m}$
 - Spill-by-spill monitoring of beam direction
- First Front detector @ 280m
 - Neutrino intensity/direction
- Second Front Detector @ $\sim 2\text{km}$
 - Almost same E_ν spectrum as for SK
 - Water Cherenkov can work
- Far detector @ 295km
 - Super-Kamiokande (50kt)

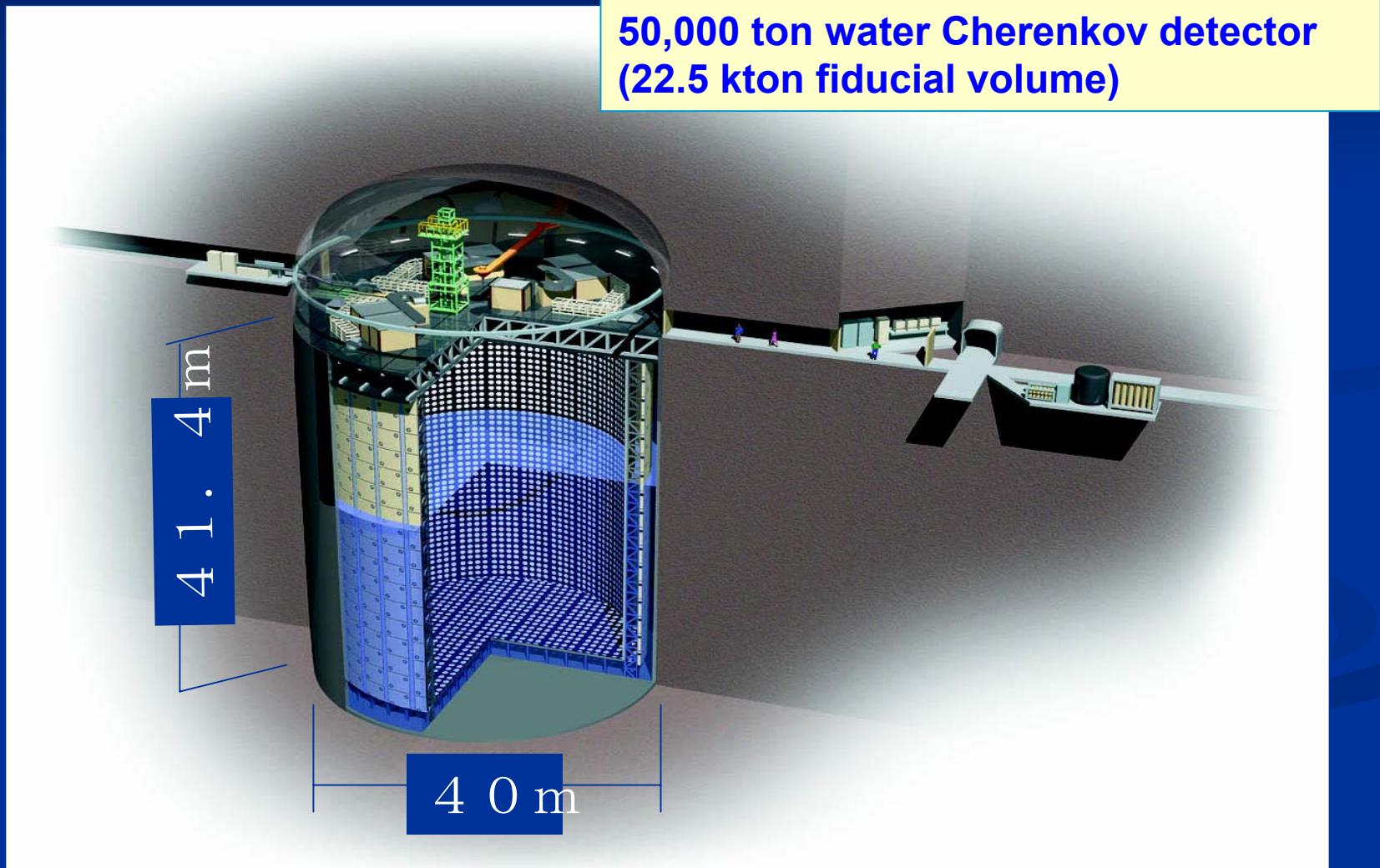


dominant syst. in K2K

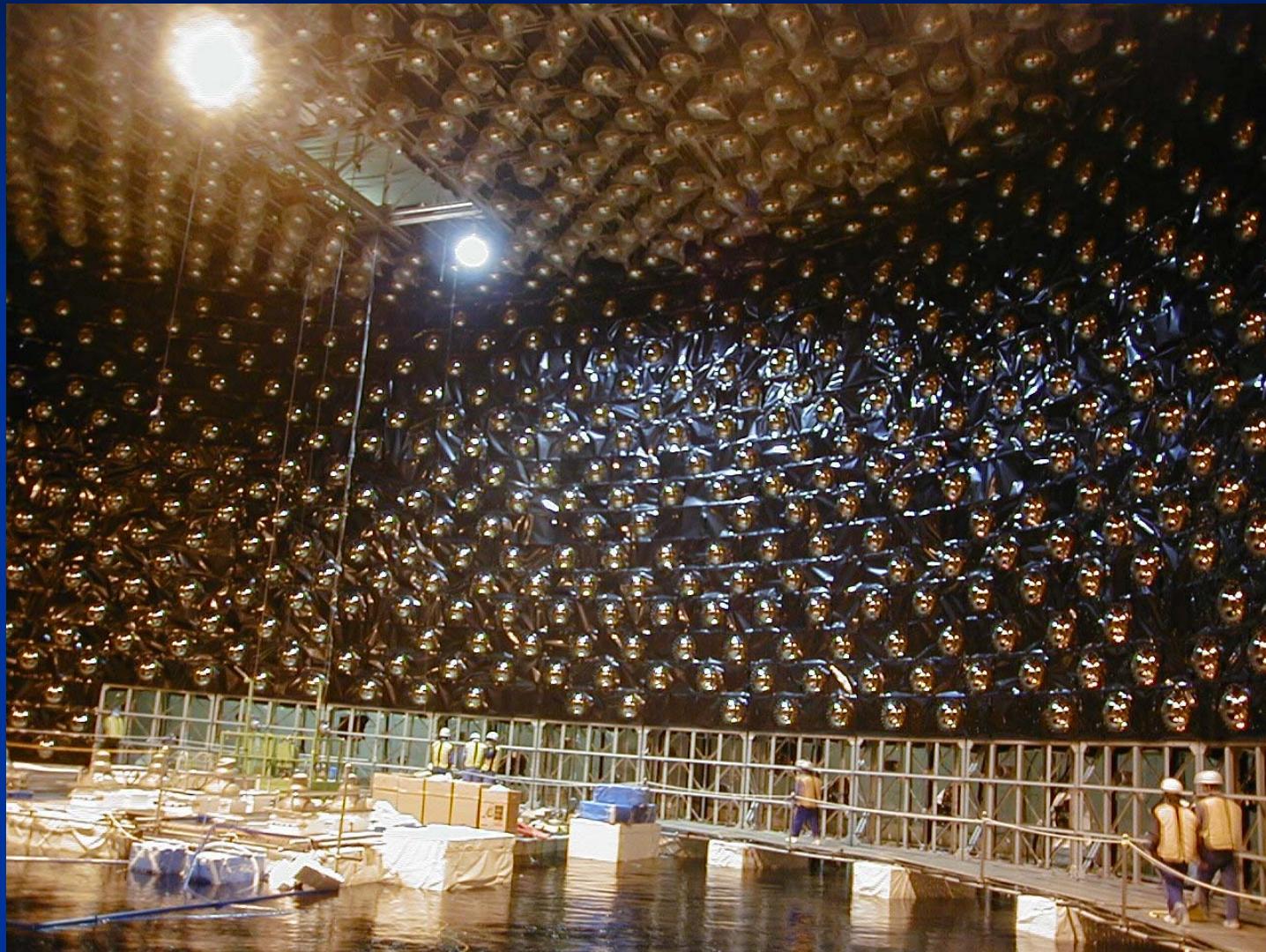
Far detector Super-Kamiokande

(since Apr 1996)

**50,000 ton water Cherenkov detector
(22.5 kton fiducial volume)**



Re-Construction of Super-K (Jul 2002)



Physics Goal

★ Precise measurement of neutrino mixing matrix

Accuracy: $\sin^2 2 \theta_{23} \cdots \cdots \cdots \cdots \cdots 1\%$

$\Delta m^2_{23} \cdots \cdots \cdots \cdots \cdots \cdots \cdots \cdots \cdots$ a few %

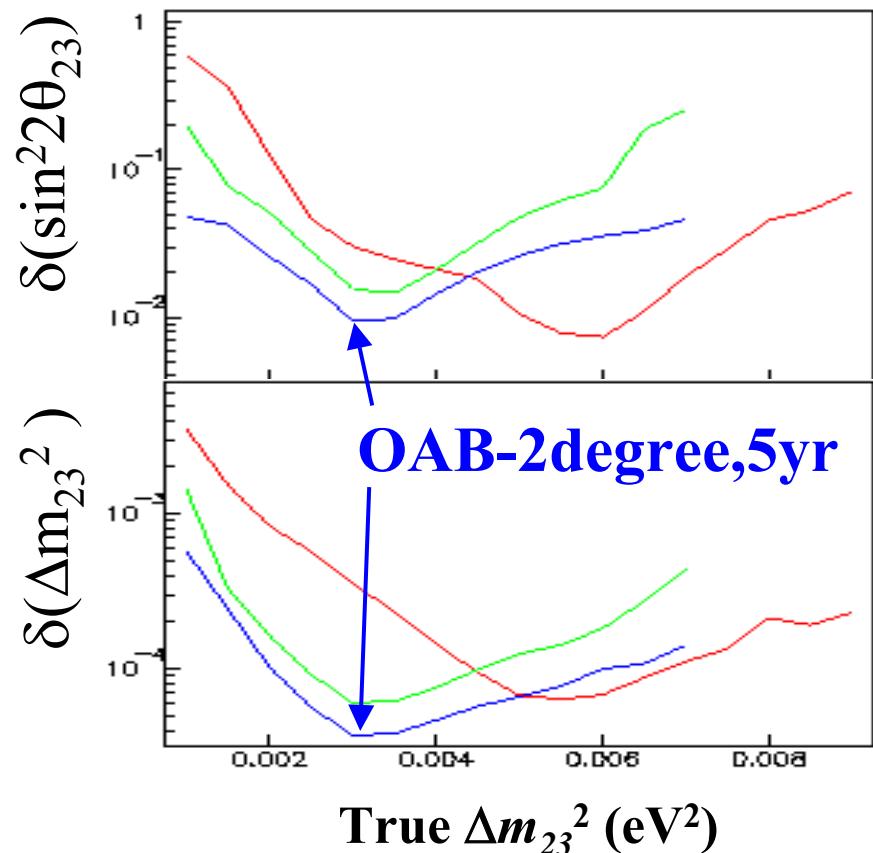
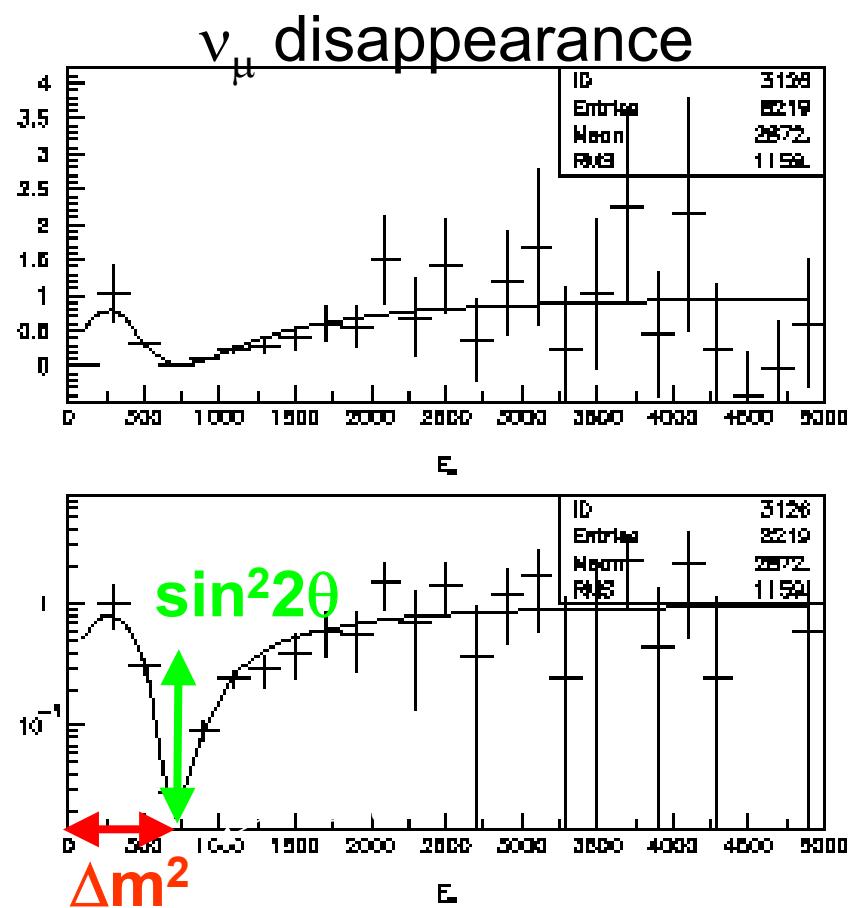
★ Discovery and measurement of non-zero θ_{13}

$\sin^2 2 \theta_{13} \cdots \cdots \cdots \cdots > 0.006$

1st Evidence of 3-flavor mixing !

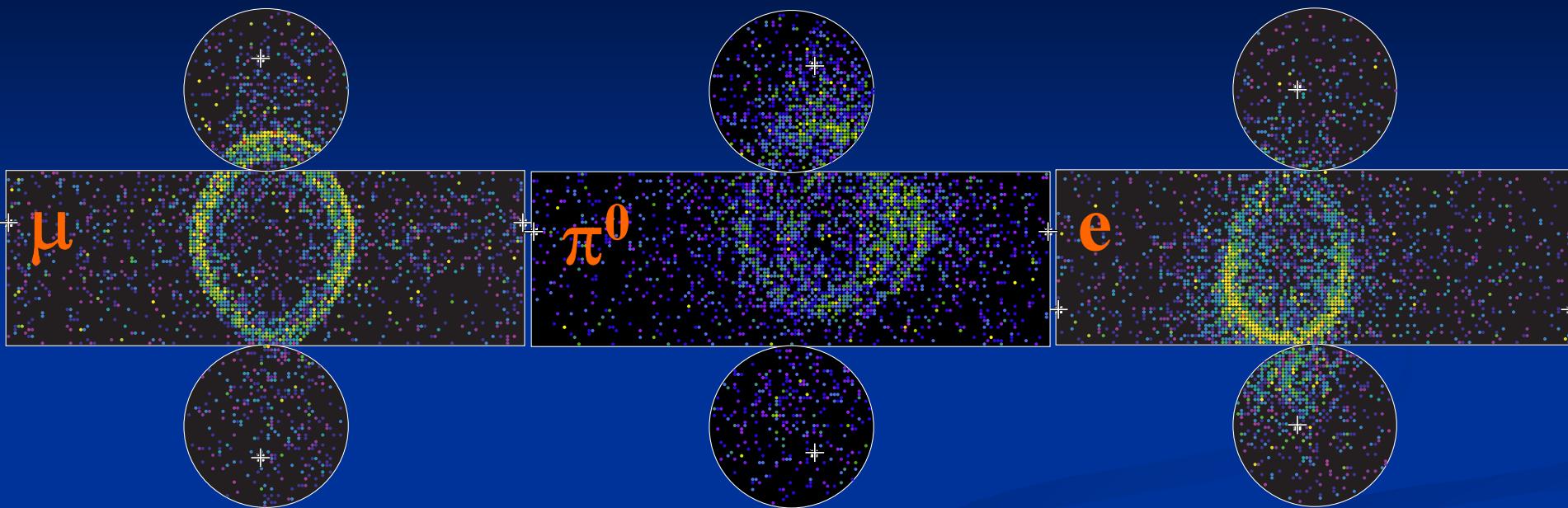
1st step to CP measurement

Measurement of $\sin^2 2\theta_{23}$, Δm_{23}^2



$\delta(\sin^2 2\theta) \sim 0.01$
 $\delta(\Delta m^2) \sim < 1 \times 10^{-4}$

ν_e appearance in JHF-Kamioka



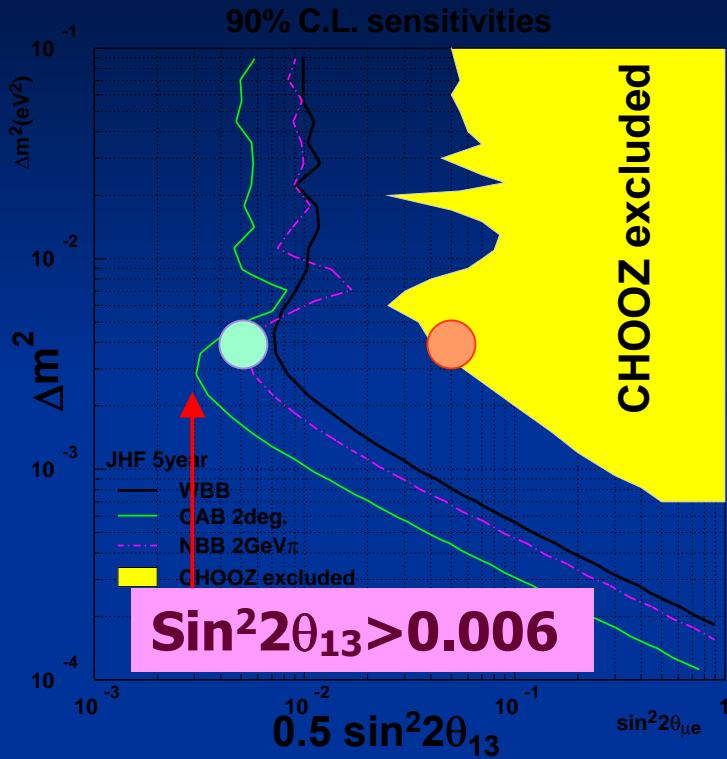
Back ground for ν_e appearance search

- Intrinsic ν_e component in initial beam
- Merged π^0 ring from ν_μ interactions

10% uncertainty for BG estimation

The 1kt π^0 data will be studied for exercise

$\sin^2 2\theta_{13}$ from ν_e appearance



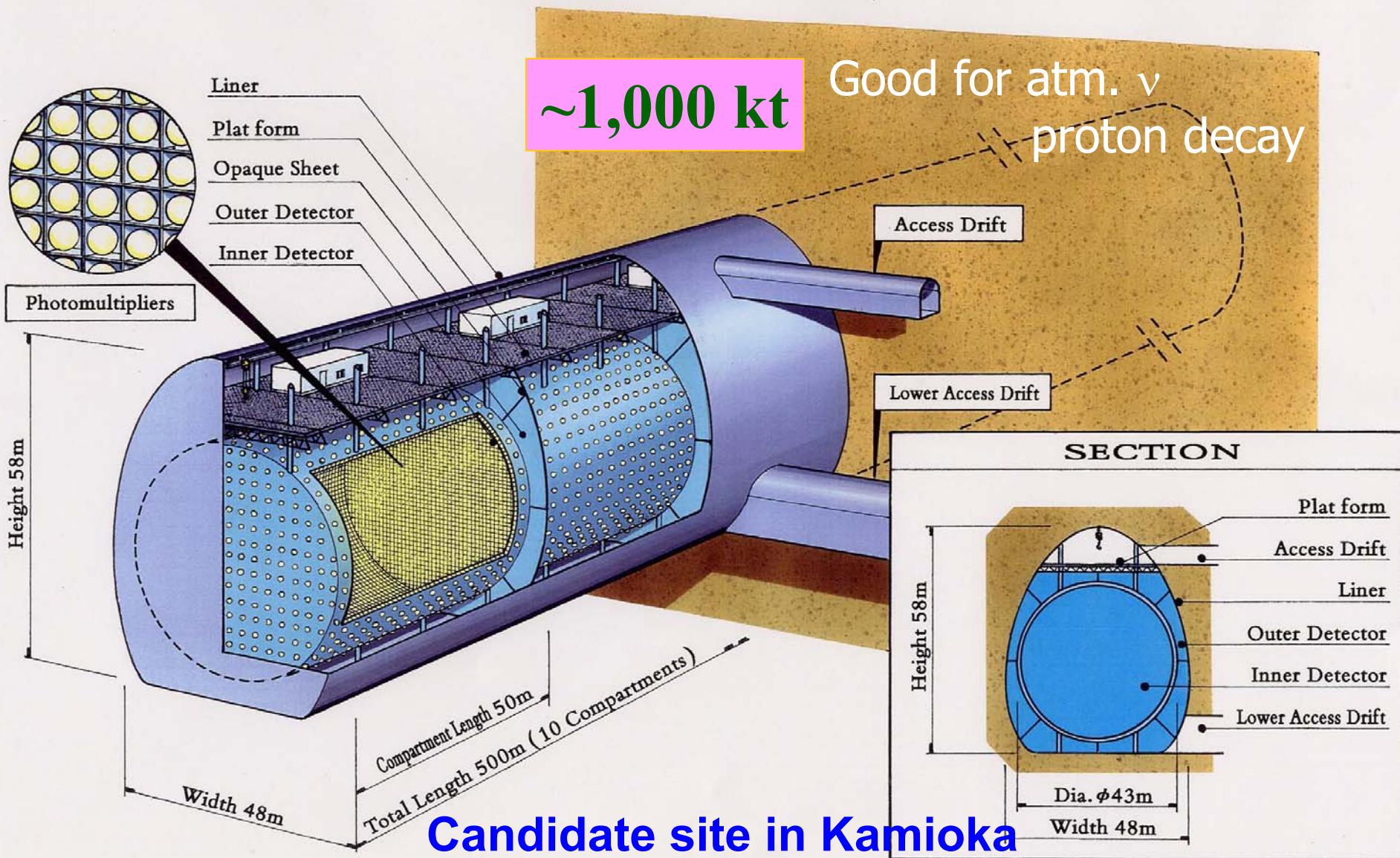
at Off axis 2 deg, 5 years

Off axis 2 deg, 5 years

$\sin^2 2\theta_{13}$	Background in Super-K (as of Oct 25, 2001)					Signal	Signal + BG
	ν_μ	ν_e	$\bar{\nu}_\mu$	$\bar{\nu}_e$	total		
0.1	12.0	10.7	1.7	0.5	24.9	114.6	139.5
0.01	12.0	10.7	1.7	0.5	24.9	11.5	36.4

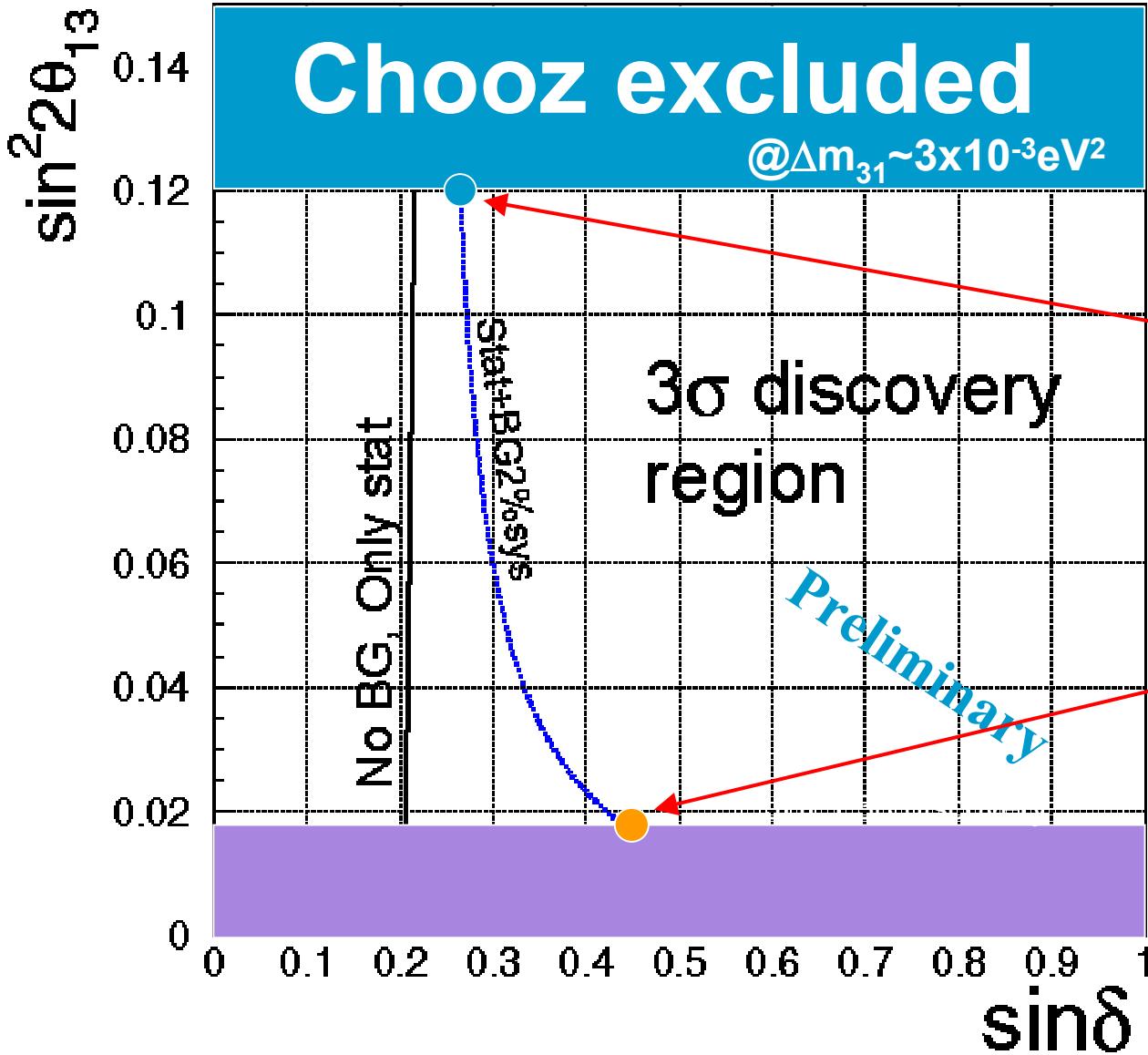
Future upgrade

Super-JHF(4MW)+Hyper-K(1Mt)



Sensitivity(3σ) to CP violation (SuperJHF+Hyper-K)

JHF-HK CPV Sensitivity



4MW, 1Mt
2yr for ν_μ
6.8yr for $\bar{\nu}_\mu$

$\delta > \sim 14 \text{ deg}$

$\delta > \sim 27 \text{ deg}$

Global Schedule

2002

2004

2006

2008

2010



JHF-v construction

physics run (OAB)

SK
rebuild

SK-half

SK-full



MINOS 2yr



OPERA 5yr



Summary

- Precision study of neutrino mixing matrix
 - Next step after the discovery
 - We may find a hint for next break-through
- JHF-Kamioka neutrino experiment (2007~)
 - JHF 0.75MW 50GeV-PS+Off Axis beam+Super-K
 - Narrow band beam at oscillation maximum ($\sim 1\text{GeV}$)
 - ν_e appearance, discovery of θ_{13} ($\sin \theta_{13} > 0.006, 90\%\text{CL}$)
 - Budget request submitted, R&D started
- Possible upgrade in future
 - 4MW Super-JHF + Hyper-K (1Mt water Cherenkov)
 - CP violation in lepton sector
 - Bonus ! (Proton Decay !)