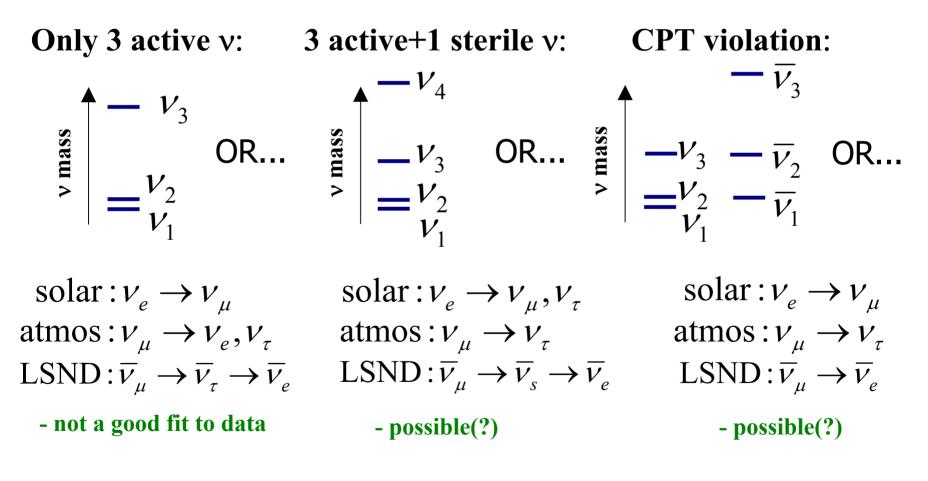


 $sin^2(2\theta)$ 

# v Oscillation Scenarios:

With current results from solar, atmospheric, and LSND  $\nu$ -oscillation searches (3  $\Delta m^2 s$ ), we have an interesting situation:



Need to definitively check the LSND result.

# Enter: MiniBooNE

### High statistics

×10 more events than LSND (~2 calendar years)

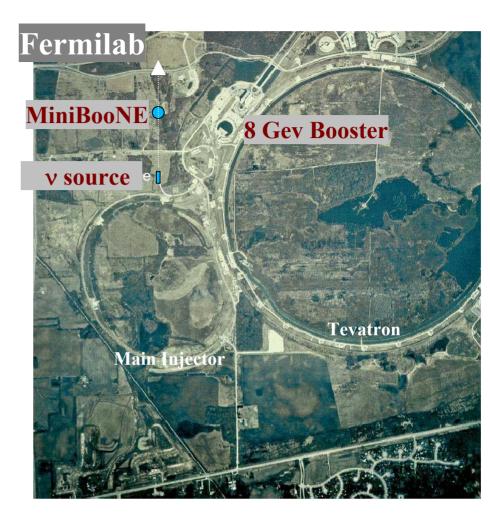
### Different systematics

× 10 higher beam energy different event signatures and backgrounds

### High significance

 $5\sigma$  over entire LSND region as a "counting experiment" (more significant when energy dependence is included)

Start to run in summer 2002



#### **BooNE** Collaboration

Y. Liu, I. Stancu University of Alabama, Tuscaloosa, AL 35487

S. Koutsoliotas Bucknell University, Lewisburg, PA 17837

E. Church, C. Green, G.J. VanDalen University of California, Riverside, CA 92521

E. Hawker, R.A. Johnson, J.L. Raaf University of Cincinnati, Cincinnati, OH 45221

T. Hart, E.D. Zimmerman University of Colorado, Boulder, CO 80309

J.M. Conrad, J. Link, J. Monroe, M.H. Shaevitz, M. Sorel, G.P. Zeller Columbia University, Nevis Labs, Irvington, NY 10533

D. Smith Embry Riddle Aeronautical Univ., Prescott, AZ 86301 C. Bhat, S J. Brice, B.C. Brown, L. Bugel,
B.T. Fleming, R. Ford, F.G. Garcia, P. Kasper,
T. Kobilarcik, I. Kourbanis, A. Malensek,
W. Marsh, P. Martin, F. Mills, C. Moore,
P. J. Nienaber, E. Prebys, A. Russell,
P. Spentzouris, R. Stefanski, T. Williams

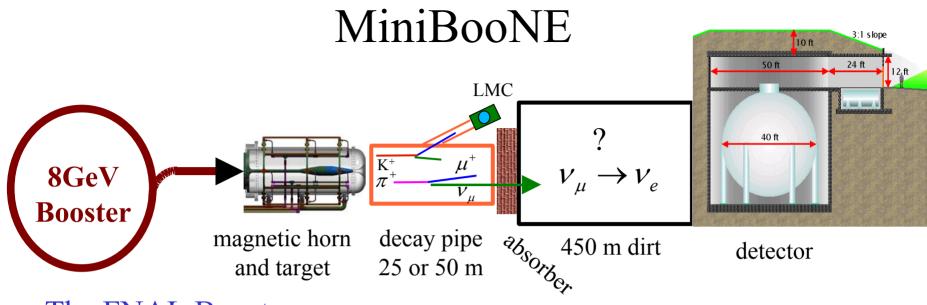
D. C. Cox, A Green, H. -O. Meyer, R. Tayloe Indiana University, Bloomington, IN 47405

G.T. Garvey, W.C. Louis, G.B. Mills, V. Sandberg,B. Sapp, R. Schirato, R. Van de Water, D. H. White *Los Alamos National Lab, Los Alamos, NM 87545* 

R. Imlay, W. Metcalf, M. Sung, M.O. Wascko Louisiana State University, Baton Rouge, LA 70803

J. Cao, Y. Liu, B.P. Roe University of Michigan, Ann Arbor, MI 48109

A.O. Bazarko, P.D. Meyers, R.B. Patterson, F.C. Shoemaker *Princeton University, Princeton, NJ 08544* 



The FNAL Booster

injects beam to the Be target

resulting mesons decay

neutrinos traverse 450 m of dirt to the oil-based Cherenkov detector

source:

 $\pi^+ \rightarrow \mu^+ v_{\mu}$  search for  $v_{\mu} \rightarrow v_e$ , which leads to  $e^-$  in detector

backgrounds are:  $V_{\mu}$  interactions with mis-id  $\mu^{-}$  or  $\pi^{0}$  as  $e^{-}$ intrinsic  $V_{e}$  in the beam from  $\mu^{+}$  or K-meson decay

#### The Booster

8 GeV proton accelerator built to supply beam to the Main Ring, it now supplies the Main Injector

Booster must now run at record intensity



MiniBooNE will run simultaneously with the other programs:
e.g. Run II + MiniBooNE
5 x 10<sup>12</sup> protons per pulse, machine running at a rate of 7 Hz (5 Hz for MiniBooNE)

MiniBooNE:  $5 \times 10^{20}$  protons on target in one year

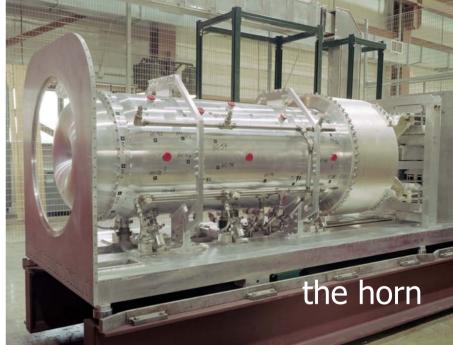
Due to radiation issues it will be a challenge to reach these goals.

#### A magnetic horn focuses the charged particles to the detector.

Initially positive particles will be focused (neutrinos)  $\pi^+ \rightarrow \mu^+ v_{\mu}$ 

then the horn current can be reversed (antineutrinos)  $\pi^- \rightarrow \mu^- \overline{\nu}_{\mu}$ 





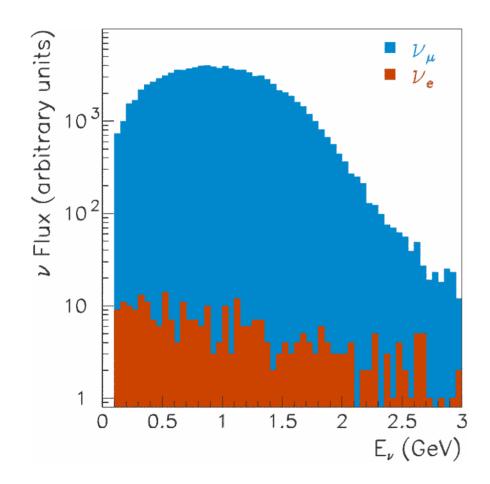
170 kA in 140 µsec pulses @ 5 Hz

Tested to 10 million pulses behaves as expected (vibration, temperature, etc.)



### Neutrino Flux at the Detector

The L/E  $\sim$ 1 m/MeV is similar to that at LSND.



-8 GeV protons on Be:  $\mathbf{p} + \mathbf{Be} \rightarrow \pi^+, \mathbf{K}^+, \mathbf{K}^0_L$ -yield a high flux of  $v_\mu$ :  $\pi^+ \rightarrow \mu^+ v_\mu$   $\mathbf{K}^+ \rightarrow \mu^+ v_\mu$ ,  $\mathbf{K}^0_L \rightarrow \pi^- \mu^+ v_\mu$ -with a low background of  $v_e$ :  $\mu^+ \rightarrow e^+ v_e \overline{v}_\mu$  $\mathbf{K}^+ \rightarrow \pi^0 e^+ v_e$ ,  $\mathbf{K}^0_L \rightarrow \pi^- e^+ v_e$ 

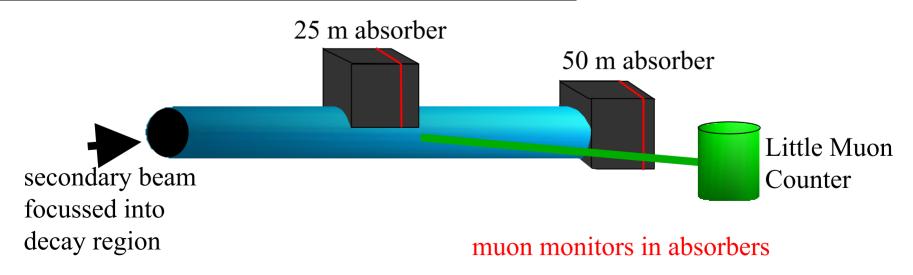
#### Flux estimate is important!

Secondary particle production

Meson production by 8 GeV protons on MiniBooNE target slug will be measured by the HARP experiment at CERN in August.







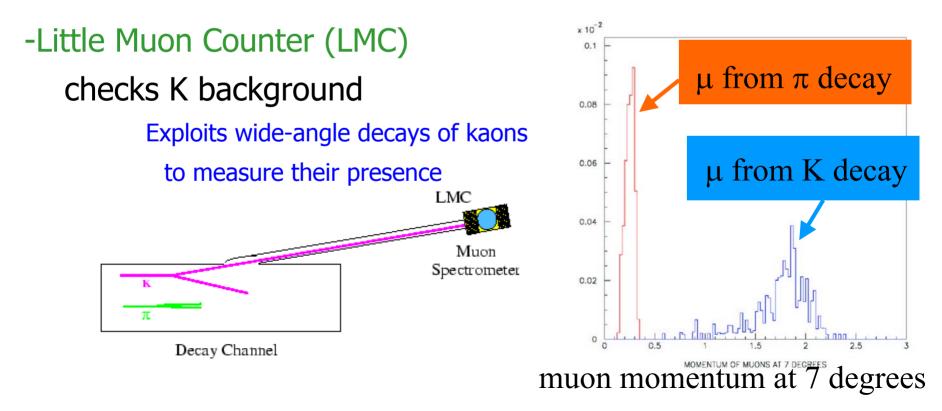
MiniBooNE secondary beam cross-checks

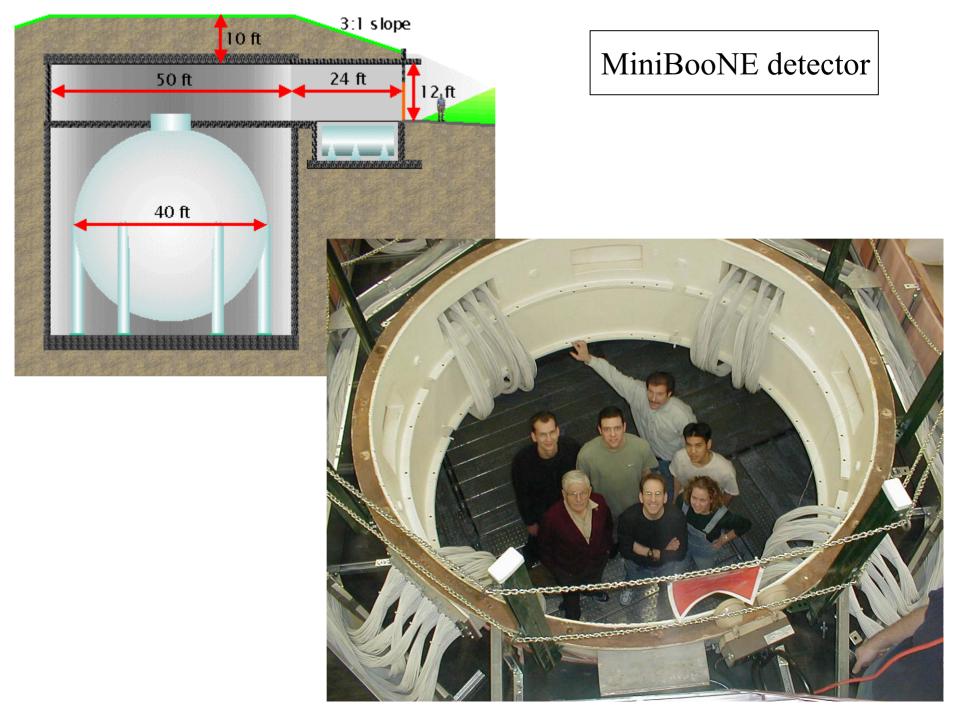
-Varying the length of the decay region from 50 m to 25 m checks  $\mu$  background

Rate of  $v_{\mu}$  from  $\pi$  depends on *L*, whereas rate of  $v_{e}$  from  $\mu$  depends on *L*<sup>2</sup>.

Therefore, if an excess is the signal, the rate will change by  $\times 2$ 

Excess from unmodeled  $\nu_{e}$  from  $\mu$  decay will change by  $\times 4$ 



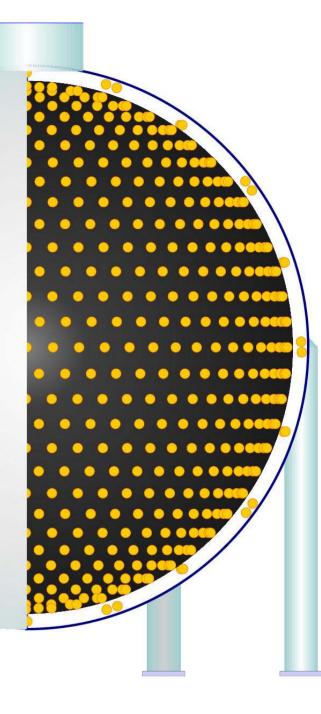


pure mineral oil

total volume:800 tons (6 m radius)fiducial volume:445 tons (5m radius)

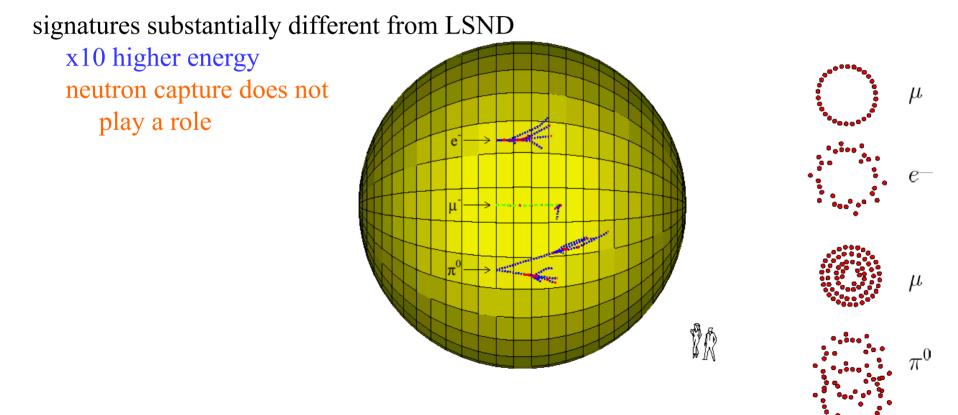
1280 20-cm PMTs in detector at 5.5 m radius
→ 10% photocathode coverage
240 PMTs in veto

Phototube support structure provides opaque barrier between veto and main volumes



Analysis:  $e, \mu, \pi^0$  discrimination

Pattern of hit tubes (with energy and time information) allows for the separation of different event types.





#### Laser calibration system

four Ludox-filled flasks attached to fiber optic cables

provide: PMT gain and timing and oil attenuation length monitoring



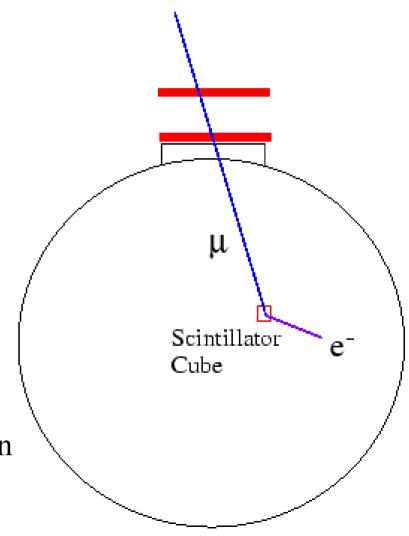
Optically isolated scintillator cubes in tank: six 3-inch (7.6 cm) cubes one 4-inch cube

Muons with known trajectory through the oil

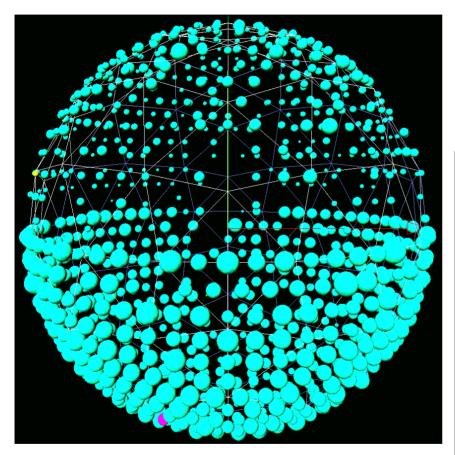
Provides: range for energy calibration cross checks on reconstruction algorithms

### Stopping muon calibration system

Scintillator tracker above the tank



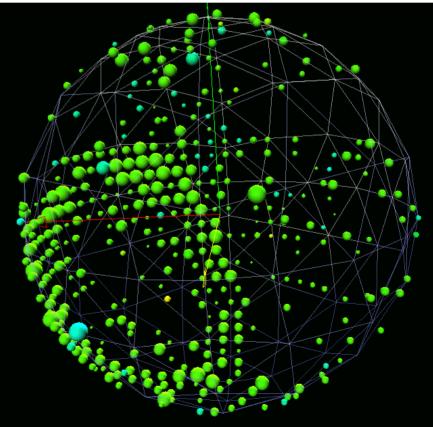
#### Calibration events are being collected



laser event, tank ~1/2 full of oil

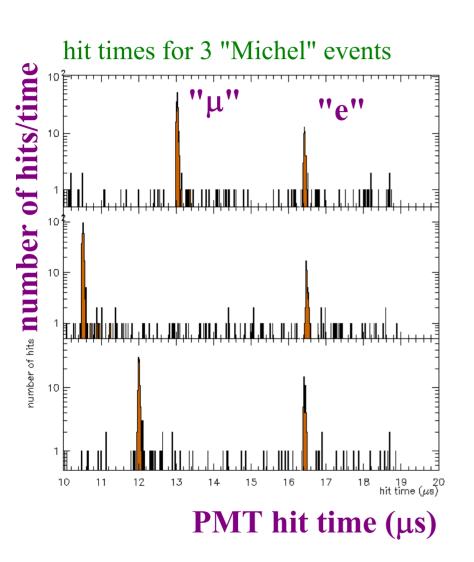
(size of each hit is proportional to charge)

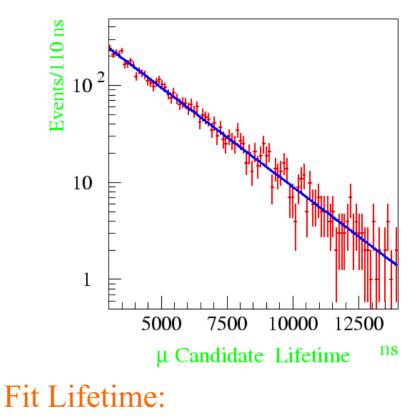
#### cosmic muon



## **Cosmic Muon Decays**

#### muon stops and the decay (Michel) electron is observed



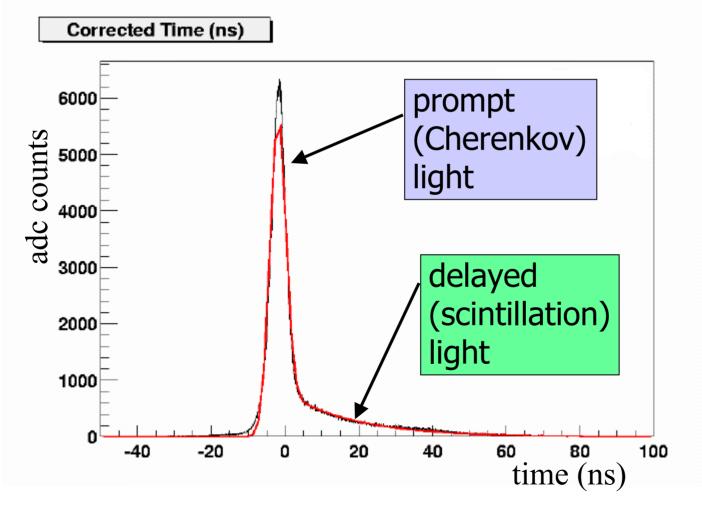


 $\tau = 2.12 \pm 0.05 \ \mu s$ 

Expected  $\mu$  lifetime in oil 2.13  $\mu$ s with 8%  $\mu^-$  capture on carbon.

#### Time spectrum of Michel electrons

Measure, e.g., time resolution scintillation time constant

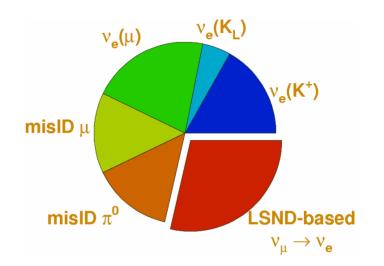


### MiniBooNE expected signal

 ${\sim}500k~\nu_{\mu}C$  charged current events

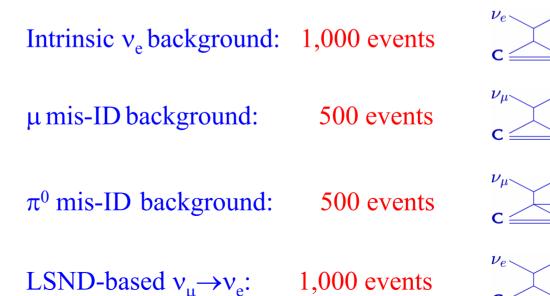
<u>Approximate</u> number of electron neutrino-like events



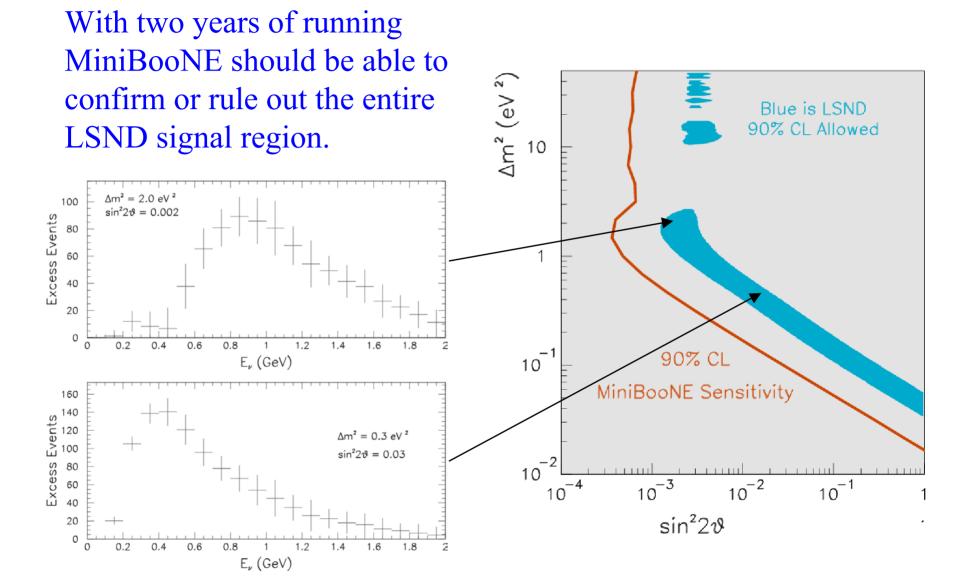


 $\nu_{\mu}$ 





# MiniBooNE expected sensitivity



#### MiniBooNE status

Neutrino beam to be delivered in August.

summer 2002:

May: detector full of mineral oil June: detector calibration proton line commissioning horn installed, hot horn handling demonstrated July: horn removed and protons delivered through target pile, study spot size, beam monitors

August: final beam/horn/target configuration and start of high intensity running



