# The UHECR Spectrum with HiRes

Douglas Bergman Rutgers University ICHEP 2002, Amsterdam 26 July 2002

## The HiRes Experiment

- Study UHECR with Air Fluorescence
  - Spectrum
  - Composition
  - Sources/Anisotropy

 Currently has the highest exposure above 10 EeV

### **HiRes** Collaboration

J.A. Bellido, R.W. Clay, B.R. Dawson, K.M. Simpson

#### **University of Adelaide**

J. Boyer, B. Knapp, W. Lee, E.J. Mannel, M. Seman, C. Song, S. Westerhoff, X. Zhang

#### **Columbia University**

J. Amann, C. Hoffman, M. Holzcheiter, L. Marek, G. Sinnis, T.N. Thompson, D. Tupa

#### Los Alamos National Labratory

J. Belz, M. Munro, M. Schindel

### The University of Montana

G. Martin, J.A.J. Matthews, M. Roberts

### **University of New Mexico**

D.R. Bergman, S. Karg, L. MacLynne, L. Perera, S. Schnetzer, G.B. Thomson, A. Zech

### **Rutgers University**

N. Manago, M. Sasaki, M. Sasano

### **University of Tokyo**

T. Abu-Zayyad, G. Archbold, K. Belov, Z. Cao, M. Dalton, J. Girard, R. Gray, W. Hanlon, B. Jones, C.C.H. Jui, D. Kieda, K. Kim, E.C. Loh, K. Martens, J.N. Matthews, J. Meyer, S.A. Moore, A.N. Moosman, J.R. Mumford, K. Reil, R. Riehle, P. Shen, J. Smith, P. Sokolsky, R.W. Springer, B.T. Stokes, S.B. Thomas, L. Wiencke

### **University of Utah**

# **HiRes** Location

- West Desert of Utah on Dugway Proving Grounds
- 100 miles WSW of Salt Lake City



# HiRes Sites



- 21 Mirrors
  - 360 deg in azimuth 3-17 deg in elevation
- Sample & Hold DAQ
- Began observation: June 1997



- 42 Mirrors
  360 deg in azimuth
  3-33 deg in elevation
- FADC DAQ
- Began observation: October 1999

# HiRes Mirror & PMT Cluster



### 5.1 m<sup>2</sup> segmented mirror

256 pixel PMT array and UV filter

# **Data Reconstruction & Analysis**

- Find Shower-Detector Plane
- Fit time vs angle to determine geometry
- Convert number of photoelectrons to number of particles in shower
- Integrate, using average energy loss, to find energy



# Data/MC Comparison: Geometry

- Data/MC comparisons constrain possible errors in aperture calculation
- Note that Psi and Rp are correlated



# Data/MC Comparison: Energy

- Energy distributions also match
  - Depends on spectrum and composition in MC
  - Input spectrum divides out of aperture



# HiRes Spectra

 Excellent agreement between HR-1 and HR-2



# **Uniform Source Model Fit**

- Fly's Eye composition suggests heavy-to-light change in ankle region
- Use this to motivate a two component spectrum
  - Galactic component, power law with linear factor from composition measurement
  - Extragalactic component assuming uniform source density as in Berezinsky et al., hep-ph/0204357



# HiRes vs AGASA

- Disagreement between HiRes and AGASA reduced:
  - Factor of 2 in flux
  - A few points over 100 EeV
- Scaling AGASA energies down by 20% makes apparent the nice agreement in ankle region



# HiRes vs AGASA

Agreement below
 60 EeV and
 disagreement
 above quite visible
 in unmodified flux
 comparison



## Conclusion

- HiRes has measured the spectrum of Cosmic Rays at the highest energies and with the highest sensitivity at these energies
- We observe a feature consistent with what's expected from the GZK process
- HiRes has observed several strong candidates at energies above the the GZK energy